

Towards calibration of the scintillation detectors using uDAQs

AIR SHOWER:
3 – 10 KM LENGTH
200 M DIAMETER



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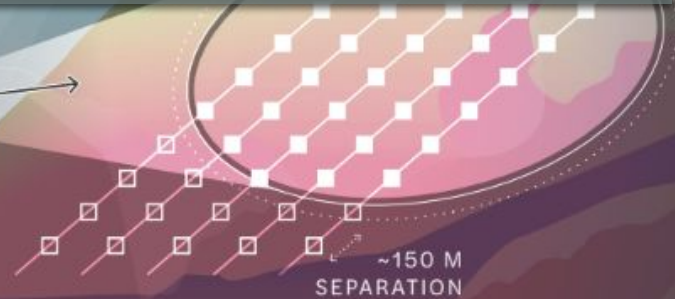
RANGE:
50 M – 5 KM

CHARGED-CURRENT
INTERACTION

ROCK

> 4 KM SHIELDING FROM
BACKGROUND MUONS

*TAMBO Collaboration Meeting
Shefali (17.08.26)*

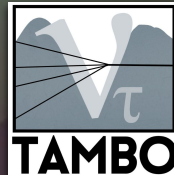


AIR SHOWER
DETECTOR ARRAY
~M² EACH



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DEEP VALLEY

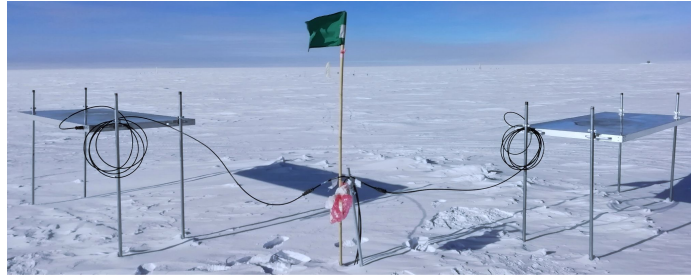


About Me

- Started as a Postdoc at the Harvard group in July 2026
- My PhD was focused on everything hardware and data analysis with the scintillation detectors for the IceCube Surface Array Enhancement at KIT, Germany
- During my masters I did the mechanical design of the acoustic calibration detector for the IceCube Upgrade RWTH Aachen University, Germany
- I like plants, and being in the nature whenever I get the chance :)



Acoustic Pinger

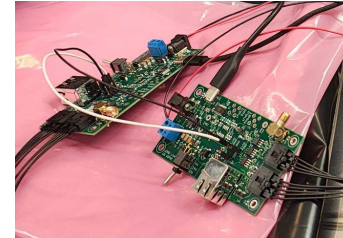
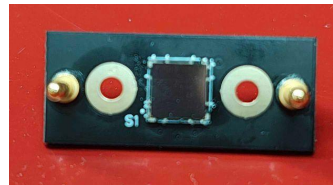
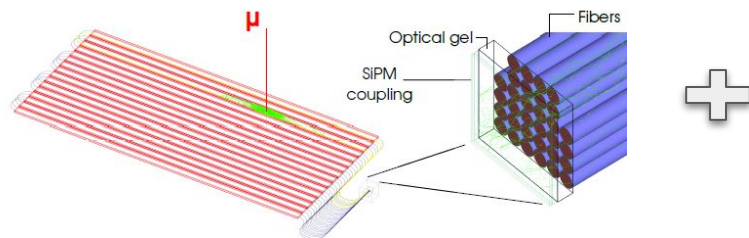


