

Can GRAND and TAMBO share a mountain?

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UNIVERSITY OF
COPENHAGEN

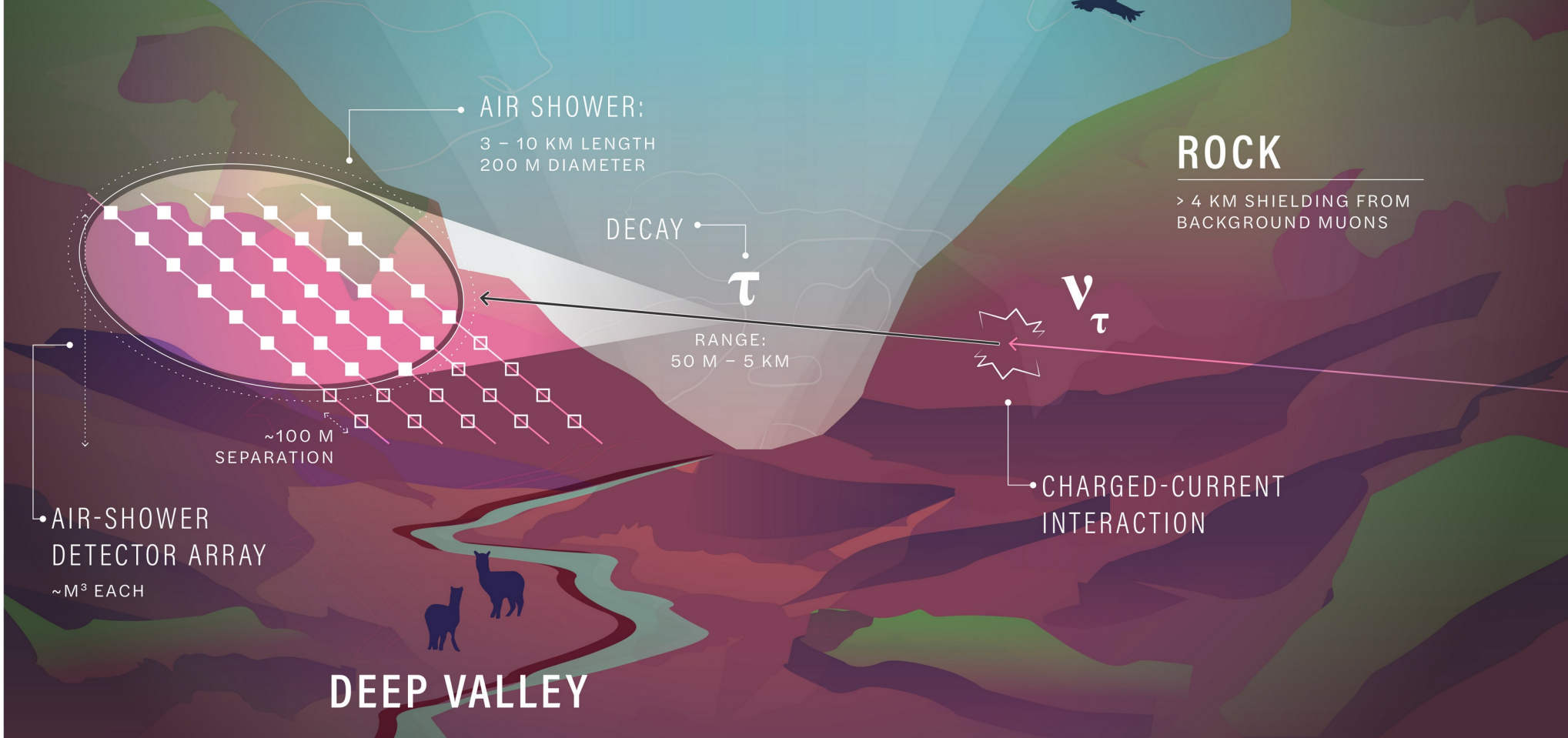


VILLUM FONDEN

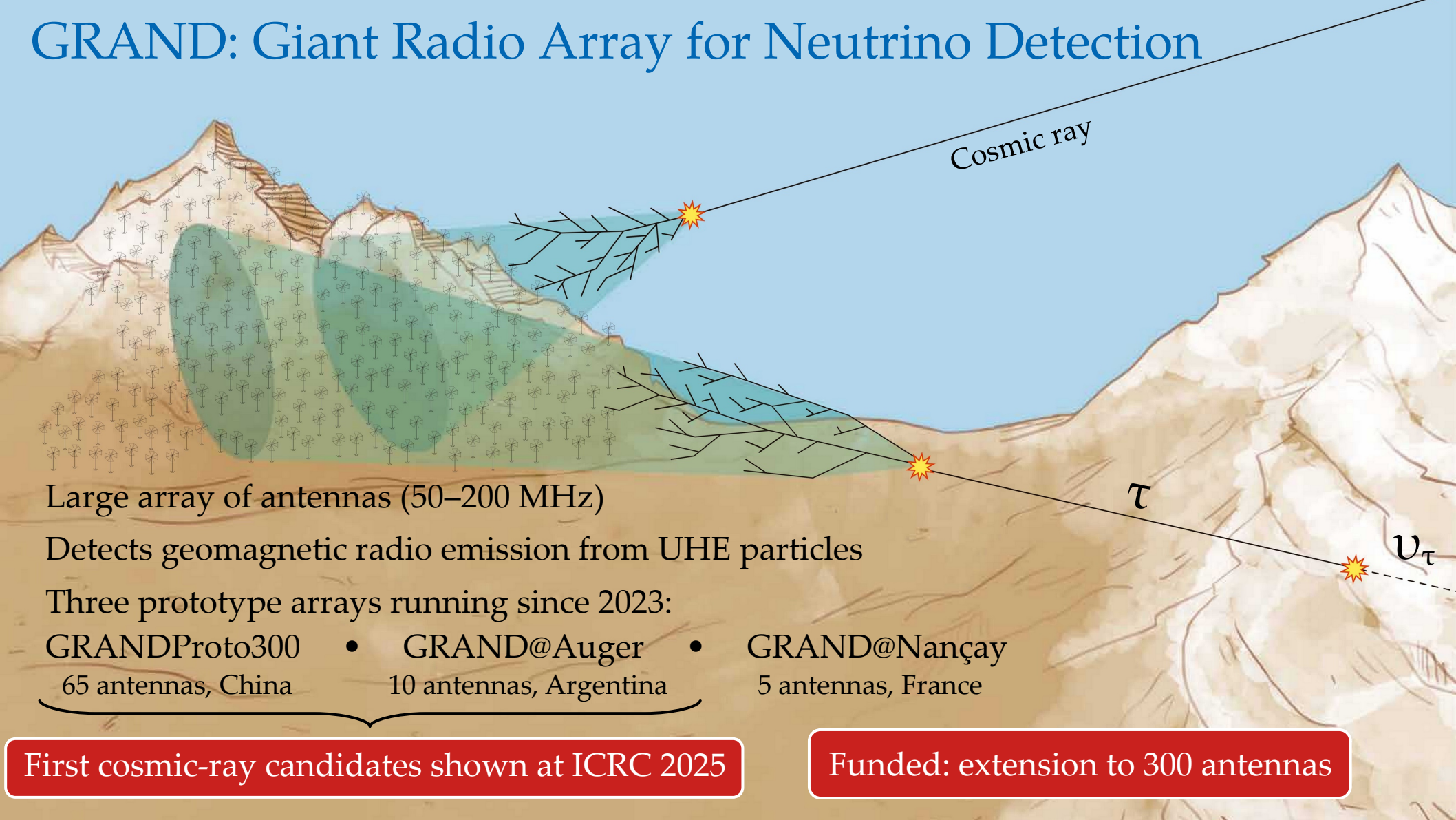


Why put two
detectors
on the same mountain

TAMBO: the Tau Air-Shower Mountain-Based Observatory



GRAND: Giant Radio Array for Neutrino Detection



Cosmic ray

τ

ν_τ

Large array of antennas (50–200 MHz)

Detects geomagnetic radio emission from UHE particles

Three prototype arrays running since 2023:

