

Hardware Overview

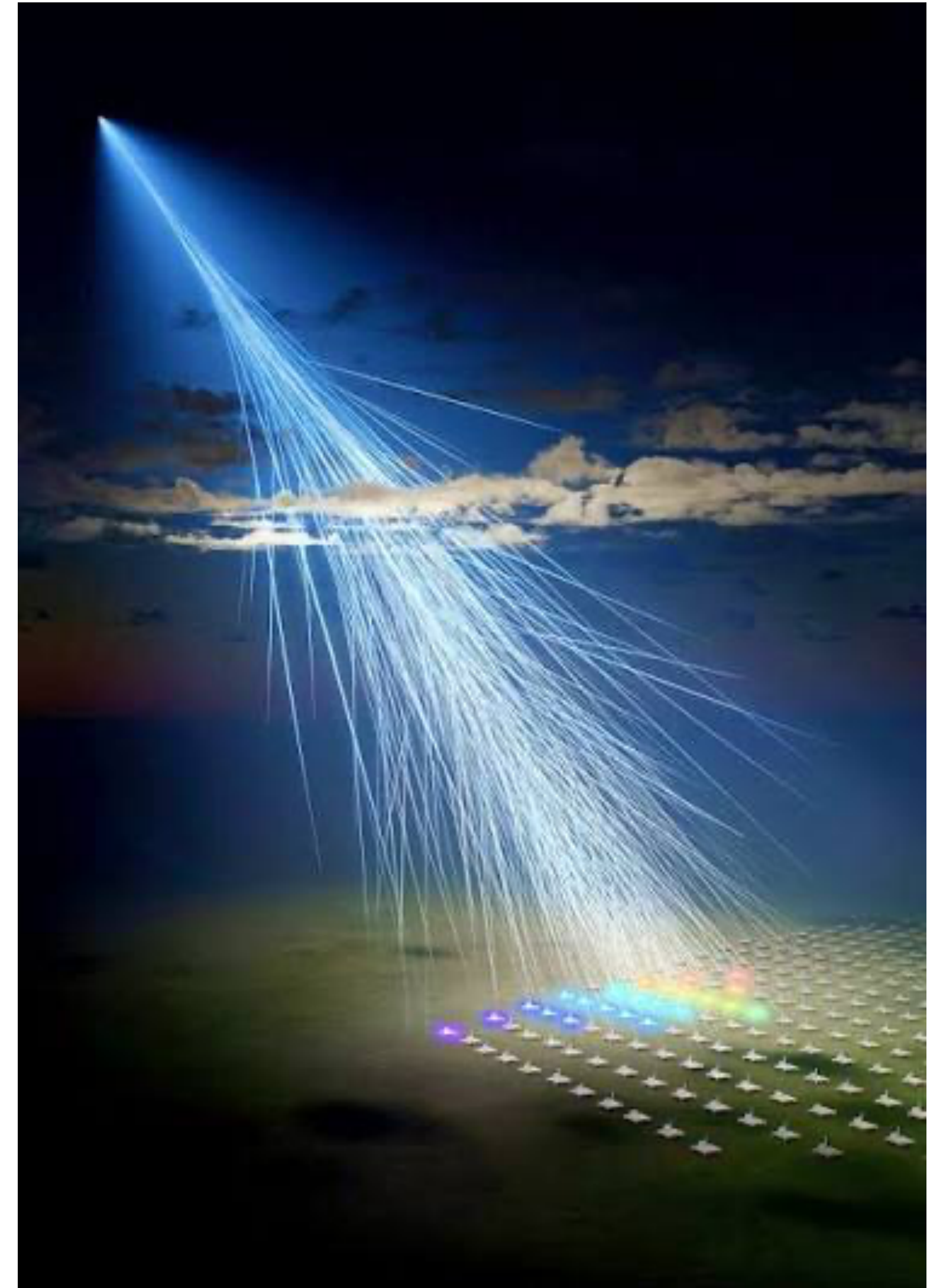
Will Thompson

TAMBO Collaboration Meeting — Summer 2026

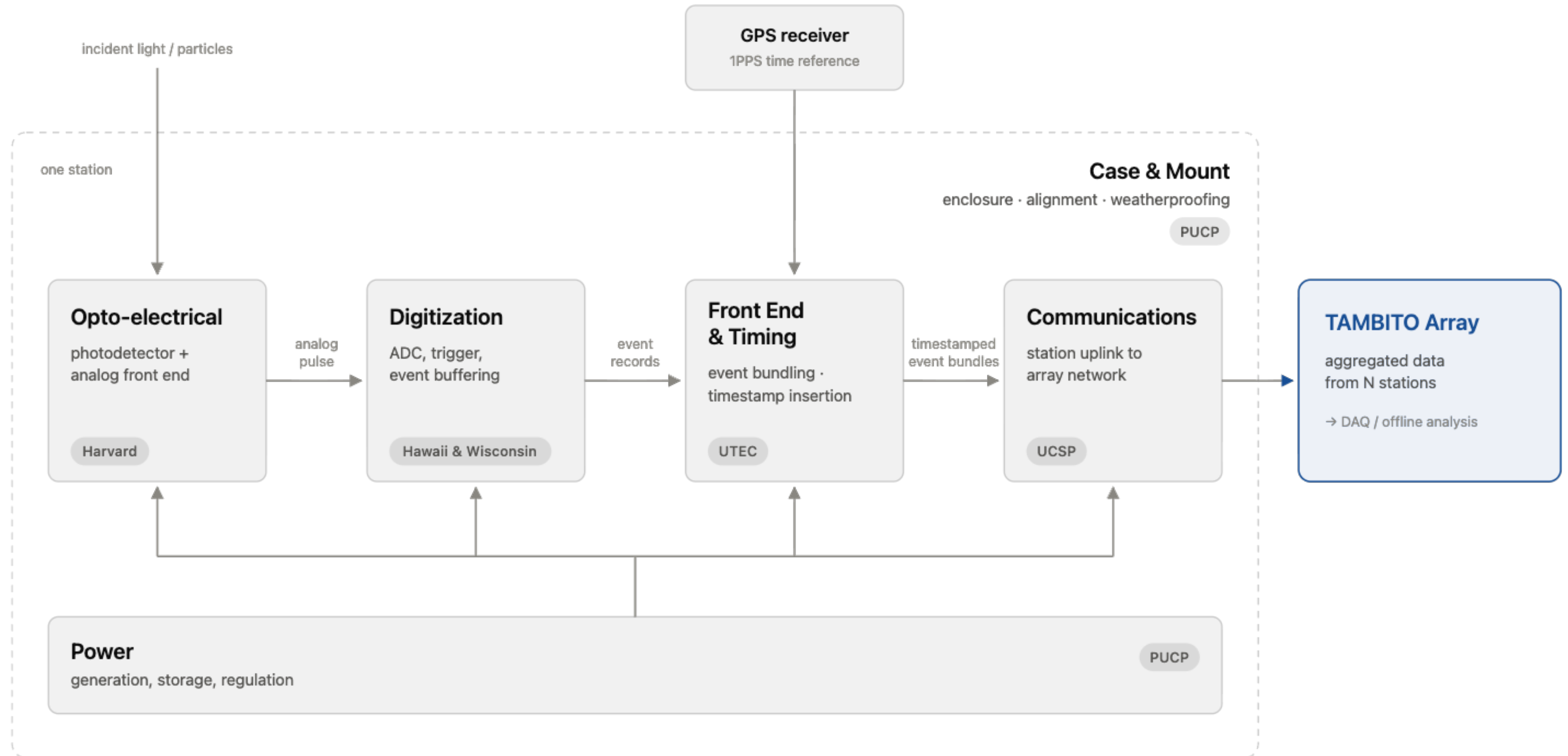
Big picture hardware requirements

Each individual detector must:

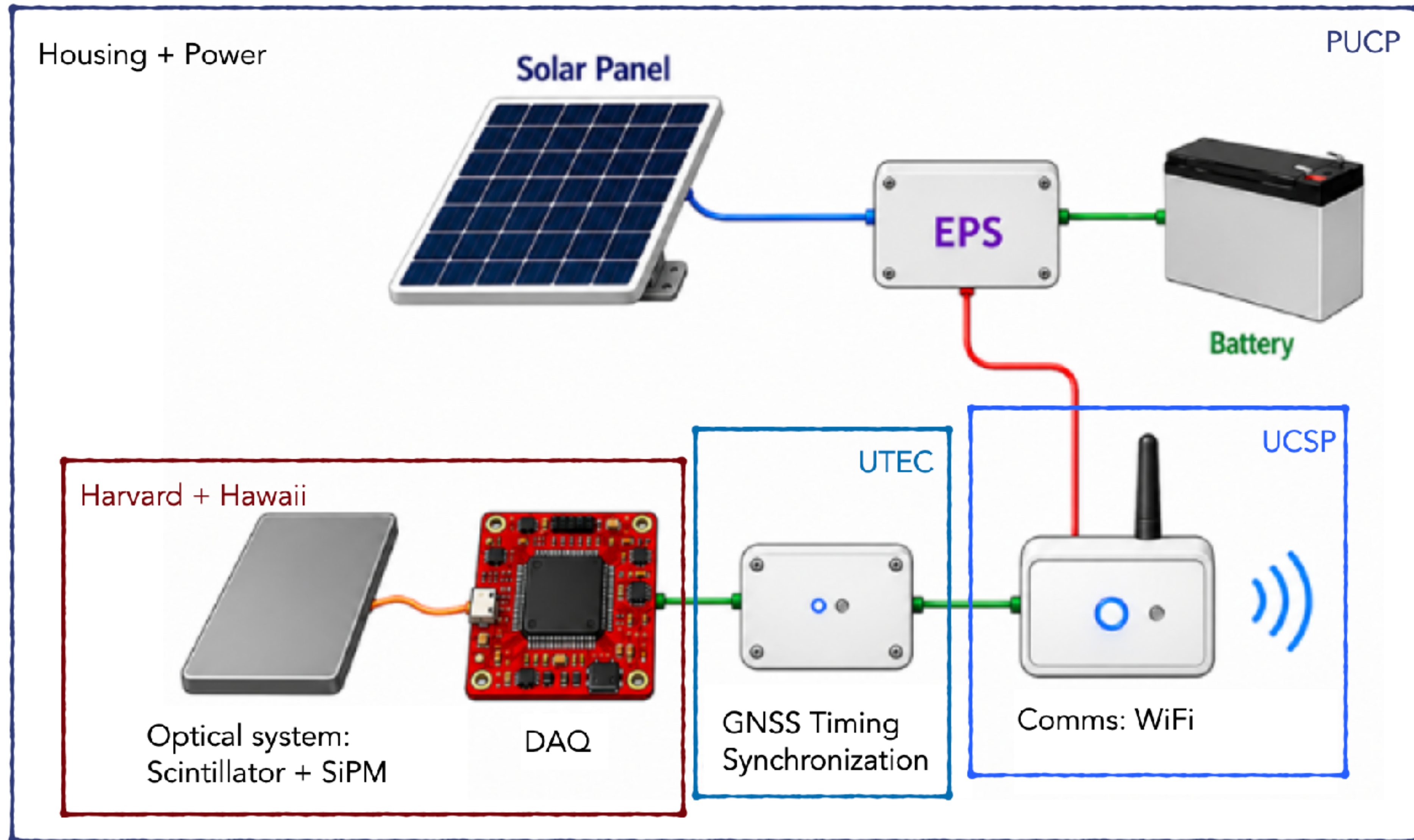
- Generate a response (electrical signal) when struck by an air-shower particle
- Digitize electrical signal
- Timestamp each event and bundle together all event information
- Send event to central DAQ system



TAMBITO array subsystems



TAMBITO array subsystems



Actually building the array requires a lot of engineering on each subsystem, so things are getting more serious

A set of requirements is essential for the engineering design!

What do our array requirements look like?

Some are very physics-driven:

- “Aperture $\geq 400\text{m}^2\text{sr} \times E_{\text{PeV}}^{1.5}$ from 1-10 PeV and 10x IceCube from 10-100 PeV”
- $\leq 1^\circ$ air shower pointing resolution

Others are more operations driven:

- Power consumption less than xxx W per station
- Station failure rate less than xxx per year
- Construction compatible with local sourcing and repair

Example: 1° pointing requirement

Precise pointing important so we can:

- Determine the direction of incoming neutrinos (e.g. what galaxy did they come from?)
- Reject cosmic ray backgrounds

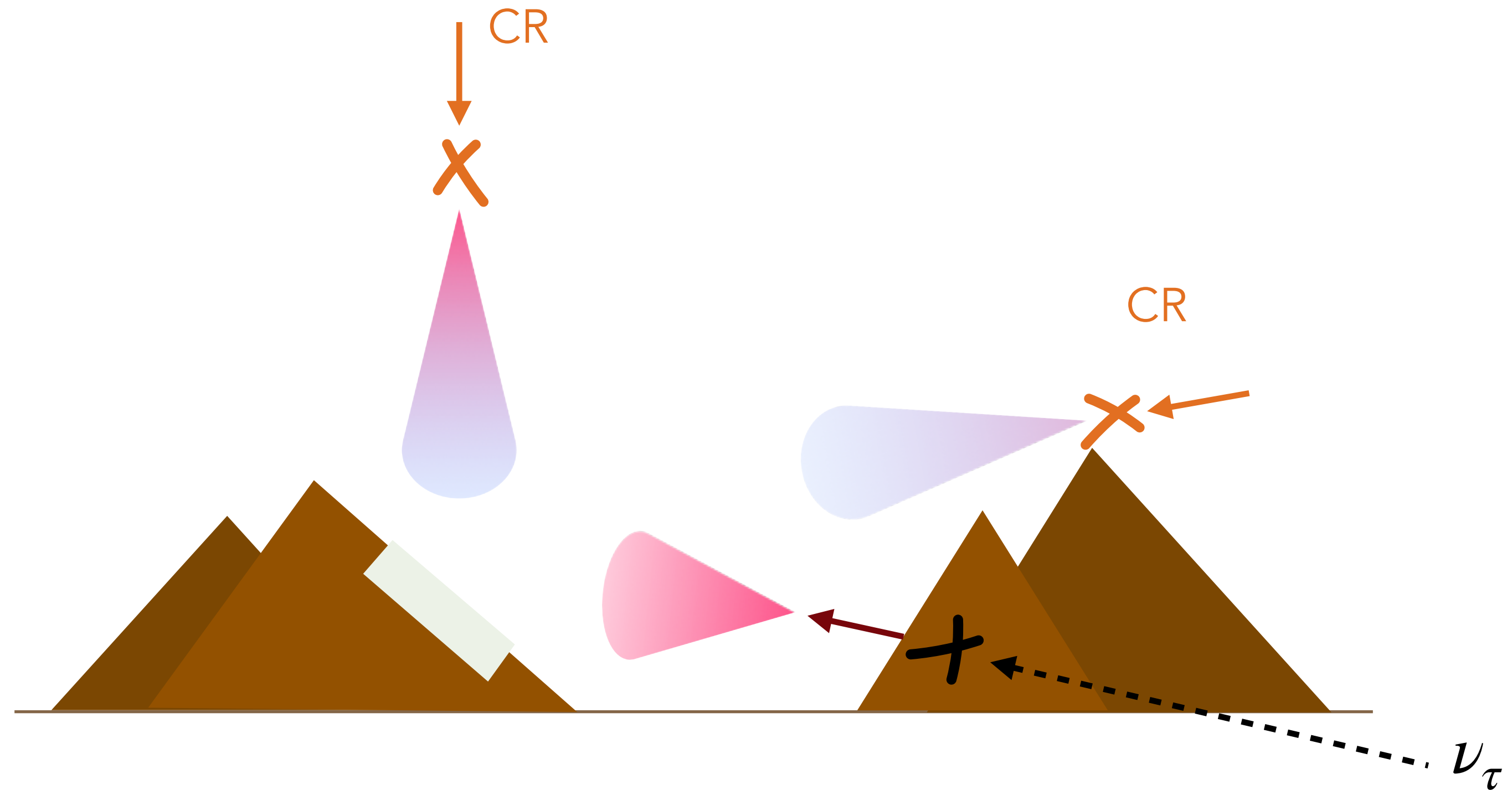
Target: $\sigma_\theta \leq 1^\circ$

Direction uncertainty goes roughly as:

$$\sigma_t \leq \frac{\sqrt{2} D \sigma_\theta \cos \alpha}{c} \approx 7 \text{ ns}$$

Assuming $D = 150\text{m}$, $\alpha = 35^\circ$

Note that 7 ns is a Gaussian 1σ uncertainty (~ 24 ns width uniform distribution)



Current timing budget by subsystem

	Front End	Scintillator	Digitization	Synchronization	Shower
Required	5.5	6.0	1.0	9.4	1.0
CBE	1.6	2.2	1.0	5.0	1.0



	Timing Error
Required	12.5
CBE	5.9

Technical Performance Reviews (TPRs)

Hardware design is progressing with TPRs every ~2 months

- These reviews make sure we are hitting our requirements and give an opportunity for outside feedback
- Our first TPR was in July, next in September

All information about the different subsystems can be found in the subsystem DSNs

¿Preguntas?