SAE Scintillation detectors



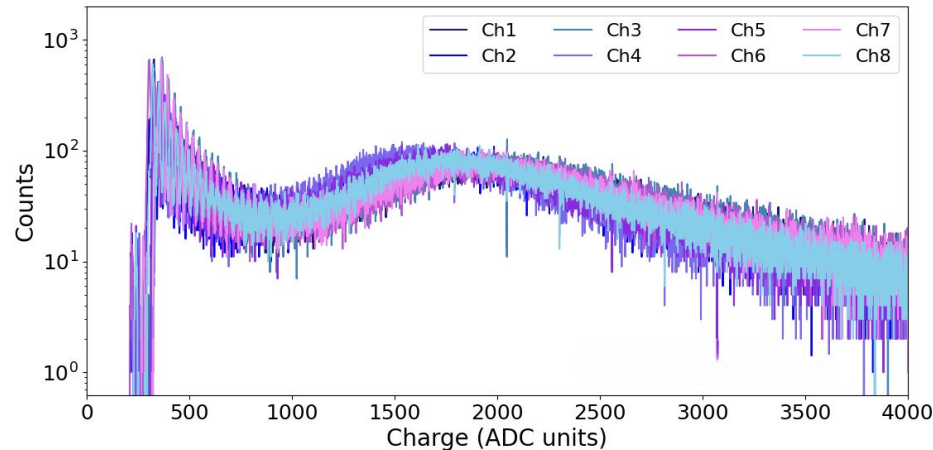
Status Quo

- In the virtual meeting and the panel review scintillation detectors readout by SiPM + uDAQ were finalised for the prototype design
- Challenges: For shower reconstruction and all the cool physics, we first need to calibrate the system
 - SiPM gain is inversely proportional to the ambient temperatures
 - We have a lot of natural background that masks our signal
 - uDAQ has limited functionality: dynamic range, buffer and more
- I have been working with uDAQs to understand the caveats and possible ways forward



How can we calibrate the detectors?

- Minimally Interacting Particles (MIPs) deposit roughly the same energy when traversing the detectors, leading to a characteristic peak
- At low temperatures gain of the SiPMs can be high enough to resolve the single photoelectron peaks

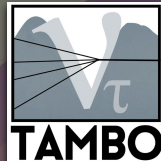


**Example histograms from 8 SAE scintillators*

** uDAQ gives a digital output after shaping the pulses.*

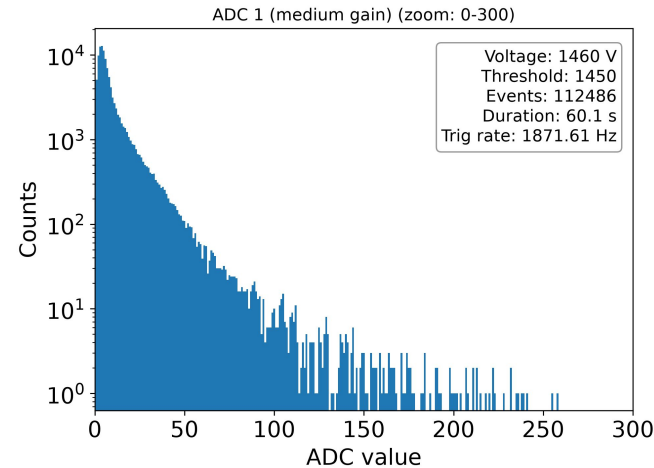
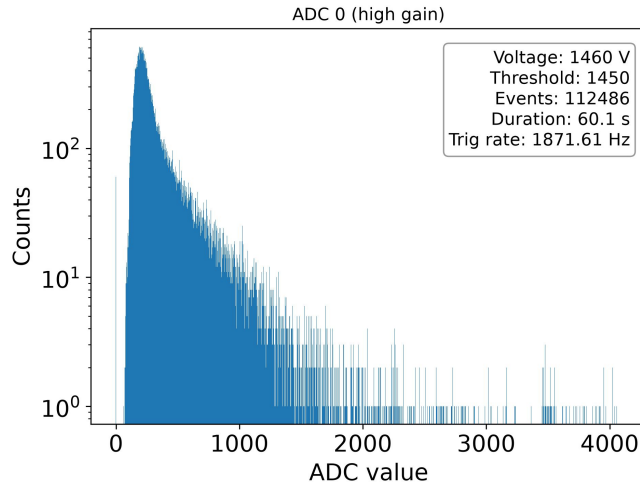