GRANDProto300
65 antennas, China

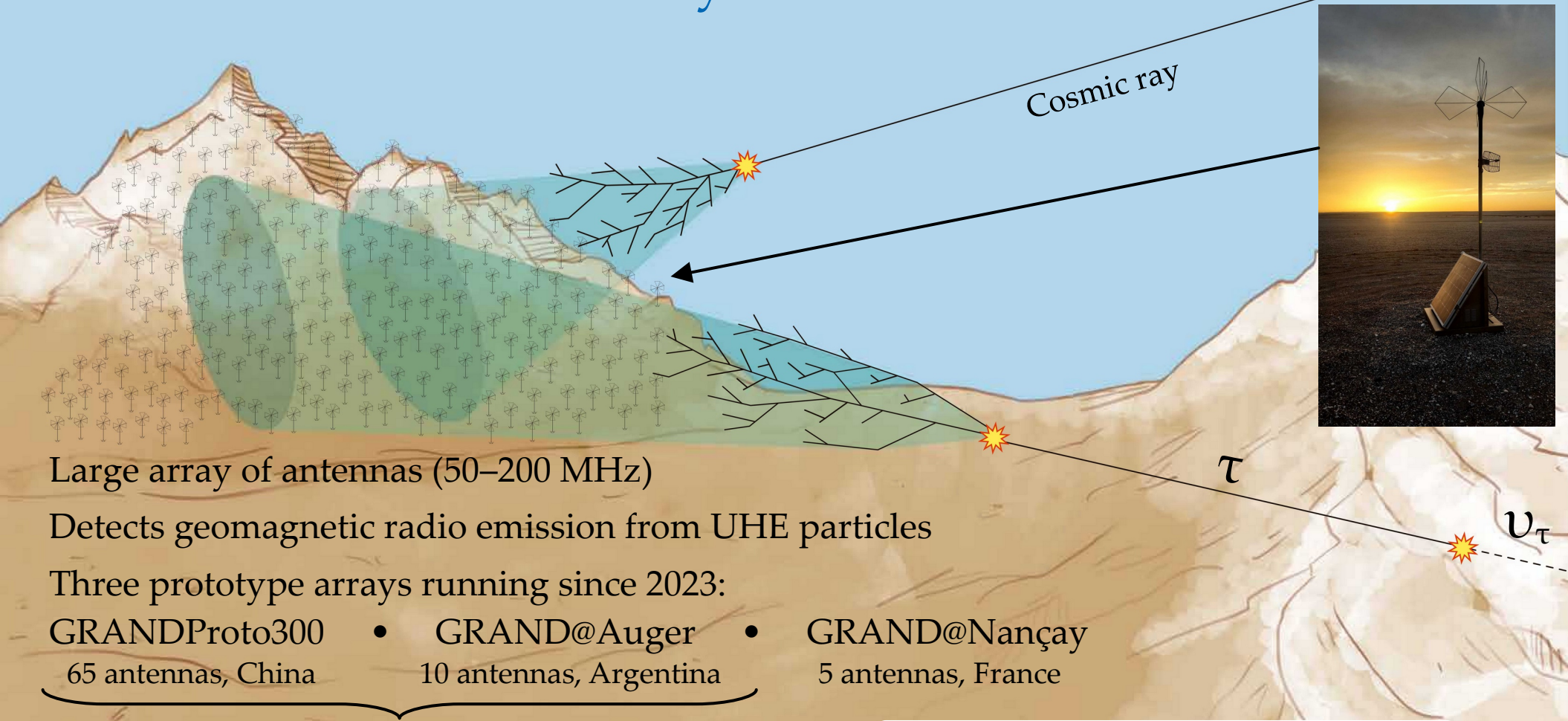
• GRAND@Auger
10 antennas, Argentina

• GRAND@Nançay
5 antennas, France

First cosmic-ray candidates shown at ICRC 2025

Funded: extension to 300 antennas

GRAND: Giant Radio Array for Neutrino Detection



First cosmic-ray candidates shown at ICRC 2025

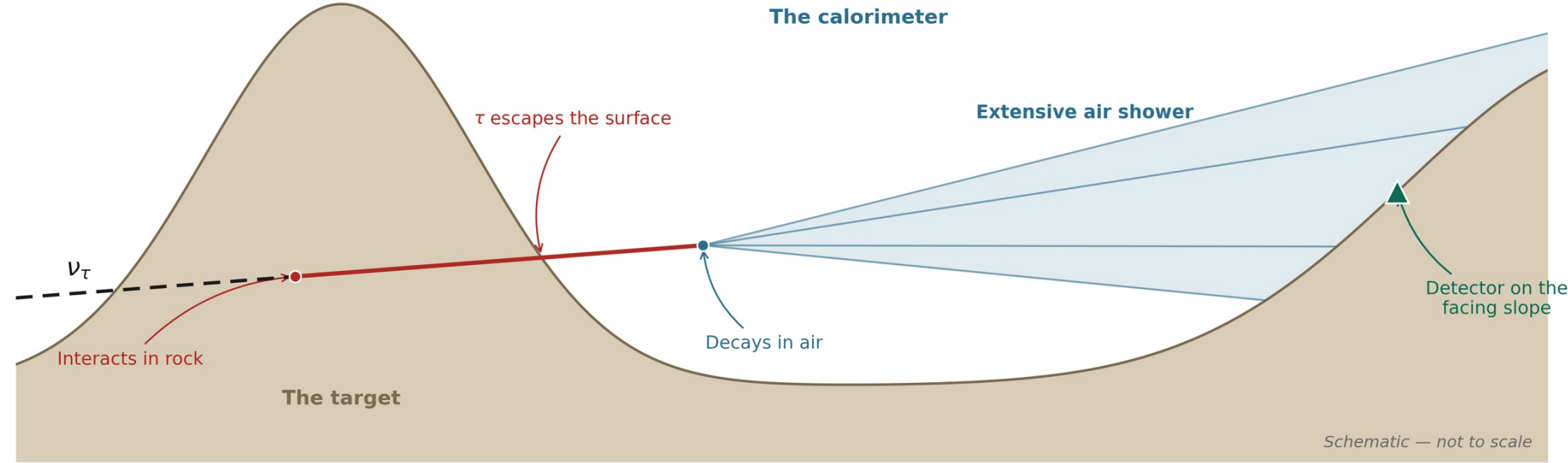
Funded: extension to 300 antennas

The messenger: Earth-skimming tau neutrinos

A ν_τ crosses the Earth, interacts in rock, and makes a τ .

If the τ escapes to the surface, it decays in air and starts an extensive air shower.

The Earth is the target. The atmosphere is the calorimeter. A mountain is the interface.



Same shower, two readings

	GRAND	TAMBO
Channel	Radio from the shower	The particles themselves
Sees of the shower	Only the electromagnetic part	All of it
Energy reach	$\gtrsim 100$ PeV	~ 3 PeV – 1 EeV

Radio reflects only the electromagnetic component of the shower; the particle detector samples all of it. Two independent channels, not one signal read twice.