Scintillator R&D at Harvard

Will, Kiara, Shefali
TAMBO Collaboration Meeting — Summer 2026

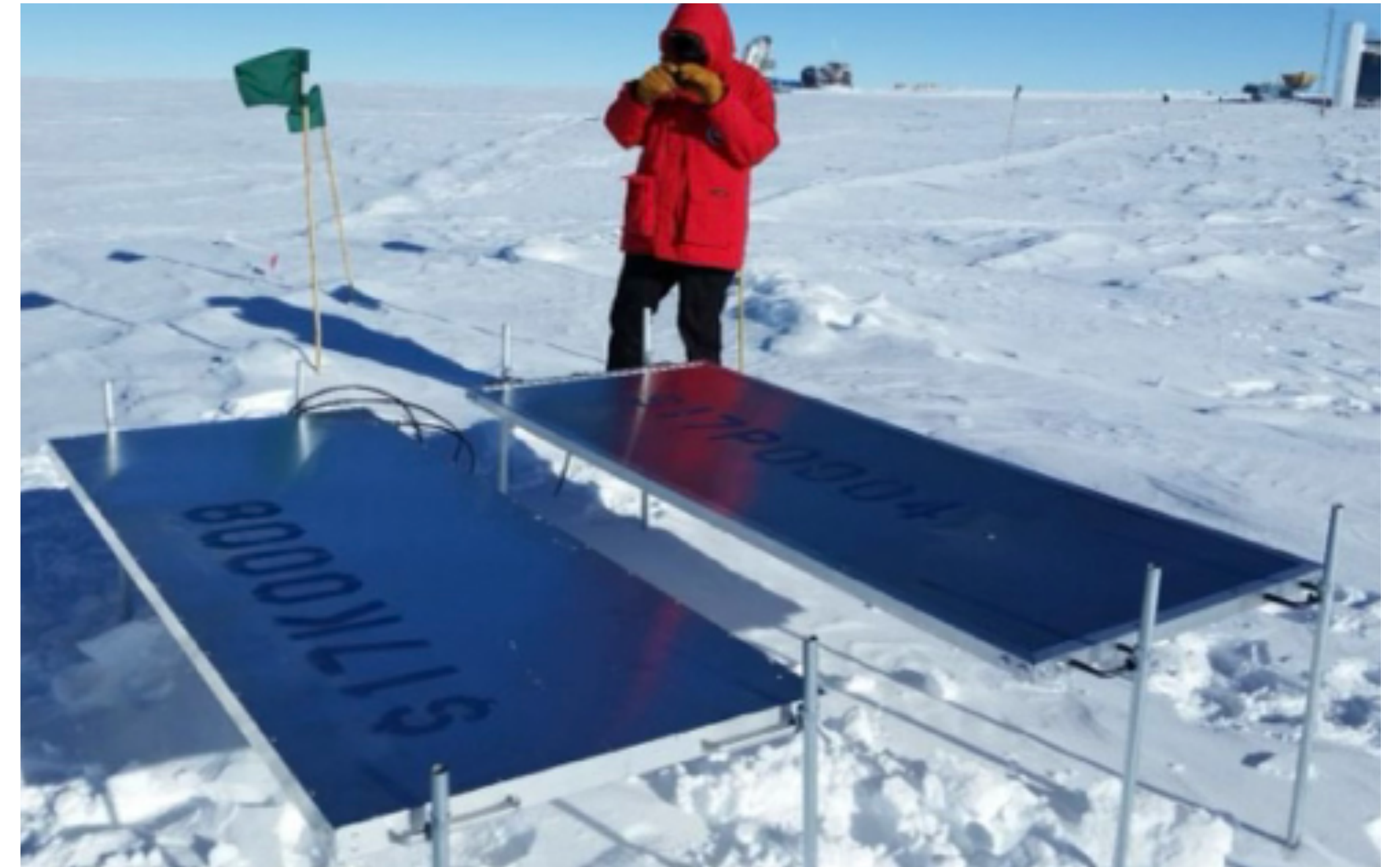
Opto-electrical system

Two main types of air shower detectors. Which should we use?



Water tanks:

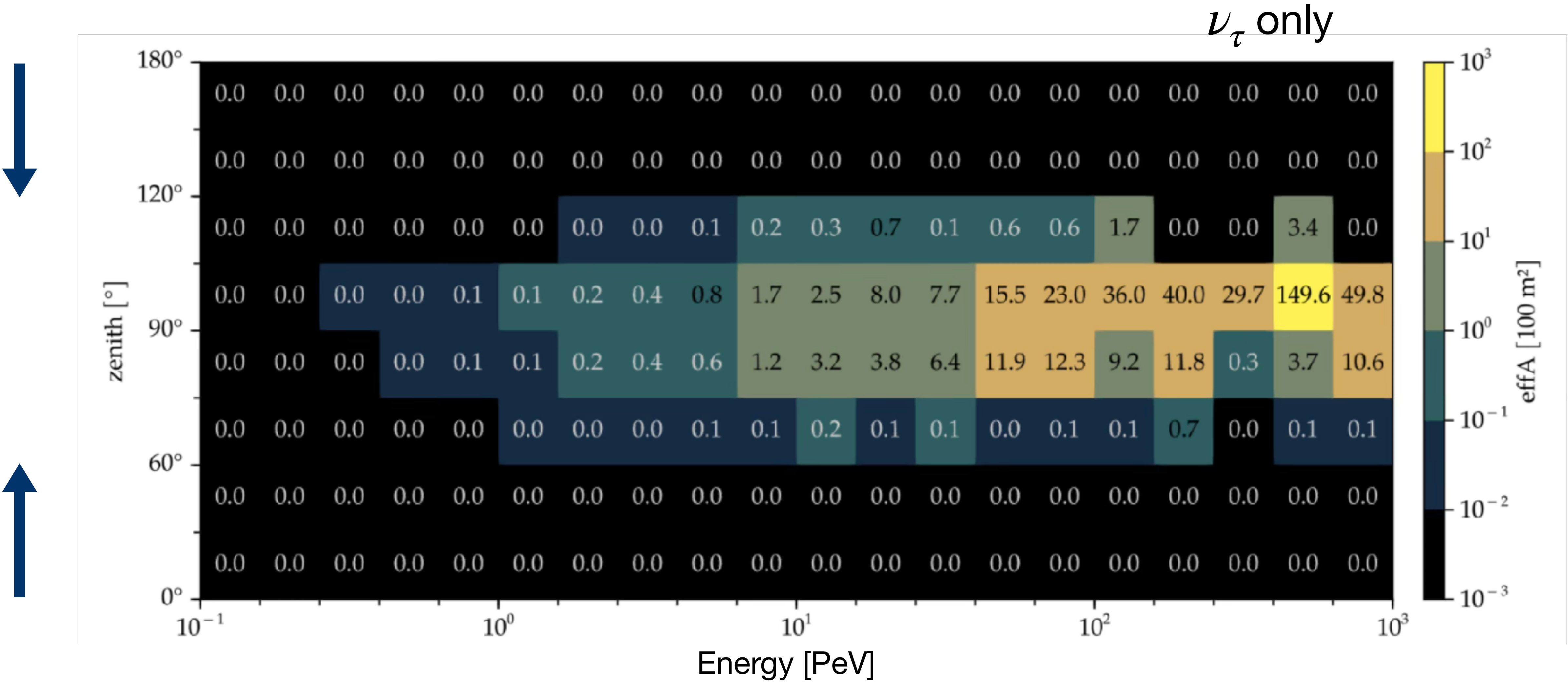
- Highly gamma-sensitive (most abundant particle)
- Expensive to purify and transport water
- Difficult to deploy in canyon



Plastic scintillator panels

- Sensitive mostly to charged particles
- Can be built off-site and transported
- Easier to deploy in canyon

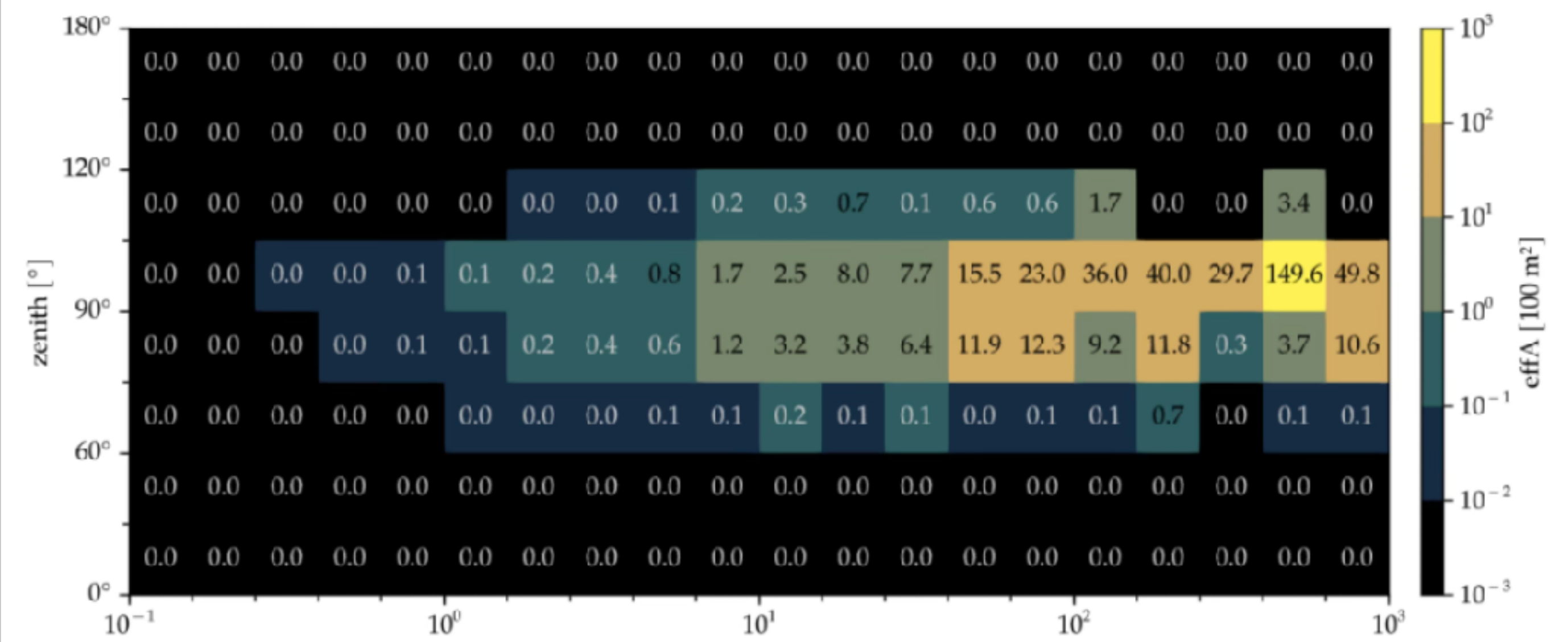
Tanks v. Panels



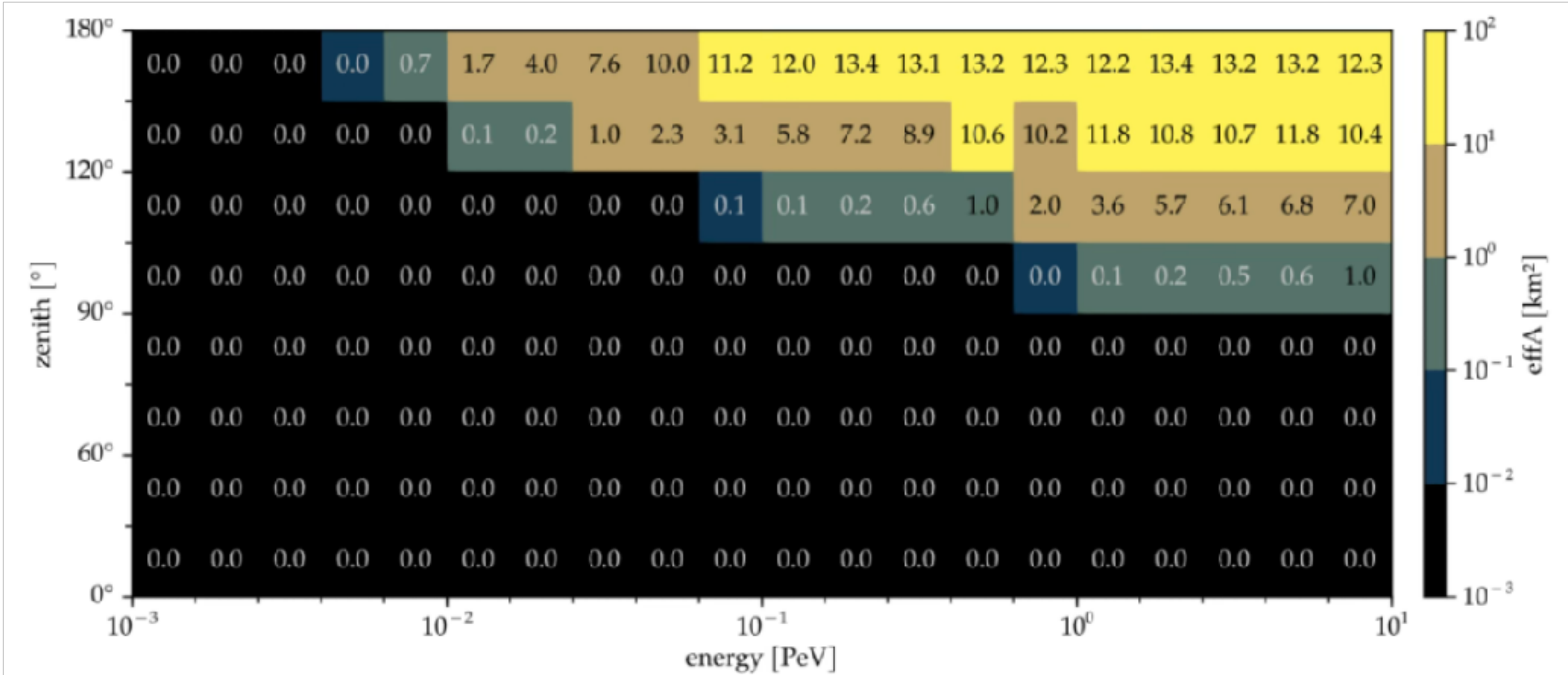
Tanks v. Panels

Filter condition: at least
one hit on at least five
modules

“water tank-like”