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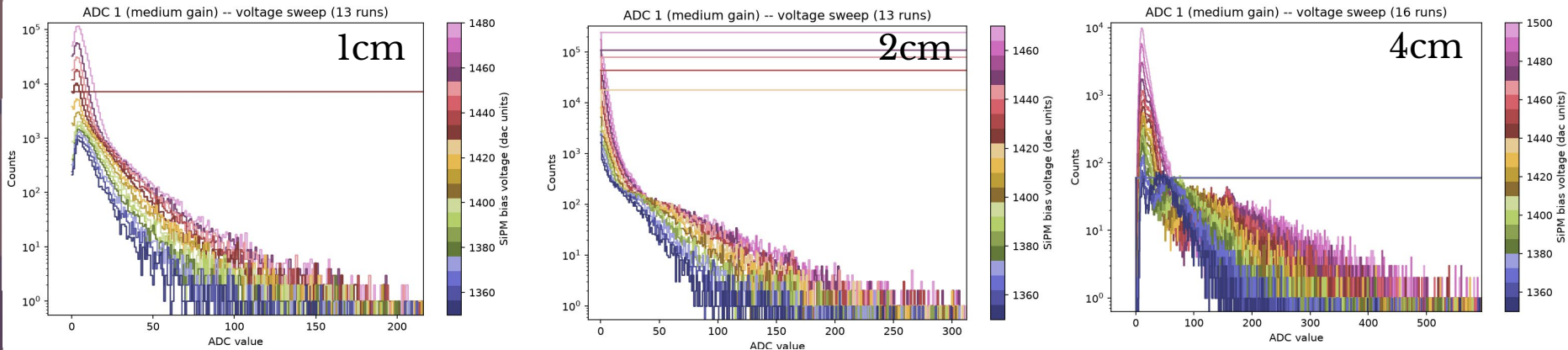
What do we see in the lab?

- The high gain channel resolution is not great
- The MIP peak distribution is completely masked by Dark Rate and low energy natural radioactivity background
- Both gains have pile up features that can further affect the dynamic range



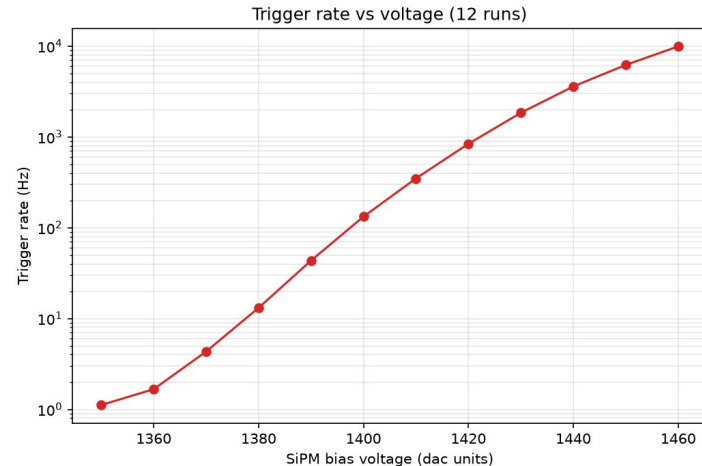
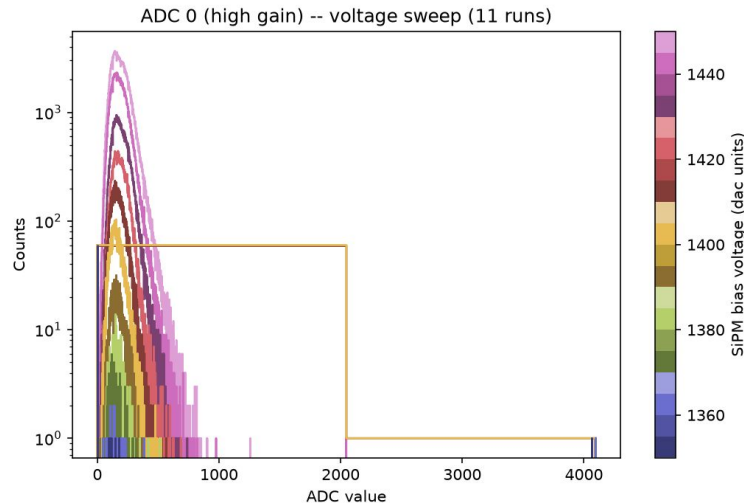
Can we see the MIP with increased thickness?