Energy ranges **overlap but are offset**; together, they measure more than either alone.

Why co-locate at all?

1. Cross-calibration

An established technique beside a new one: same sky, same energy decade, two channels that sample different parts of the same shower.

2. A joint spectrum wider than either alone

PeV to EeV, from one site.

3. Shared access, power, data, permits, community relations

At 4,000 m in the Andes these are not a footnote in the budget.

4. Hybrid events

If the arrival directions and energies genuinely coincide.

Why Peru: each experiment's requirements

	GRAND needs	TAMBO needs
Ground	Gentle slope, 3–25°	A canyon wall, 20–60°
Spacing	Antennas every 1 km	Detection units every 150 m
A target	Massifs 10–40 km away	A far wall 2–5 km away
A view	$\pm 3^\circ$ about the horizon	$\pm 20^\circ$

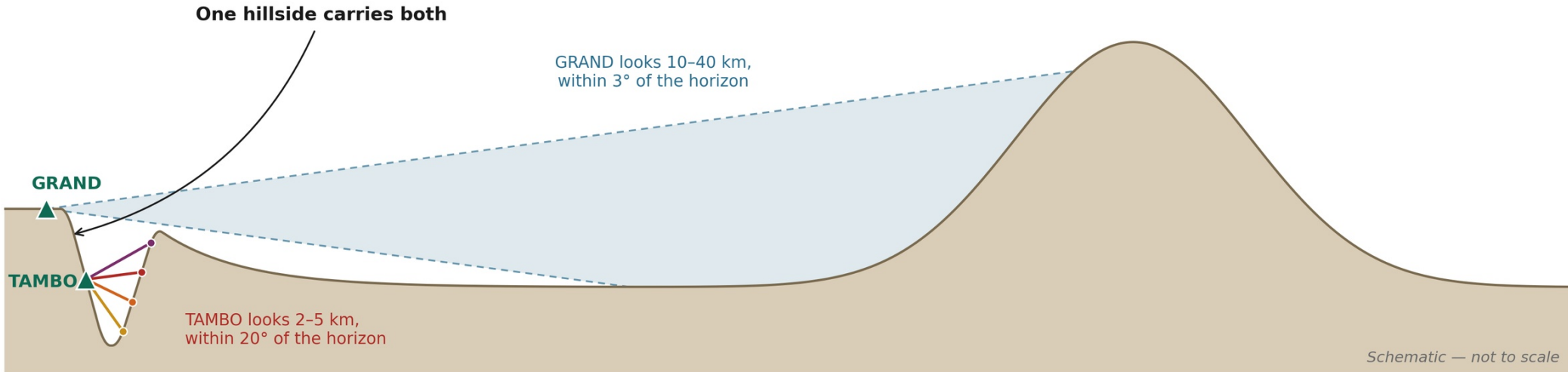
The Andes supply both: high plateau + canyons among the deepest on Earth, within tens of km of each other. Colca Canyon is TAMBO's chosen site: high, dry, sparsely populated.