ν_τ rate: 0.47 / yr



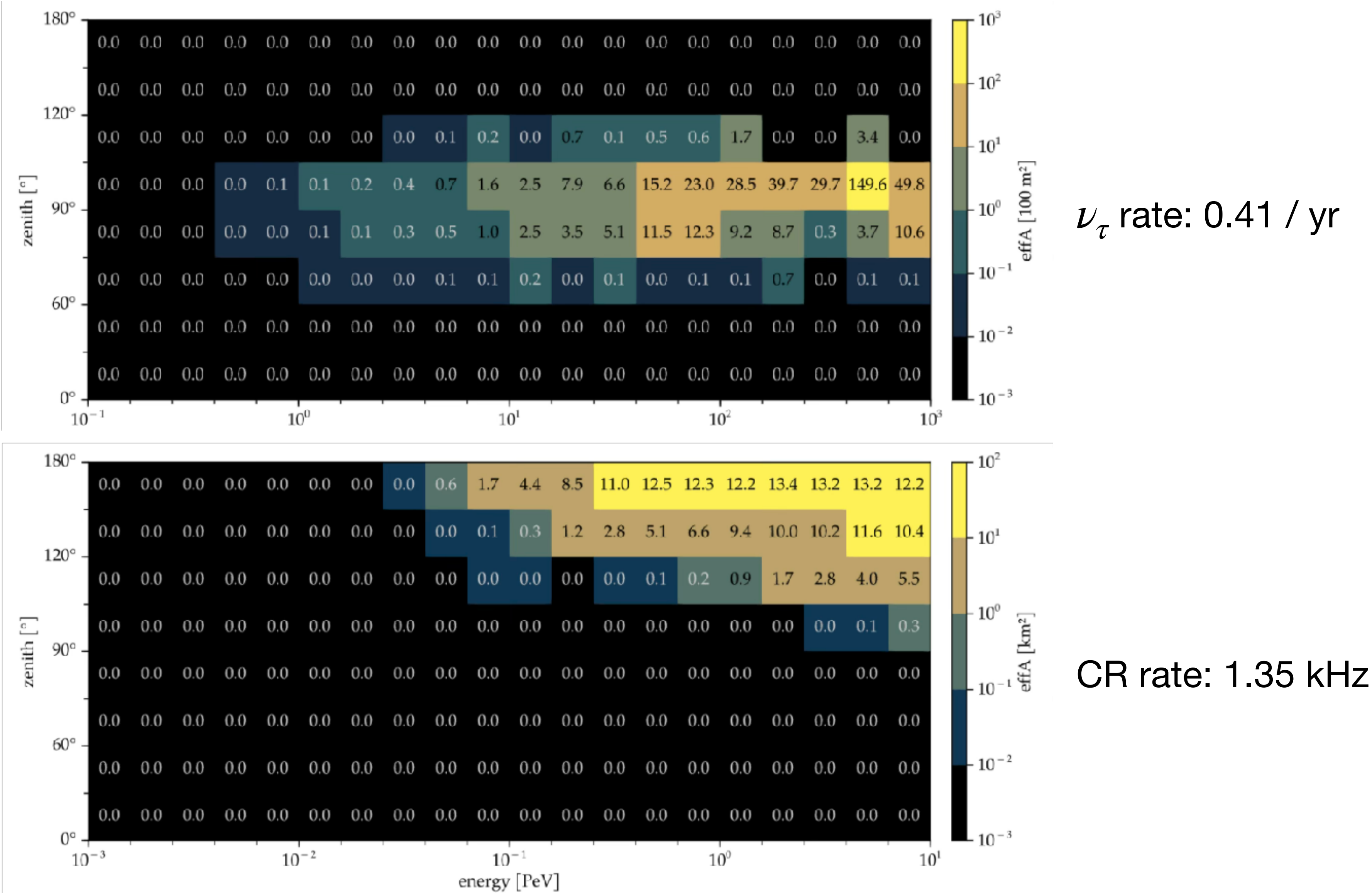
CR rate: 30.2 kHz

Tanks v. Panels

Filter condition: at least one (electron or muon) hit on at least 5 modules

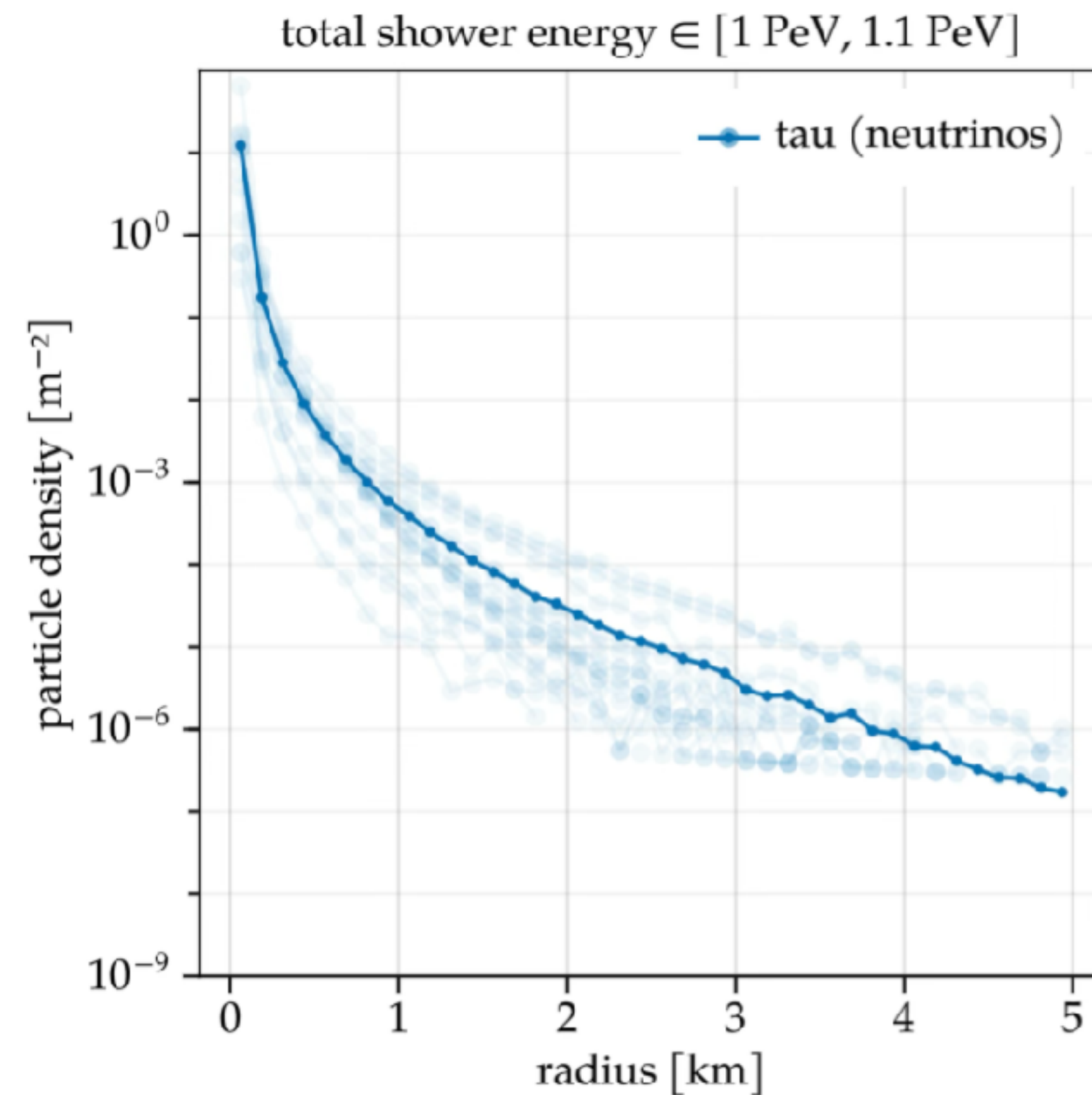
“scintillator-like”

Upshot: big impact on CRs, ν_τ less affected

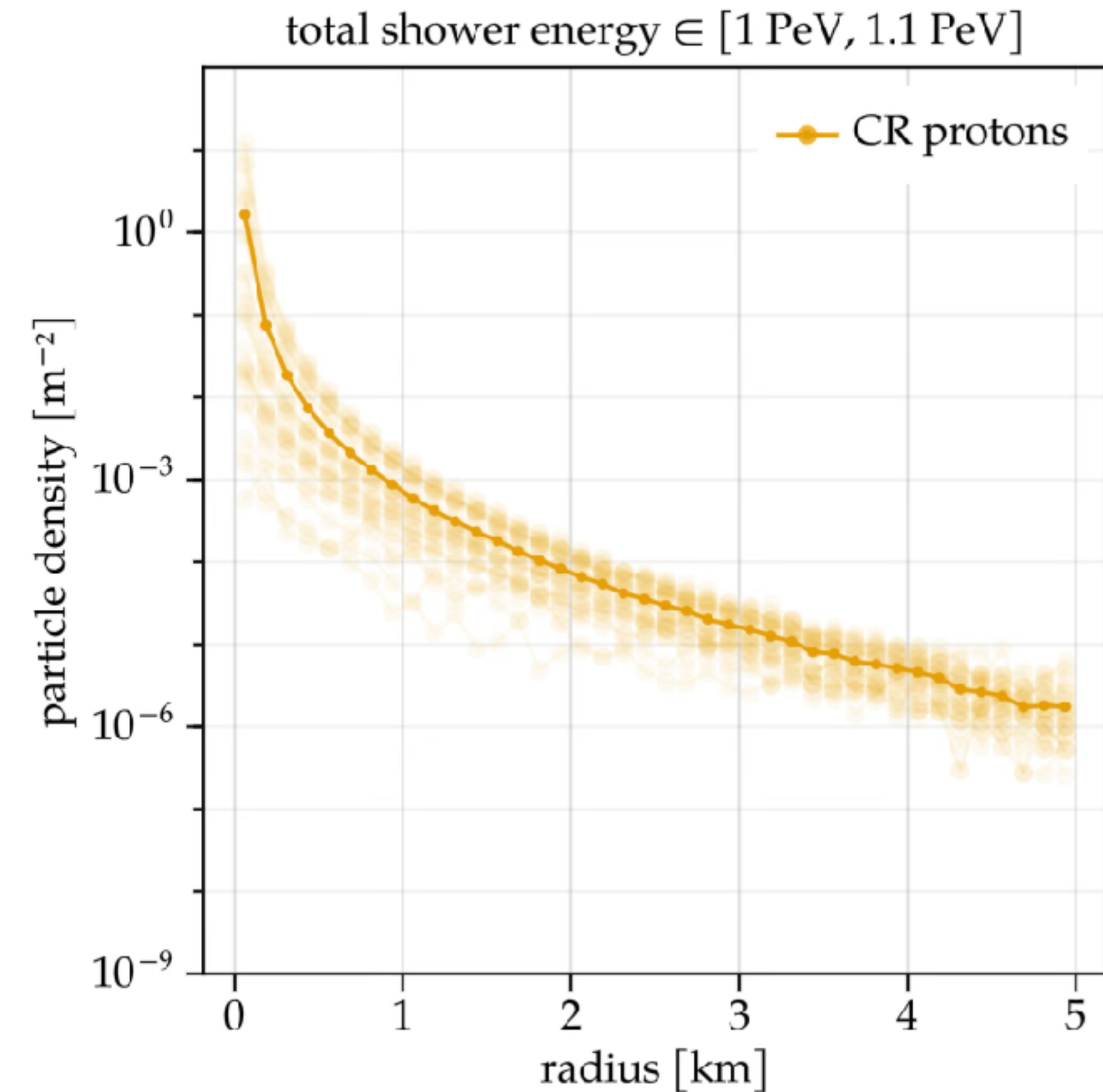


Tau showers have greater charged particle density than CR showers

* charged particles only



Average # in $r < 125\text{m}$: $12/\text{m}^2$



Average # in $r < 125\text{m}$: $1.6/\text{m}^2$

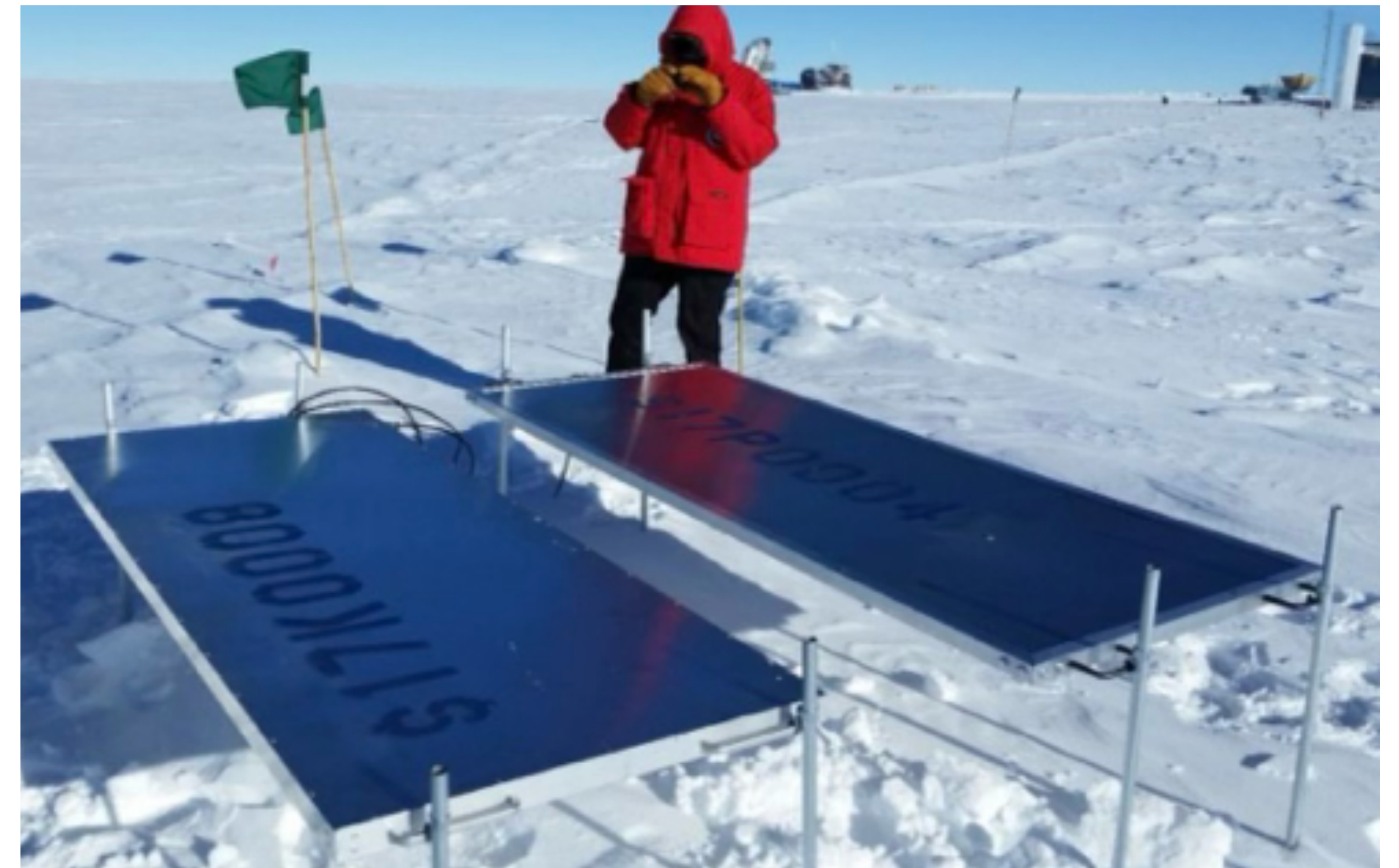
Opto-electrical system

Two main types of air shower detectors. Which should we use?