- Supplied Bias Voltage scans performed with 3 different thickness compositions
 - Increased thickness allows higher energy deposit.
 - The MIP distribution becomes more evident already at 2cm for significantly lower bias voltage
 - Higher voltage introduces higher dark rate noise
- Useful but very expensive, and therefore impractical to use higher thickness



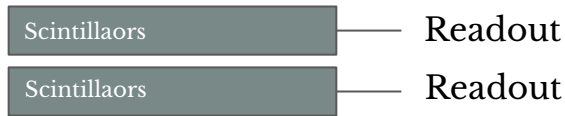
What does the SiPM see without scintillators

- Voltage scan with the same SiPM used previously, but without the incident fibre bundle
 - With 1cm thickness we are in the order 4 region
 - Lower temperatures will improve this picture

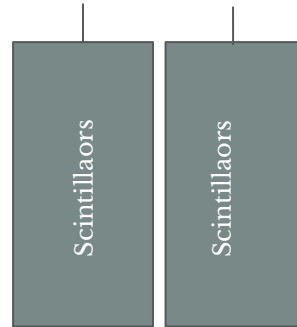


Coincidence Measurements

- Exploring coincidence measurements for calibration in 3 different configurations
 - Half length panels with 4 scintillation bars each used
 - Performed measurements in the lab as well as outside to account for overburden



Stacked (1)



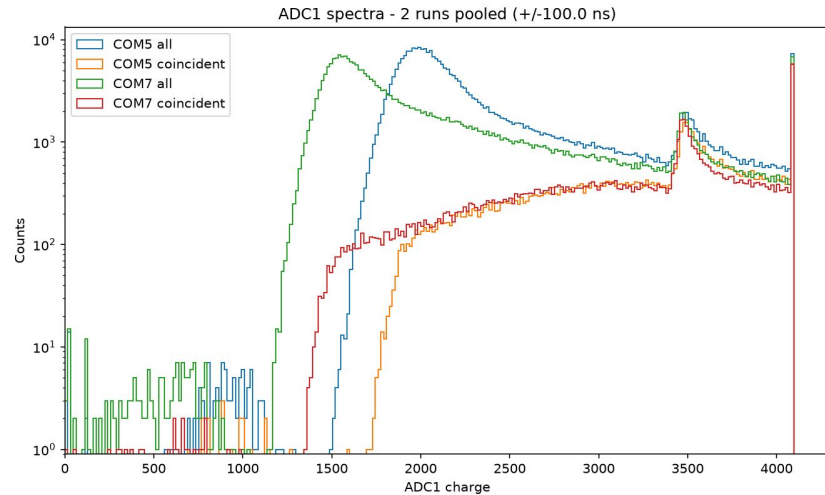
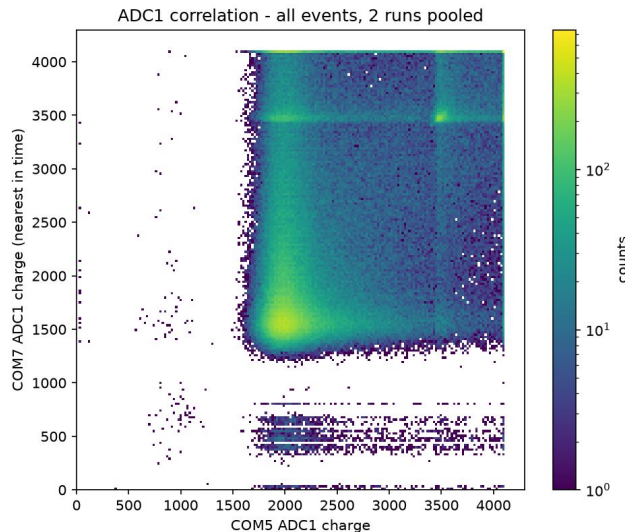
Sideways (2)



Stacked Hybrid (3)