Water tanks:

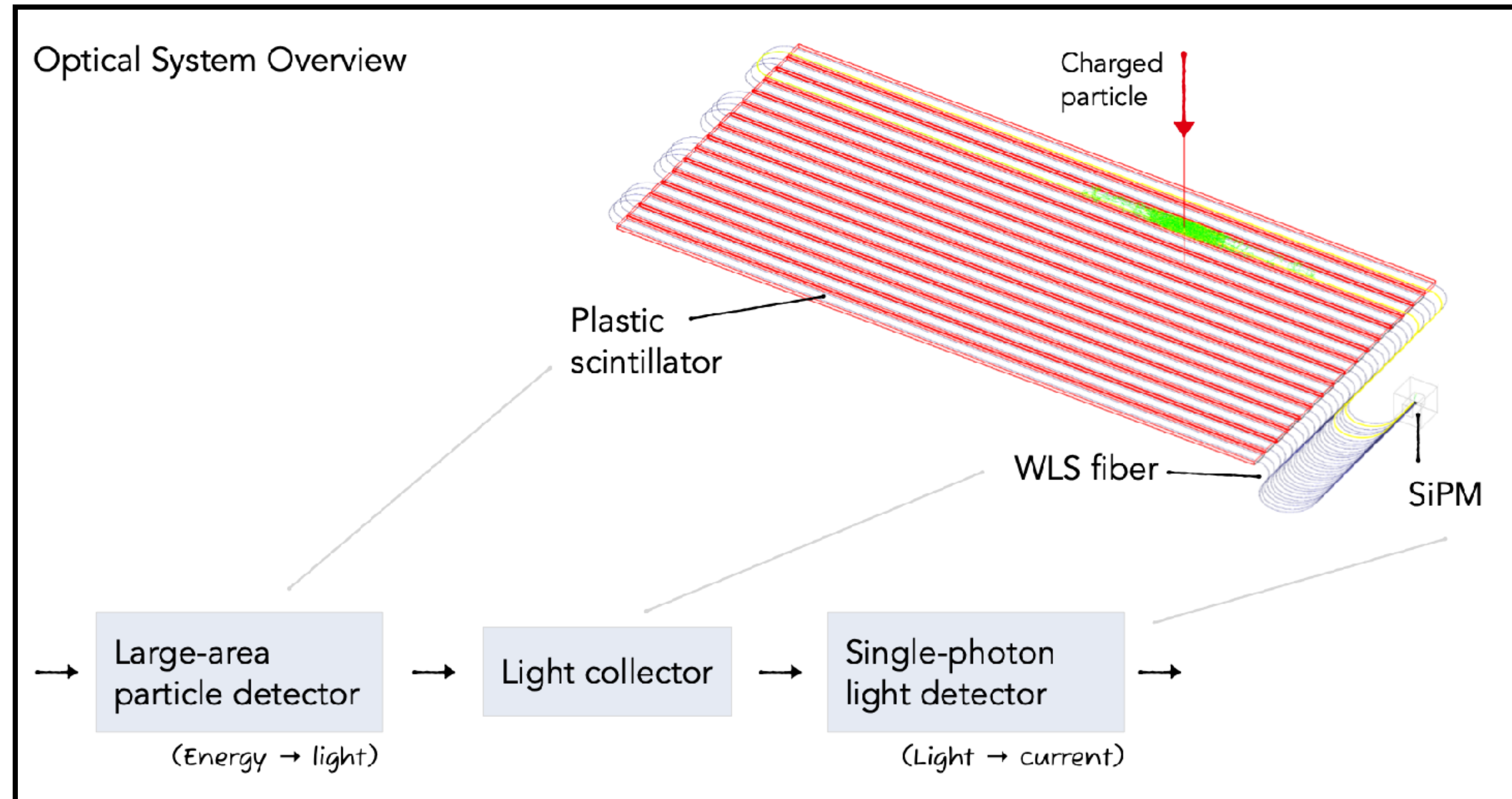
- Highly gamma-sensitive (most abundant particle)
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Plastic scintillator panels

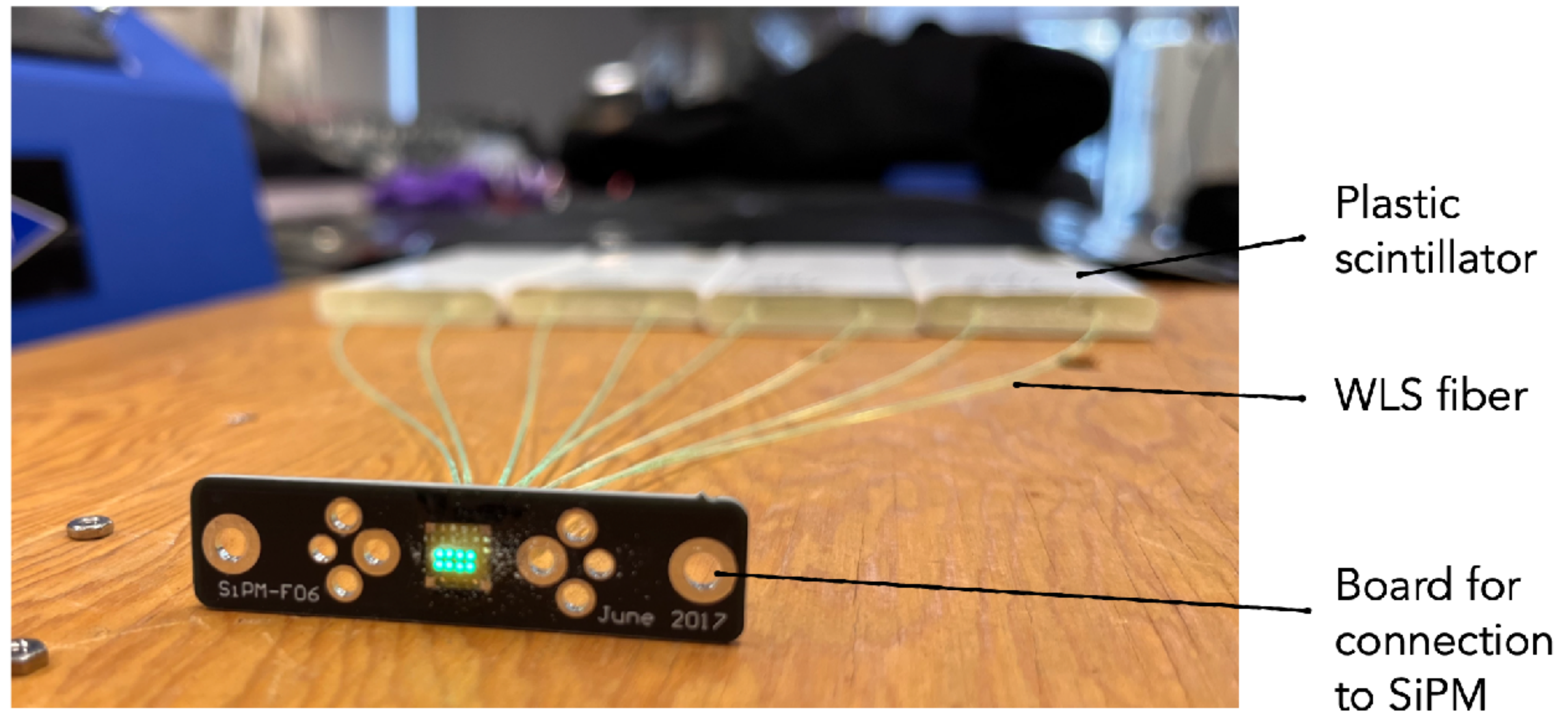
- Sensitive mostly to charged particles
- Can be built off-site and transported
- Easier to deploy in canyon

How does a scintillator panel work?



Stolen from Kiara

How does a scintillator panel work?

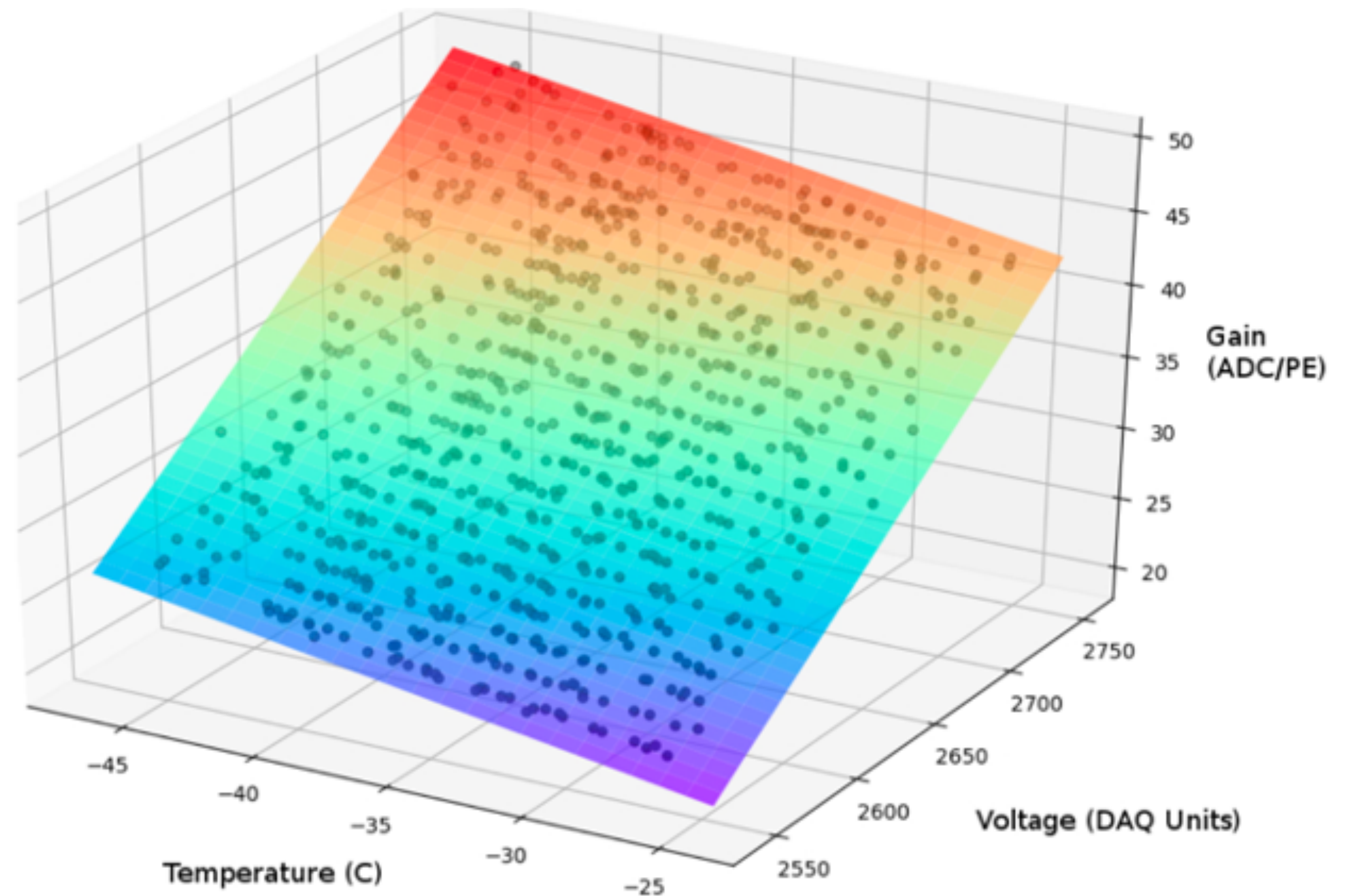


Scintillator calibration

Key requirement: must be able to calibrate detectors *in situ*

Why?

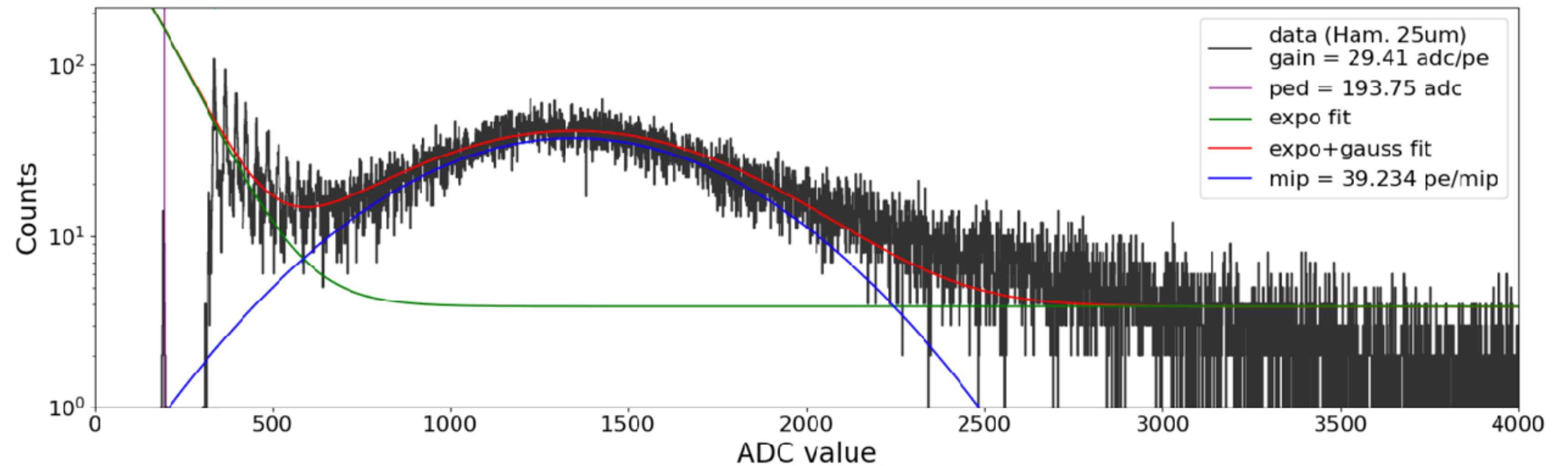
- SiPM gain known to vary with temperature
- Will be corrected in real-time via bias voltage adjustment, but want to verify this correction



Scintillator calibration

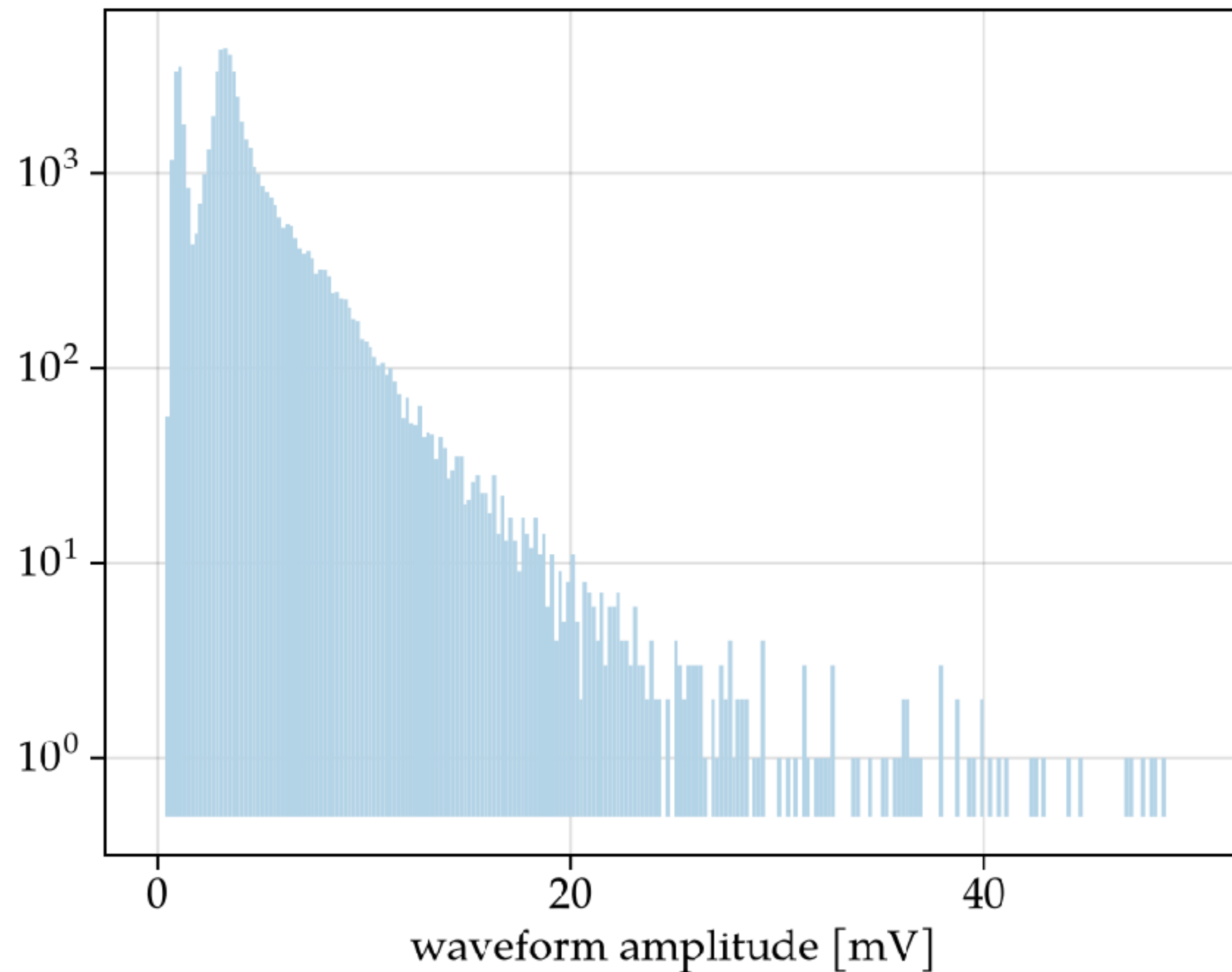
Essentially only way to calibrate *in situ* is with muons

Detectors must be designed so that we can see the muon peak



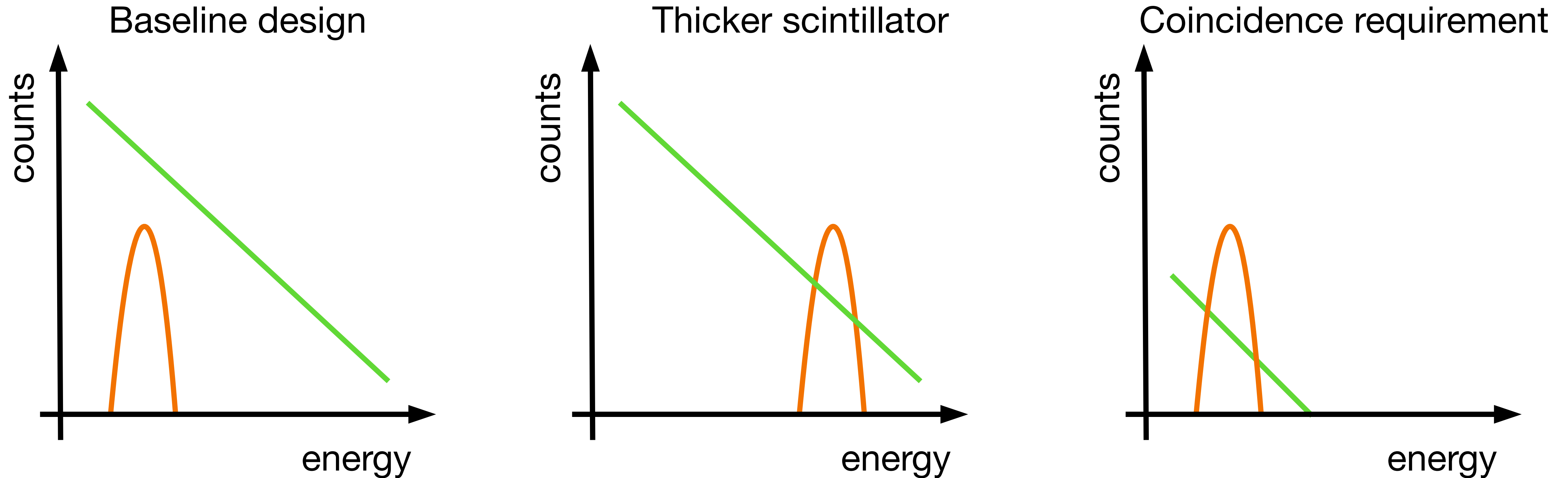
Example muon peak from IceTop panel

What do we see with one “baseline-style” detector?



Not much, actually
What's the deal?

Radioactive backgrounds can mask muon signal



If radioactive backgrounds are high they can mask our muon peak!

Prototype concepts: improving μ peak visibility

Concept #M1

- Stacked, fully overlapping detectors
- Use coincidence between detectors to extract muon peak
- Pros: simple, only one style of detector, increased thickness also increases sensitivity to tau showers
- Cons: heavier, more expensive, additional electronics



Prototype concepts: improving μ peak visibility

Concept #M2

- Stacked, partially overlapping detectors
- Use coincidence for muon peak, but partial overlap gives more detector area
- Pros: simple, one detector style, more area compared to M1
- Cons: more expensive, additional electronics, inconsistent response across detector



Prototype concepts: improving μ peak visibility

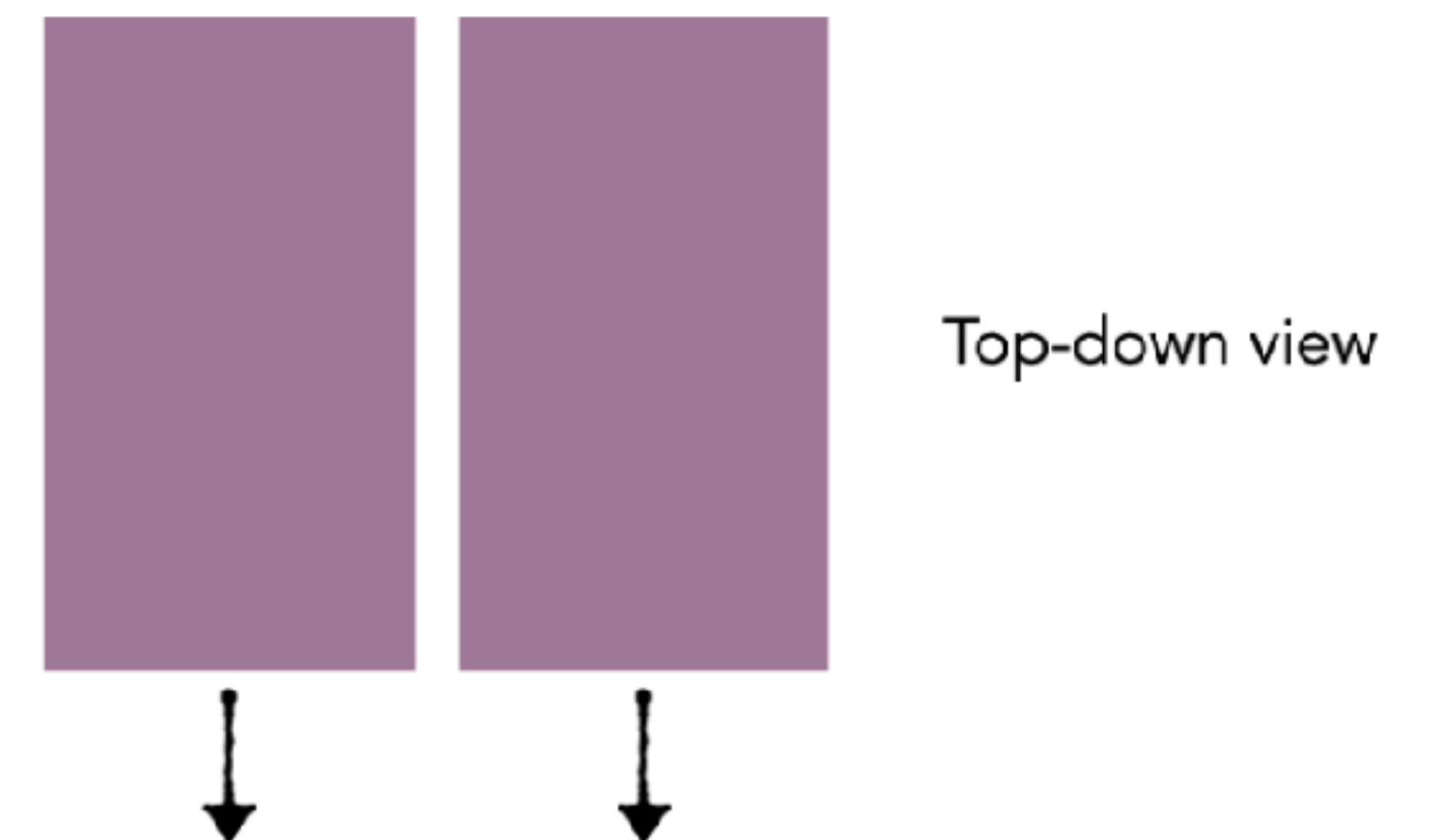
Concept #M3

- Small coincidence tagger on top of scintillator
- Same as first two concepts, but small scintillator saves money
- Pros: light, cheap



Concept #M4

- Adjacent detectors
- Tag coincident muons from same shower
- Pros: no loss of area from overlap
- Cons: lower rate?