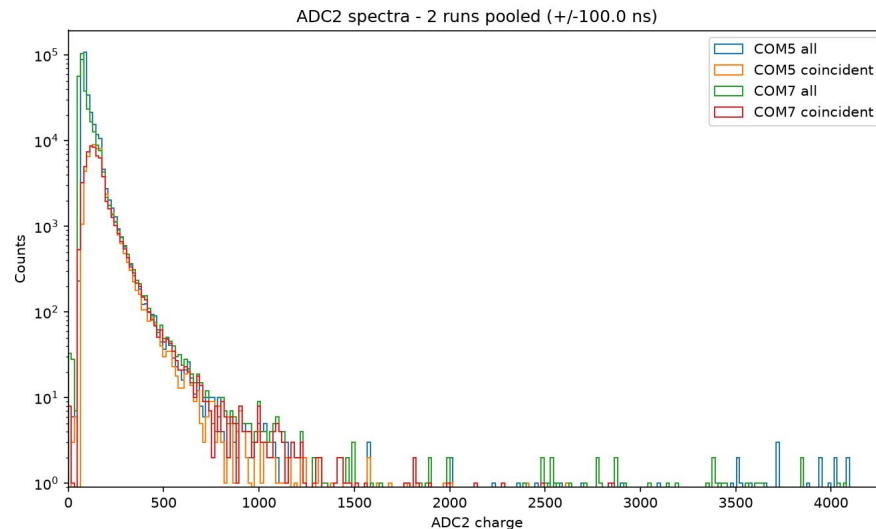
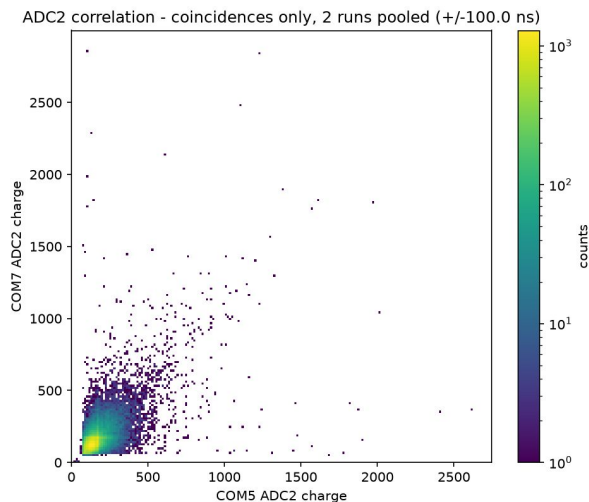
Coincidence Measurements

- Preliminary studies of the coincidence rate in the 3 configurations
- Config 1: Stacked on this slide. **Coincidence rate: ~22Hz**
 - High gain channel cannot resolve the MIPs
 - Pile up around 3500 ADCs



Coincidence Measurements

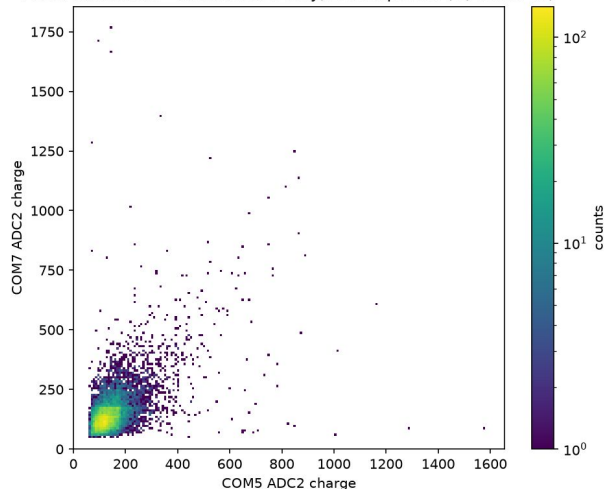
- Preliminary studies of the coincidence rate in the 3 configurations
- Config 1: Stacked on this slide
 - Medium gain allows a look into the MIP peak