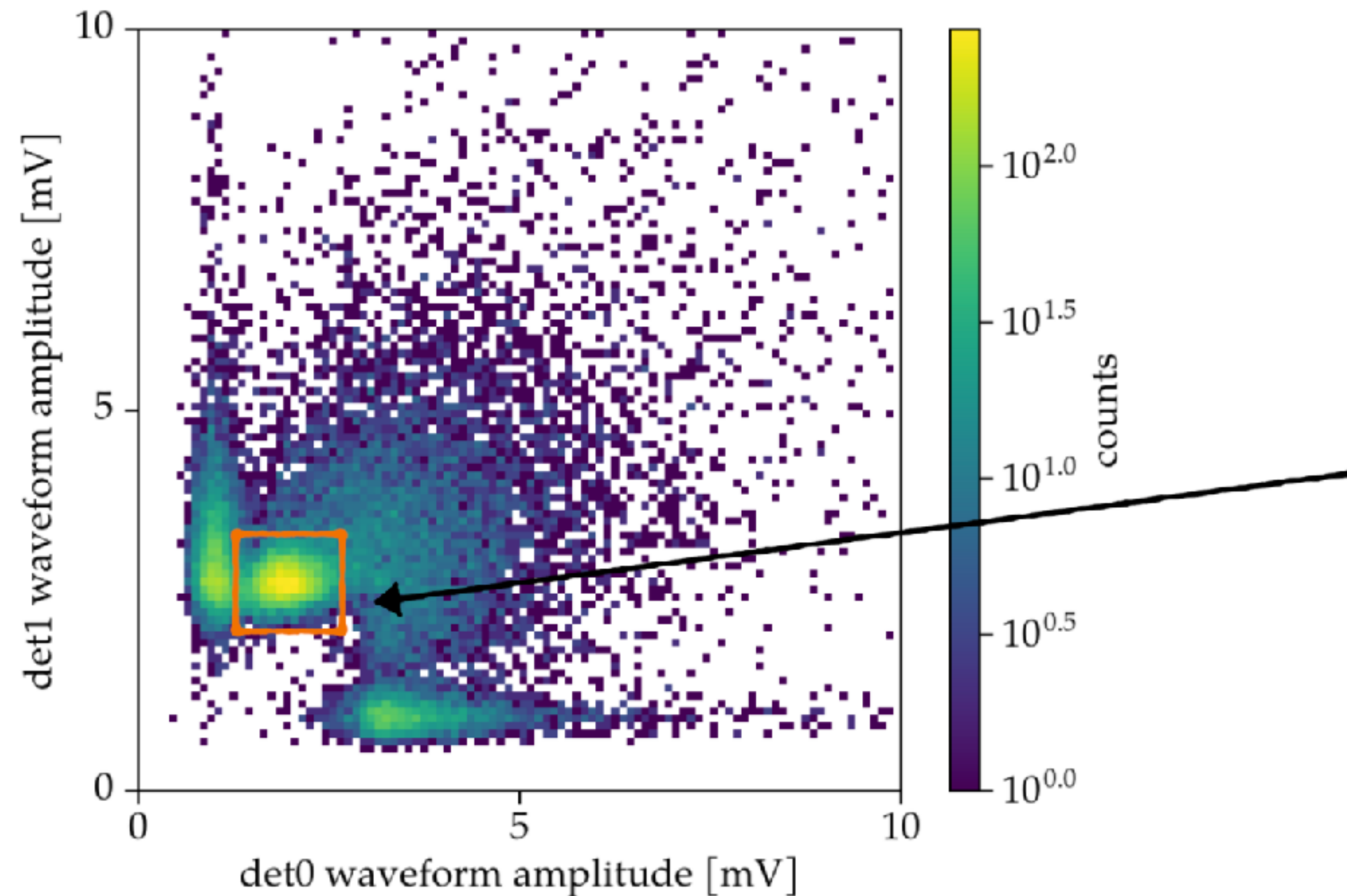


Prototype concepts: improving μ peak visibility

Concept #M5

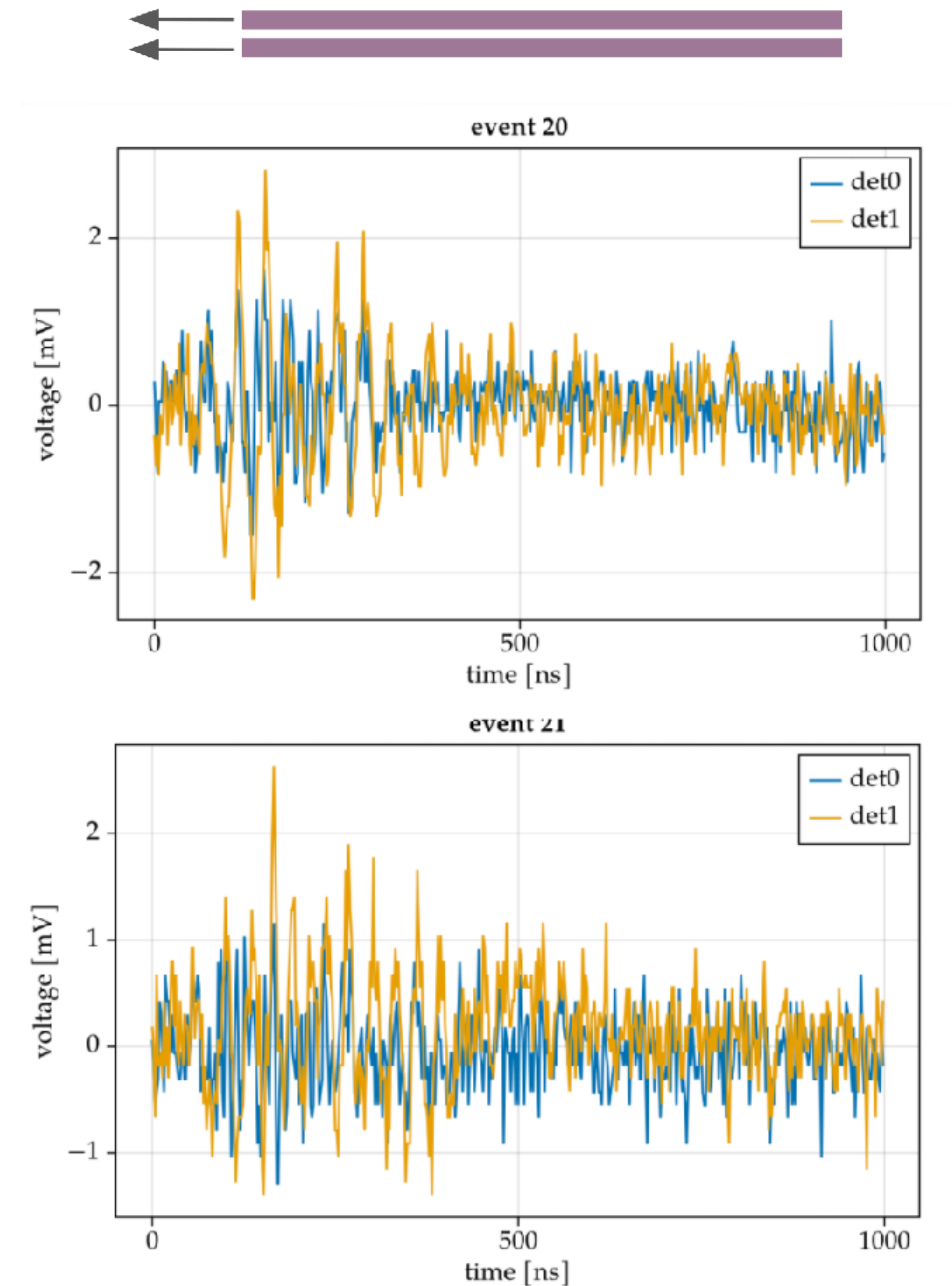
- Thicker (2 cm) scintillator
- See muon peak in single detector (?) with nothing extra
- Pros: simple, increased sensitivity to taus
- Cons: more expensive, heavier

Concept #M1

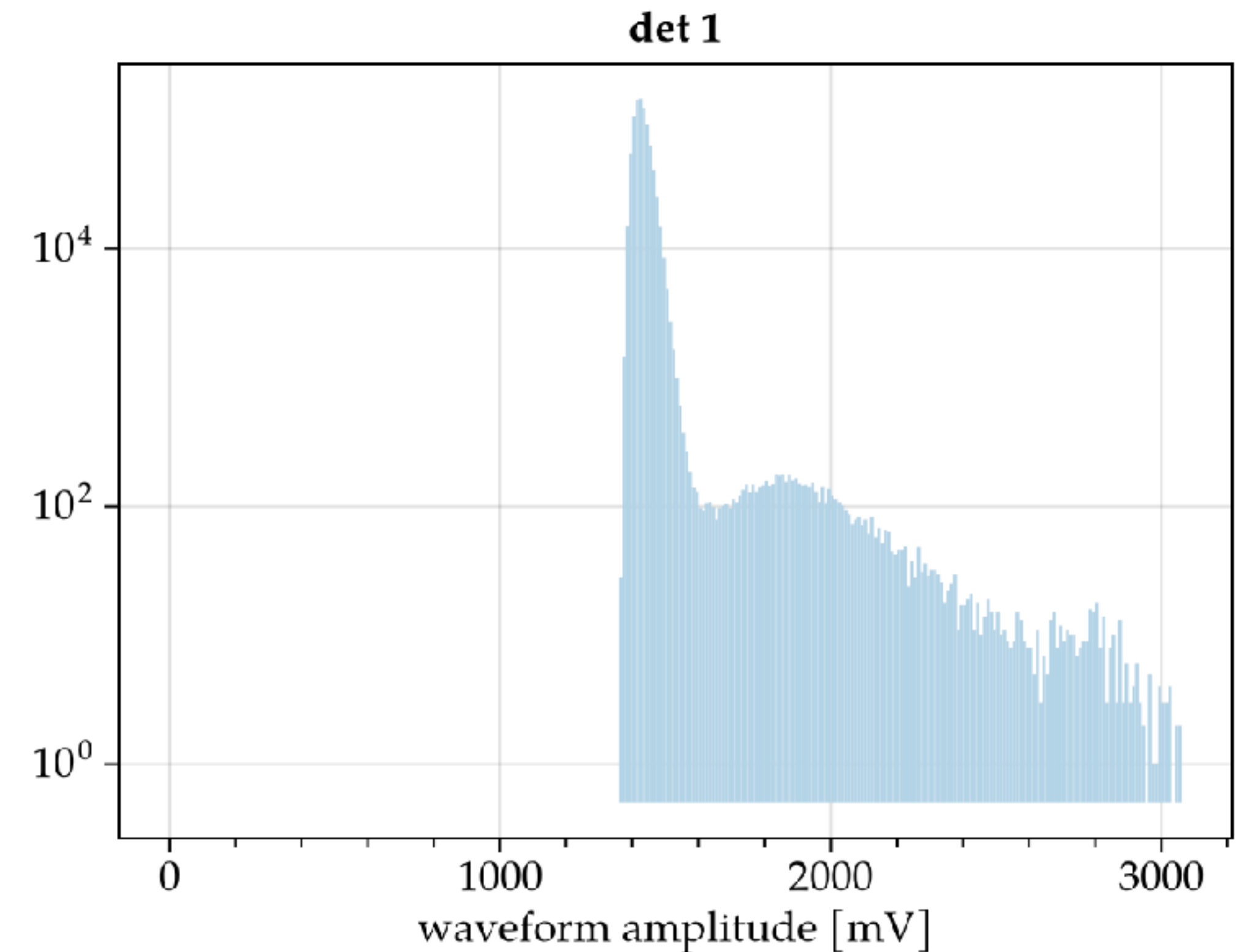
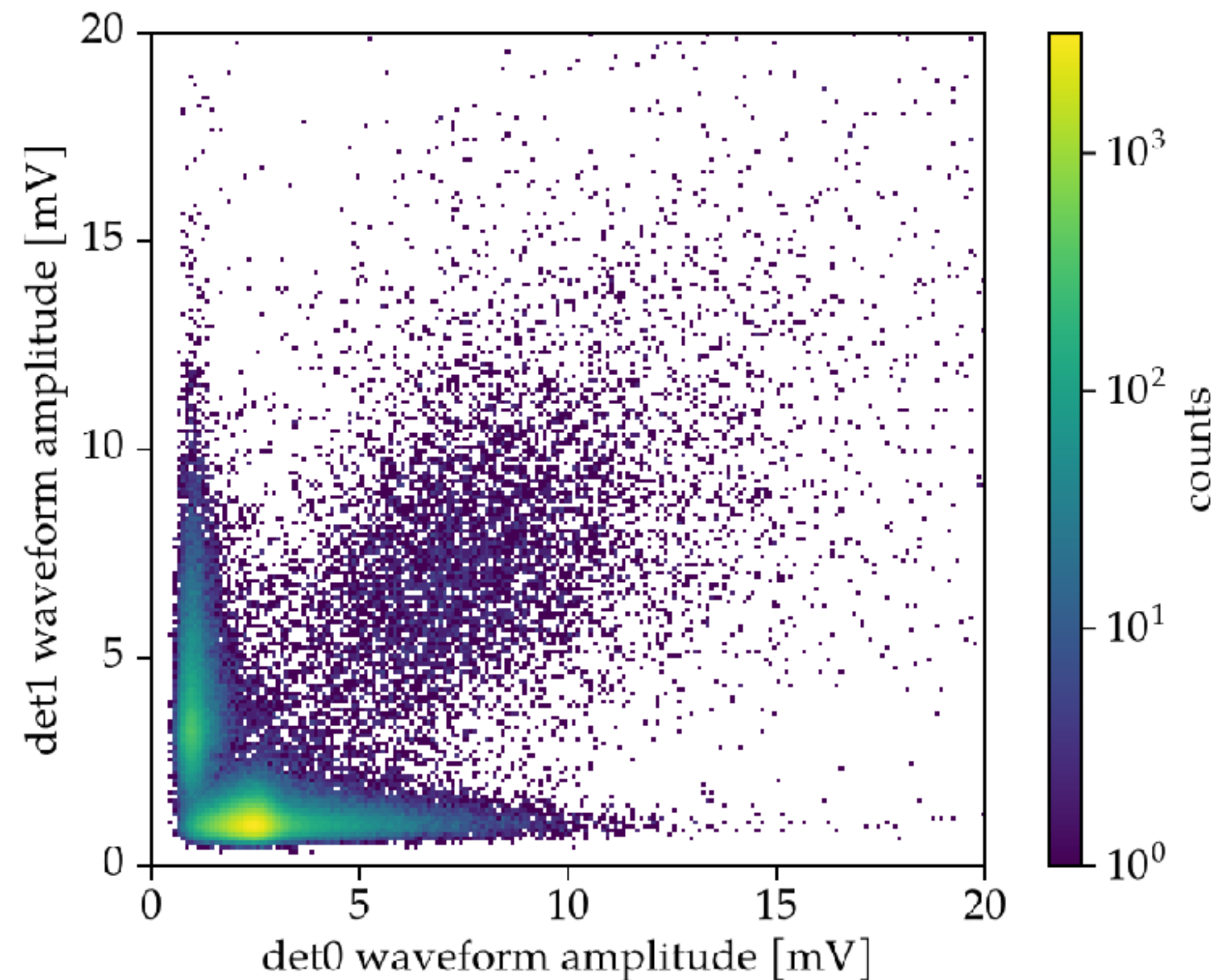


Run contaminated with EMI from scanner robot(!)

Nonetheless, can still see it's difficult to distinguish muon peak from backgrounds



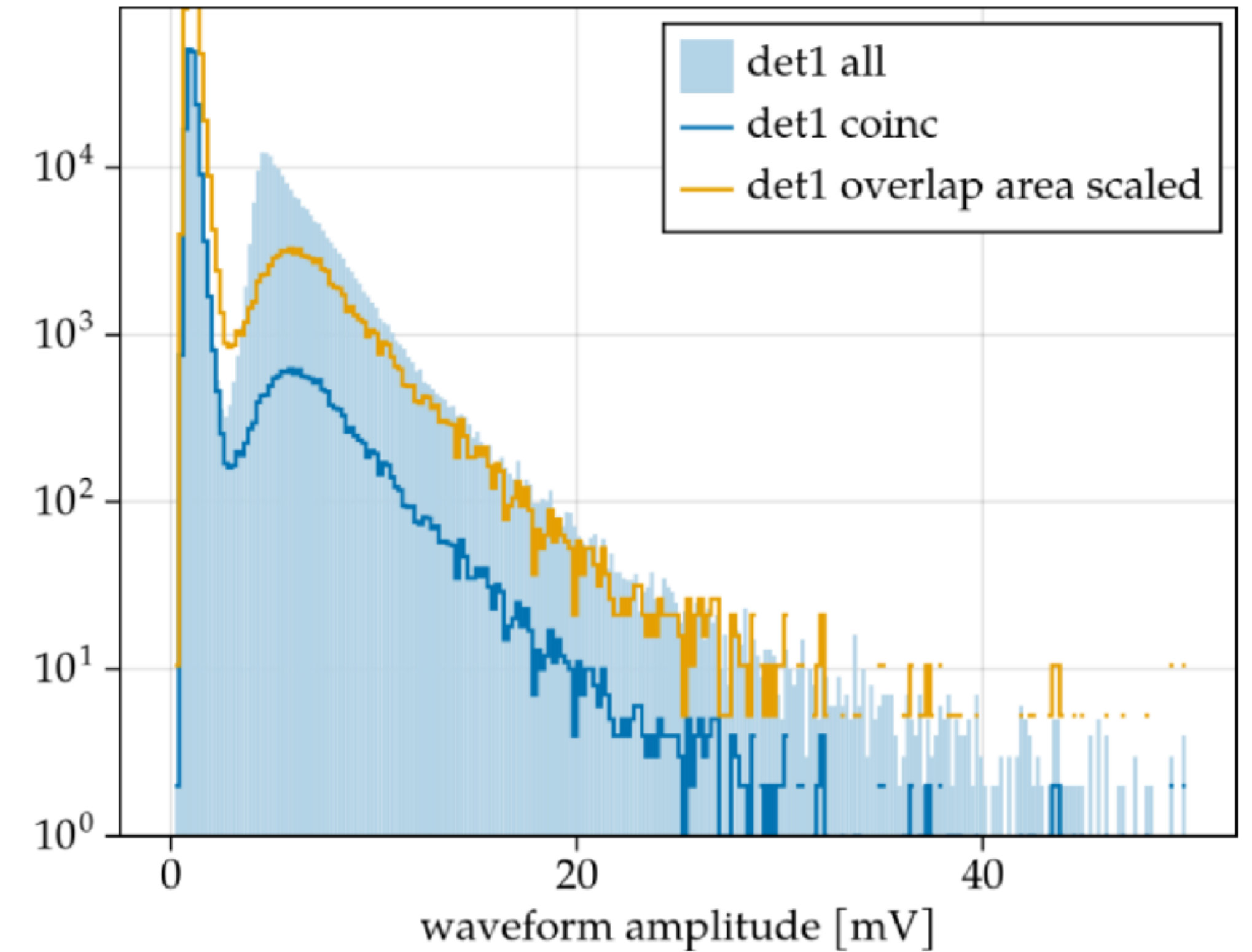
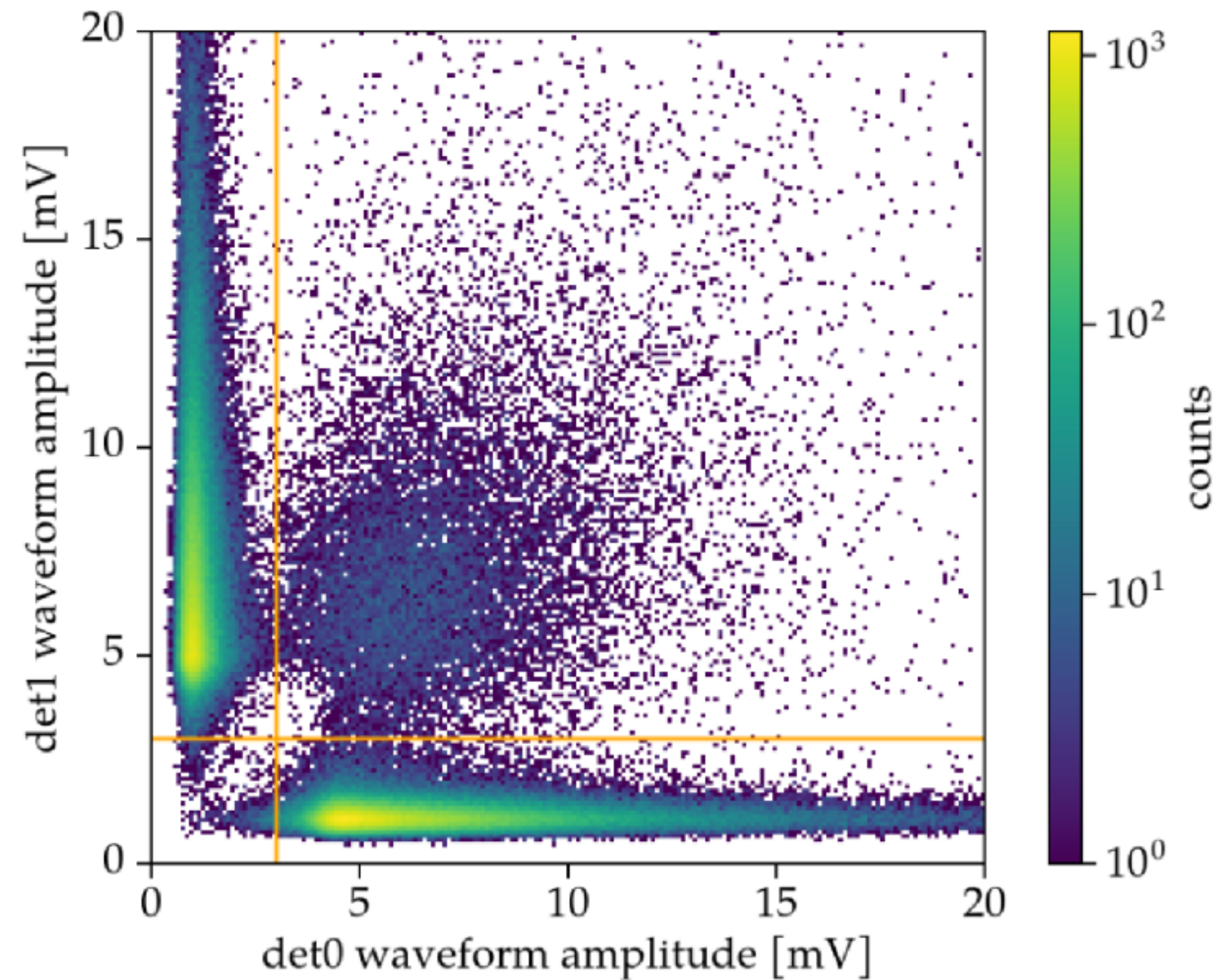
Concept #M1



At least that's what we thought before we made this measurement with half-length panels

Maybe due to less attenuation in fiber? TBD

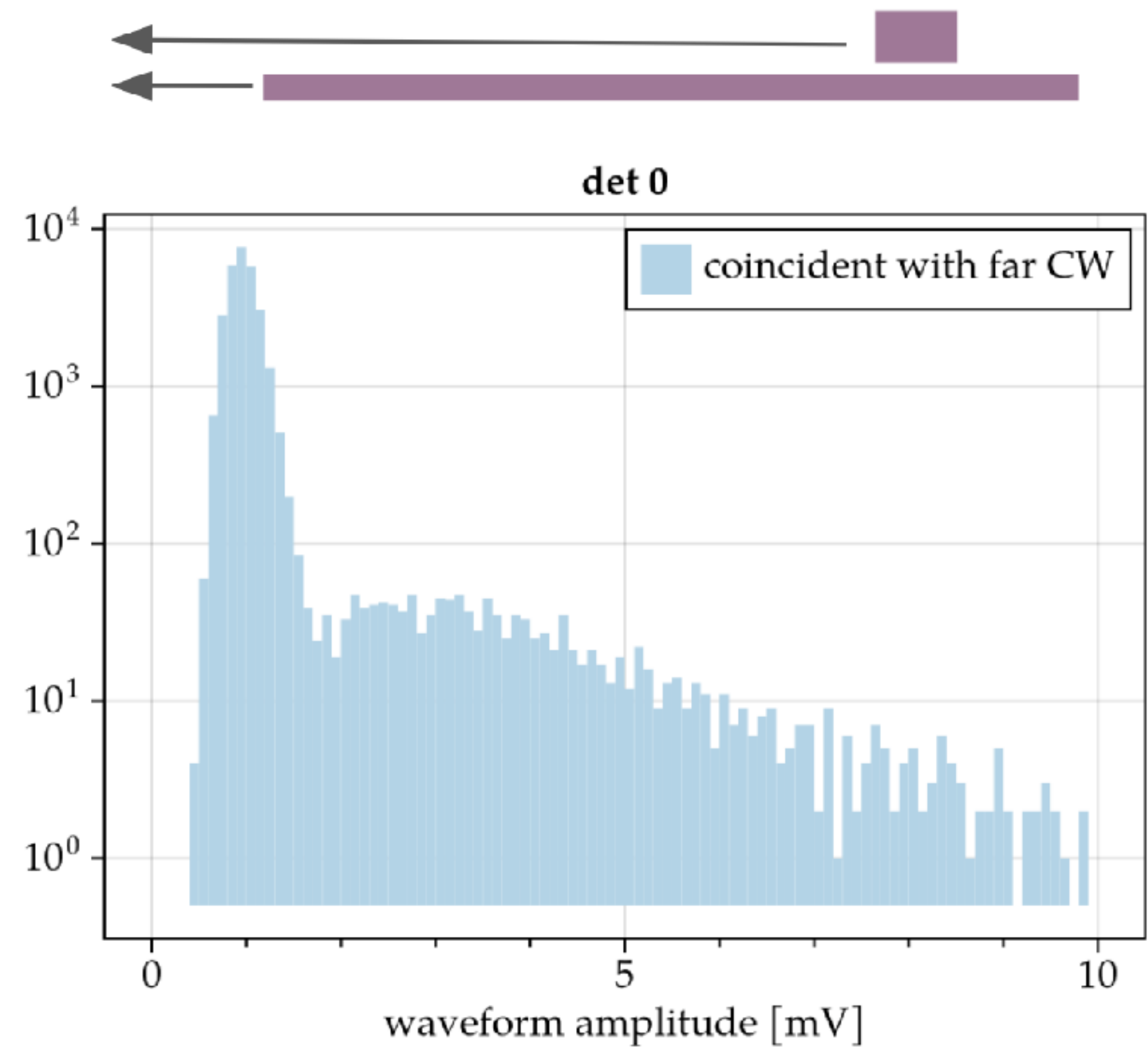
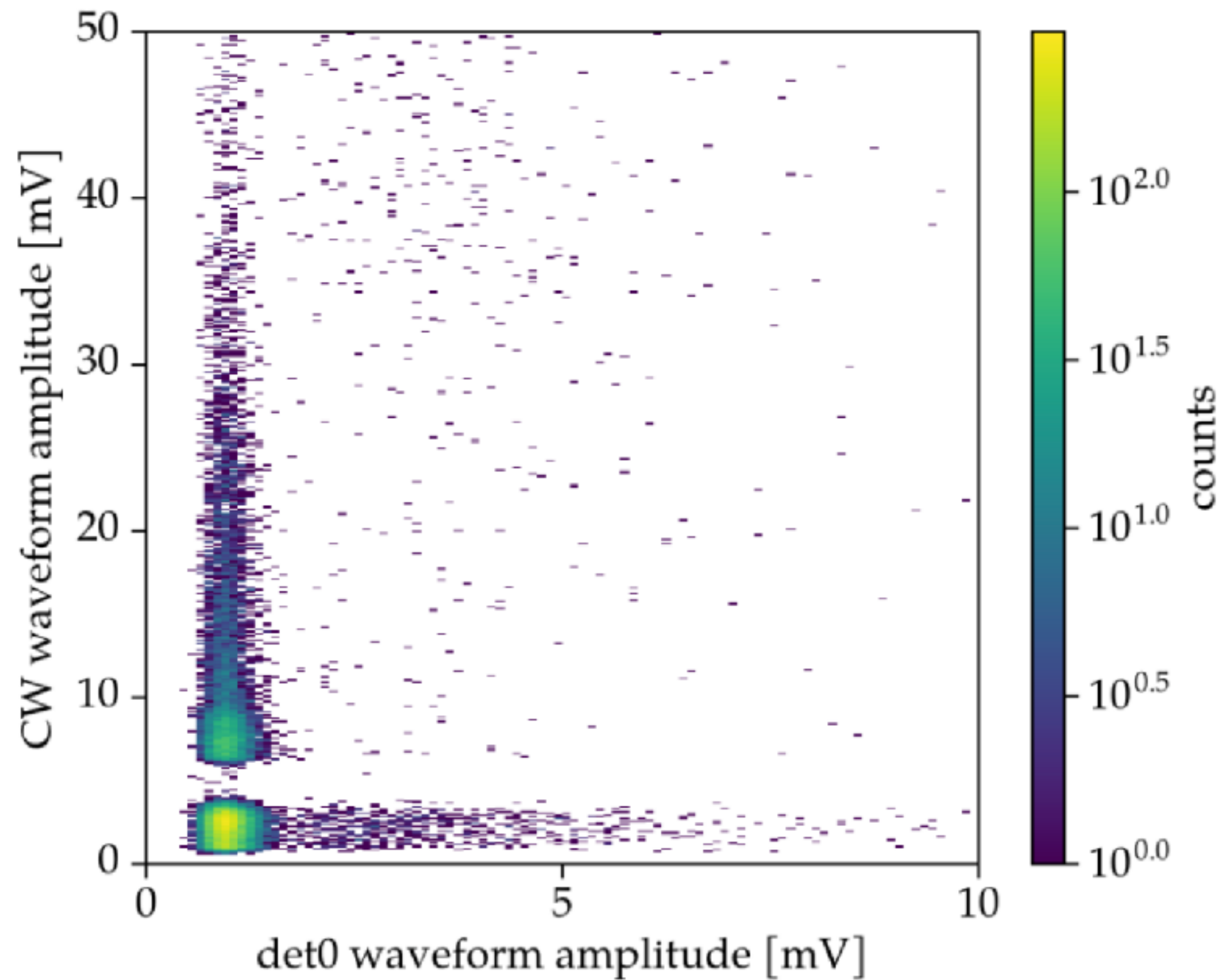
Concept #M2