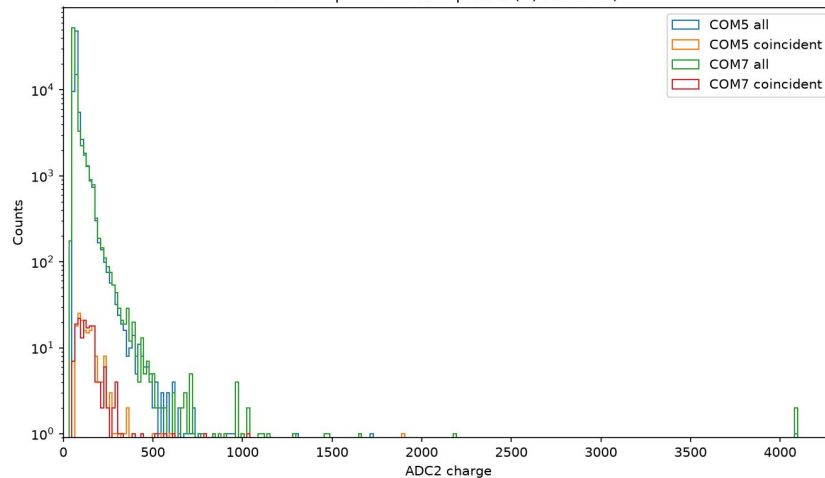
Coincidence Measurements

- Configuration 2: Side by side :Coincidence rate: $\sim 0.7\text{Hz}$
 - Allows us to cover more effective area, and useful for detector response studies
 - Rate should improve with size and lower temperature

ADC2 correlation - coincidences only, 2 runs pooled ($\pm 100.0\text{ ns}$)



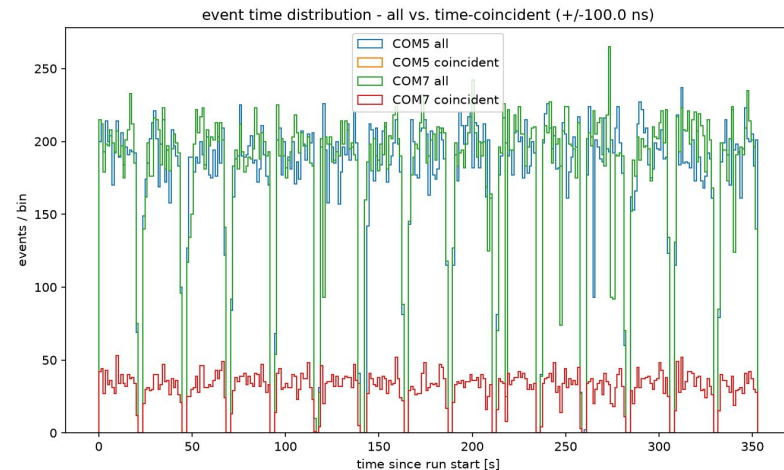
ADC2 spectra - 2 runs pooled ($\pm 100.0\text{ ns}$)



Summarized Rates

- There is clearly an improvement in the number of MIPs detectable without the overburden
- The rate in sideways coincides is limited by its size
- The rates are also however limited by the electronics:
 - Buffer flushing highly time expensive (~3s deadtime), needed every 4-6s for lower thresholds

Configuration	Rate
Stacked (lab)	22.41 Hz
Side-by-side (lab)	0.79 Hz
Stacked with MTA (lab)	0.71 Hz
Stacked (outdoors)	25.67 Hz
Side-by-side (outdoors)	0.69 Hz

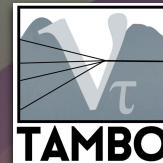


Takeaways and Outlook

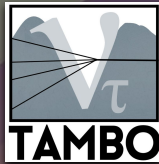
- The measurements with the complete prototype are challenging but we are making progress :)
- Calibration can potentially be achieved with optimal effective area in the 2nd configuration as well as the tagger combination with longer calibration runs (~20mins).
- SiPM dark rates need to be studied more extensively, preferably in a temperature range
- Setup a SiPM calibration test bench
- uDAQ limitations need to be addressed, firmware and hardware
 - Dedicated work will start soon in the Hawaii group