Separation between muons & background is meh

Seems to work as a proof of principle though

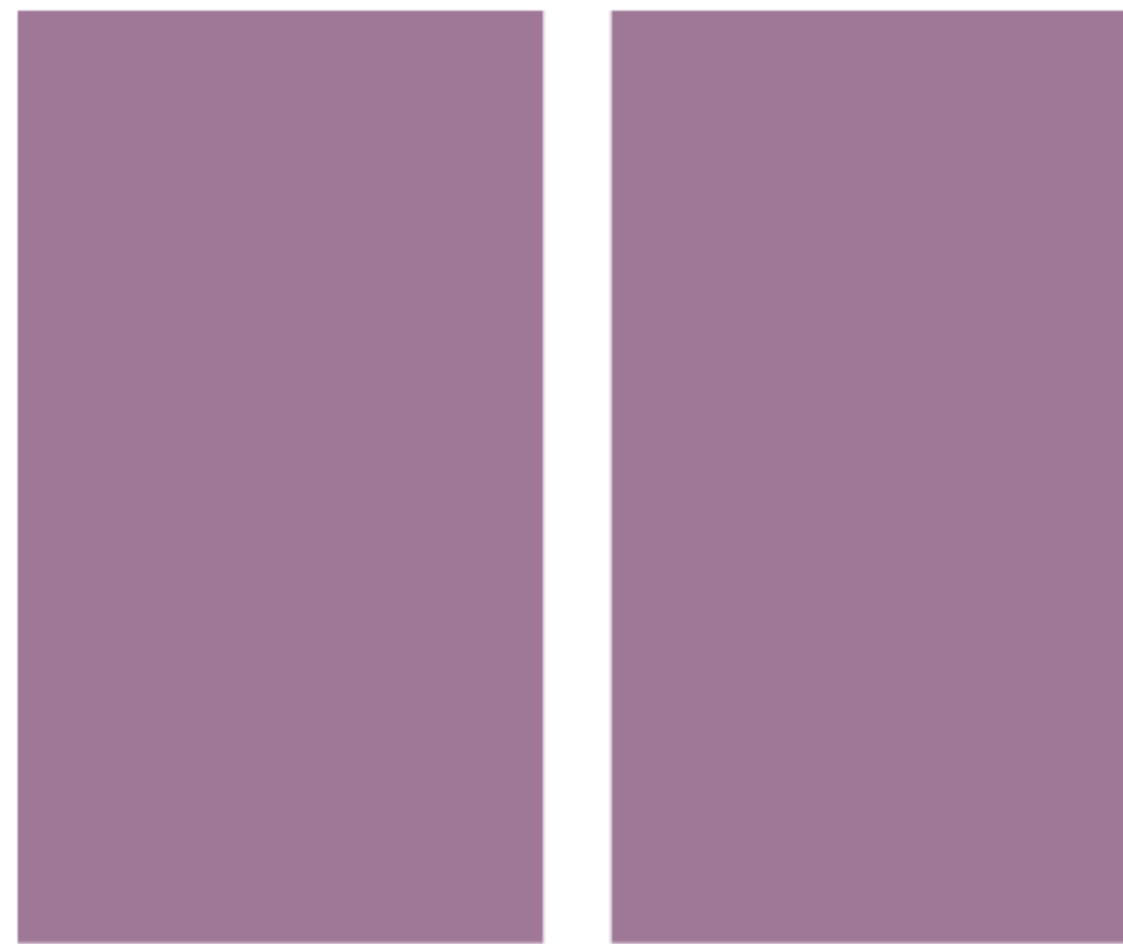
Concept #M3



Higher light yield of CW makes it easier to pull out muon peak

Should try again with half-length detector

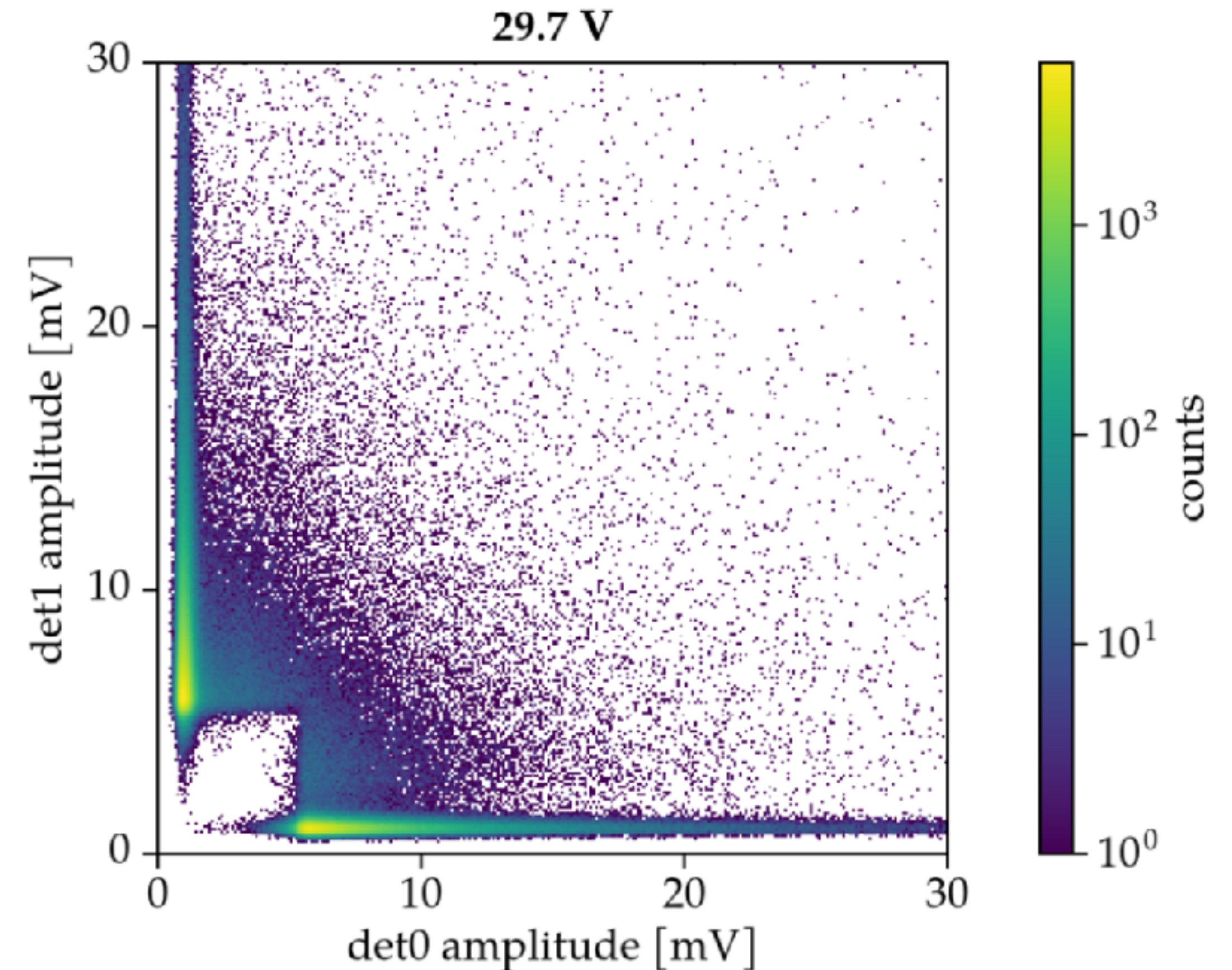
Concept #M4



Ch1

Ch0

Top-down view



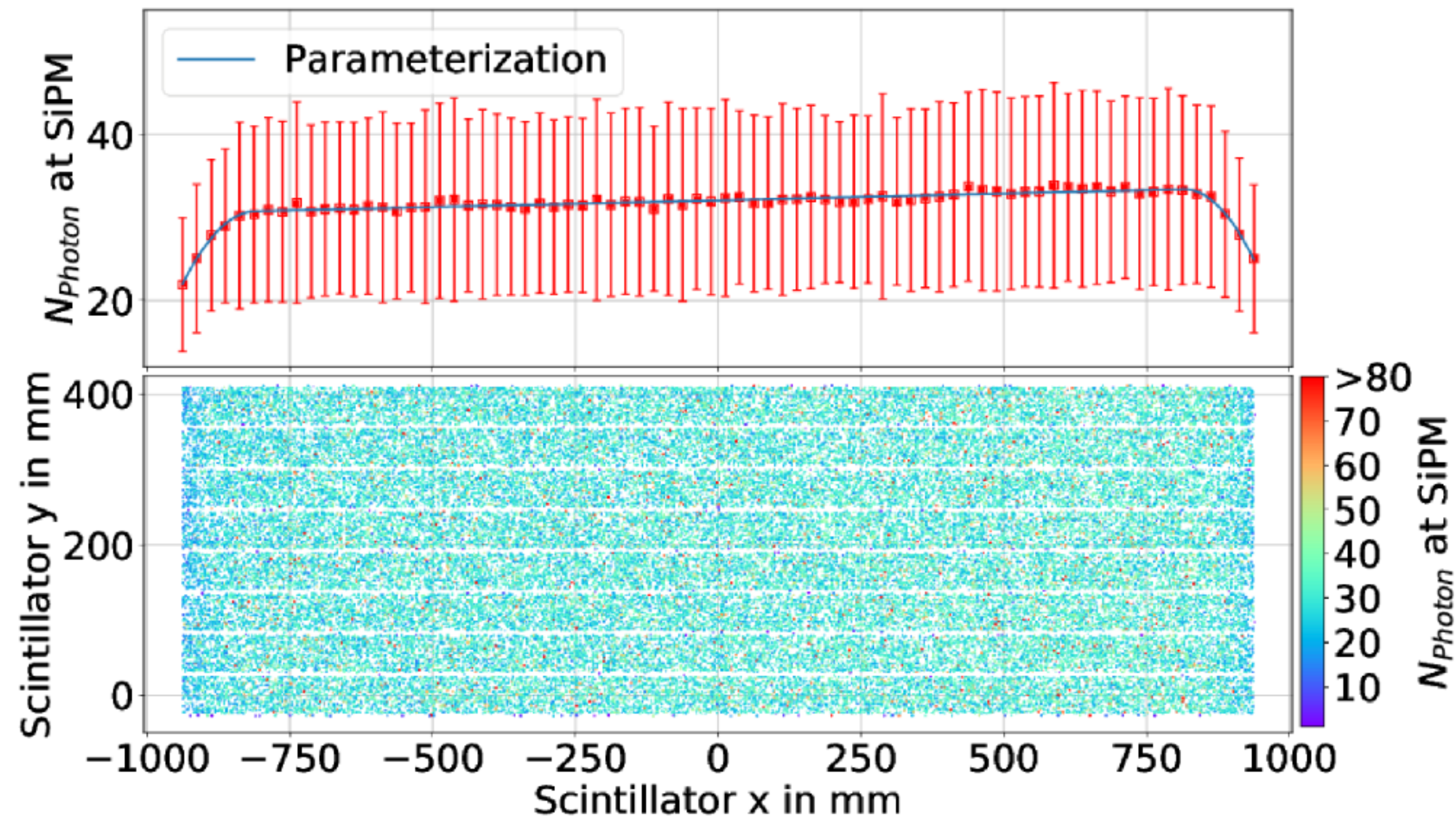
Not a very clear peak here, maybe we need to be outside?

Better evaluation of this in progress

Concept #M5

To be tested (maybe)

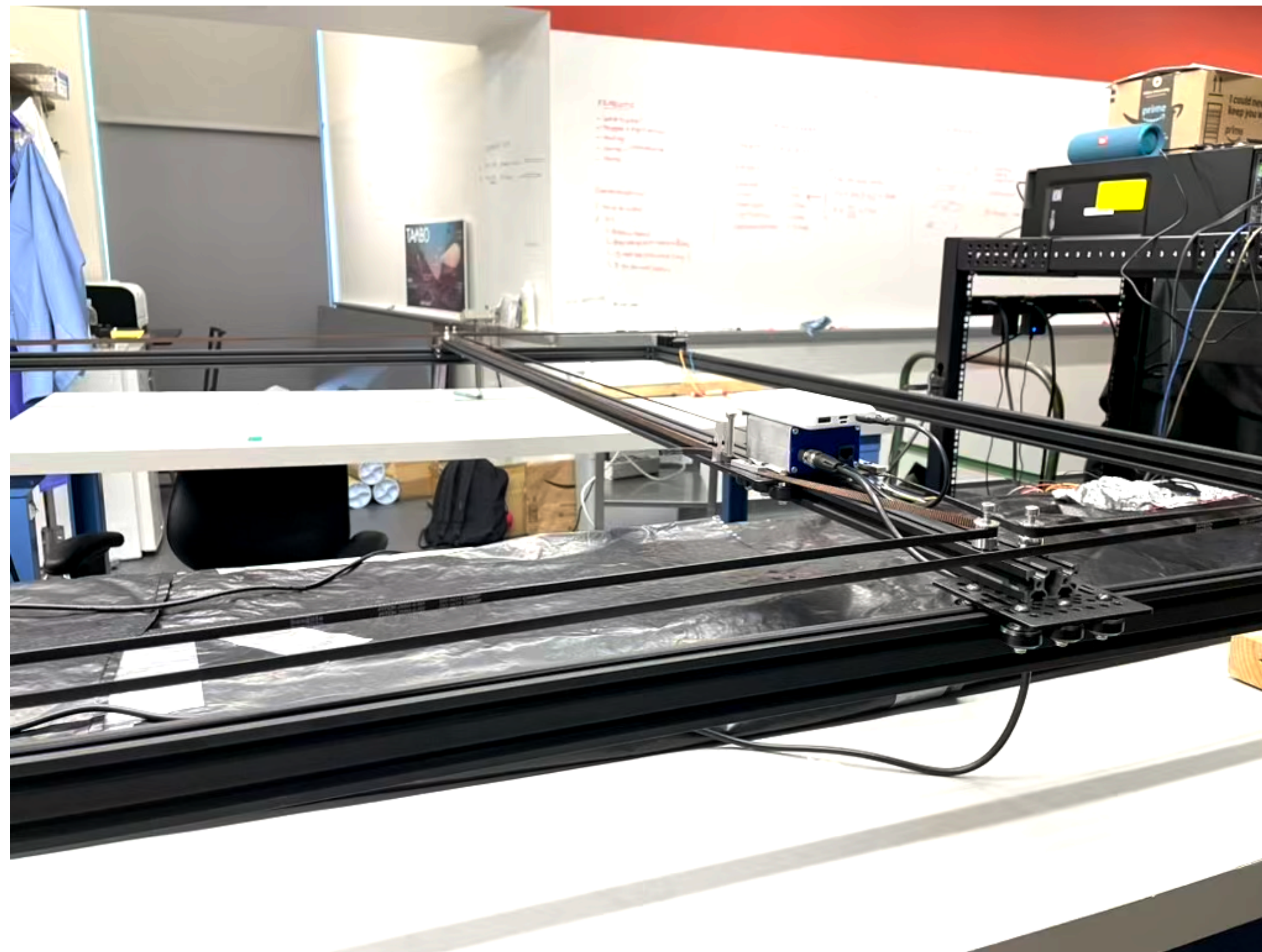
Prototype concepts: optimizing scintillator length



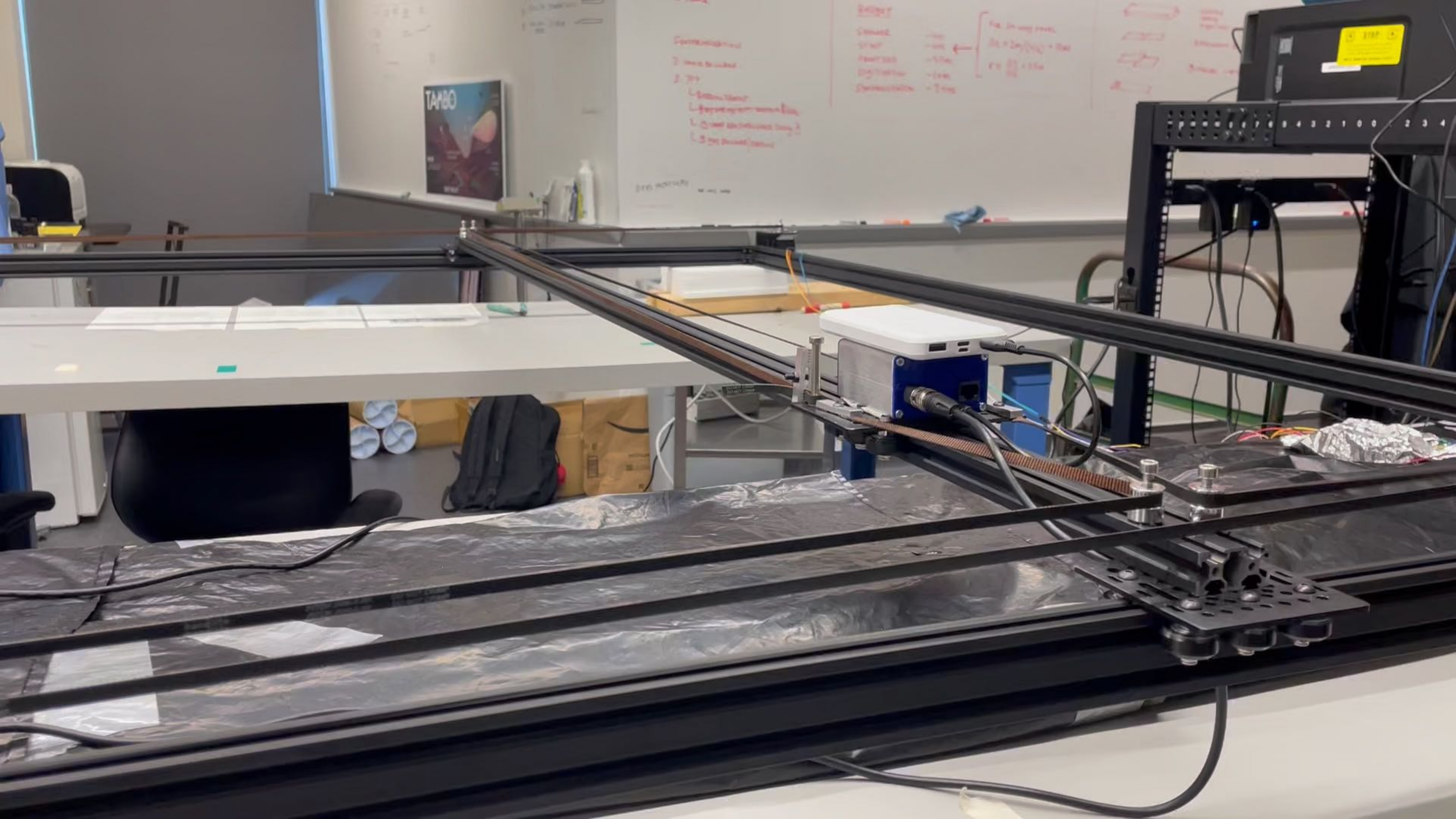
Some approximate ratios:

-934mm / +771mm	~31.5%
-778mm / +771mm	~7.5%

TAMBO Scanning Robot



Ashley (summer undergrad) build a CoreXY scanner for detector characterization



¿Preguntas?