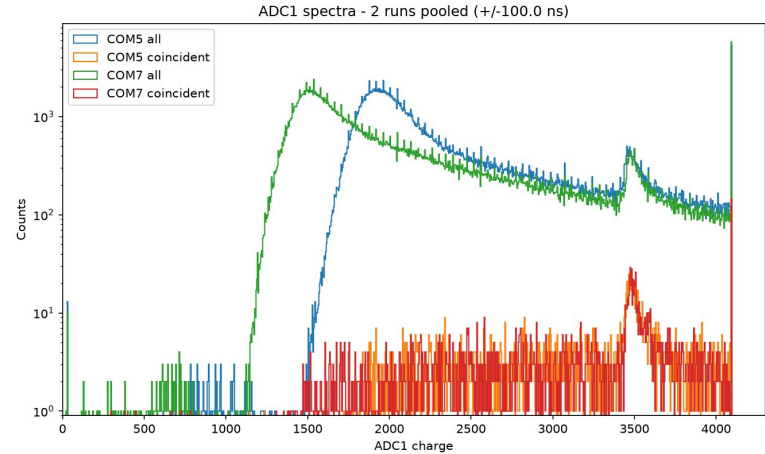
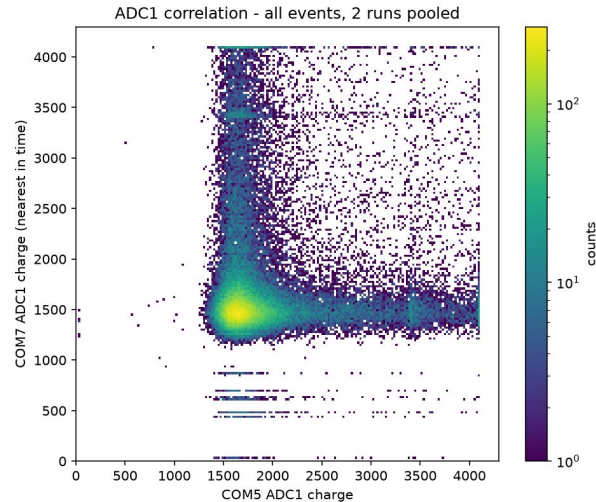
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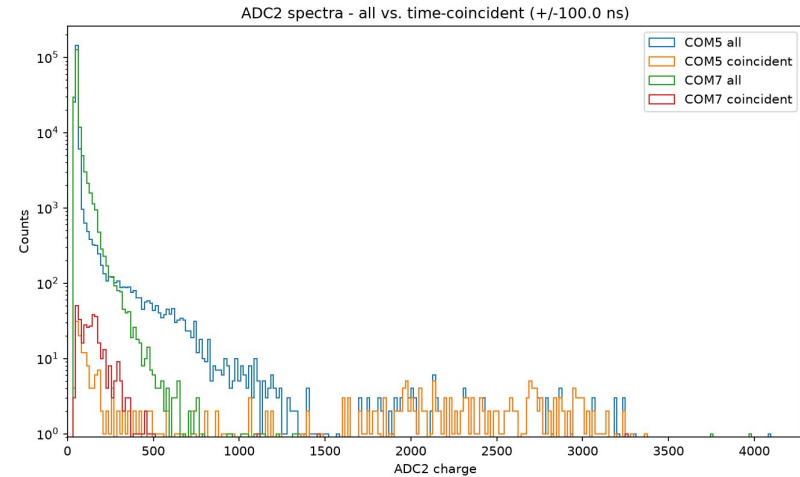
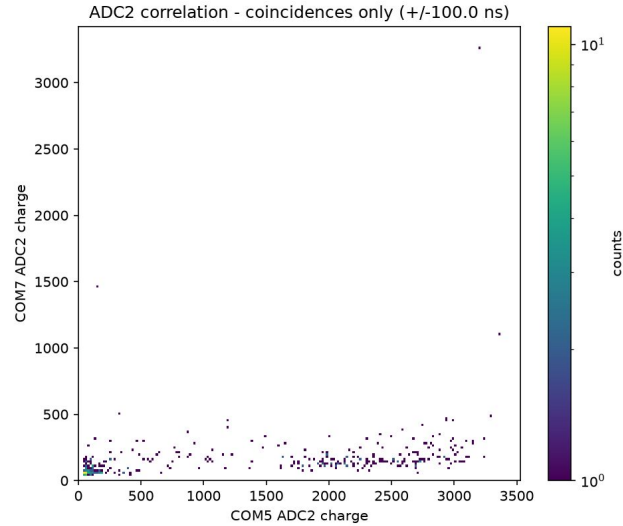
BACKUP SLIDES



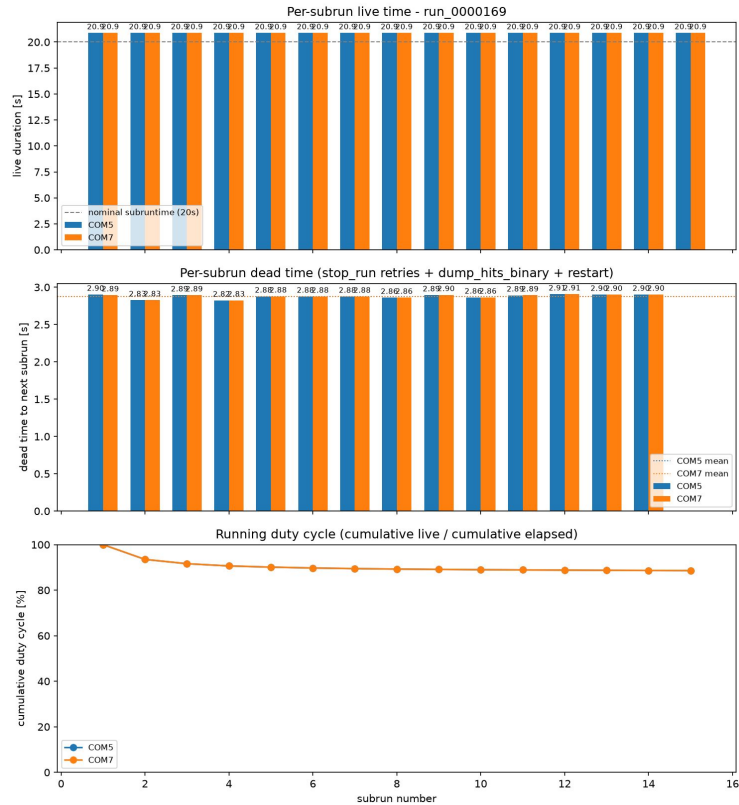
Coinc measurements in side-ways config



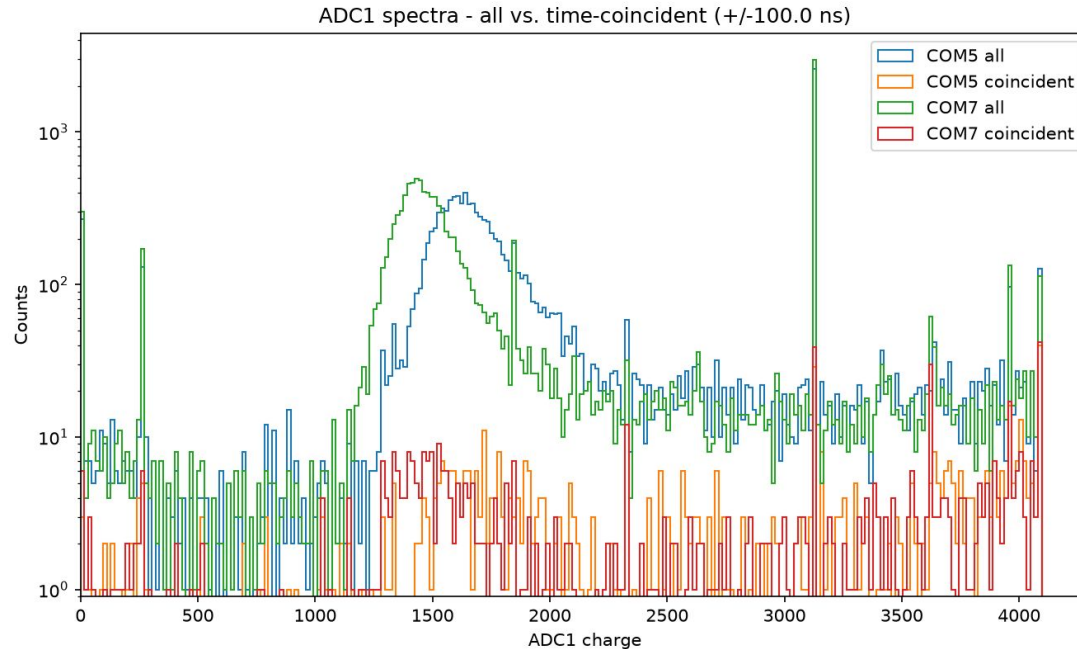
Coinc measurements with MTA



Some deadtime/Runtime stats



Timing related issues in the hitbuffer output



Average coincidence rates