The two lists name the same resource only once: the ground itself. Hold that thought.

What competes, and what does not

Standing. A patch of ground has **one slope**. GRAND builds on 3–25°, TAMBO on 20–60°. The only ground both can build on lies in **20–25°**: a five-degree sliver.

Looking. GRAND watches massifs **10–40 km** out, within 3° of the horizon. TAMBO watches a canyon wall **2–5 km** out, within 20°. Different targets, different ranges, different angles. Nothing to compete for.



Systematically
searching for
array sites

What do we want answered?

From this patch of ground, is there a **target surface** at the **right range**, in the **right direction**, at the **right relative orientation**, with the **right matter behind it**?

One engine serves both experiments: **they differ in their numbers, not their structure.**

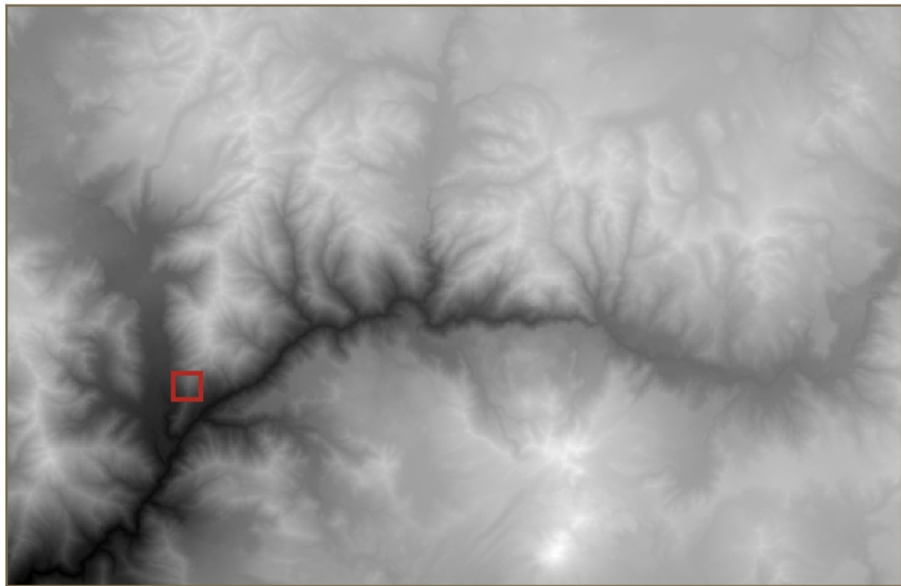
(Adding an experiment is a configuration file, not a code redo.)

The data: digital elevation model (DEM)

A DEM is a raster: a grid of pixels, each containing the elevation of the ground it covers.

We use SRTMGL1 DEM files (1 arc-second, ≈ 30 m per pixel) from OpenTopography (70–130 Mpx for each of Peru's department we explore).

(a) The Colca DEM — $3,061 \times 1,981$ pixels (6.1 Mpx)



(b) One pixel = 30 m of ground = one elevation (m)

2786	2810	2844	2868	2896	2918	2951	2973	2994
2782	2807	2849	2881	2916	2935	2949	2970	3009
2779	2812	2820	2877	2922	2940	2960	2980	3010
2777	2811	2820	2867	2894	2935	2955	2986	3014
2770	2804	2817	2846	2884	2917	2952	2987	3008
2748	2786	2817	2835	2855	2892	2926	2970	2994

Oroscope: a new tool to search for astroparticle array sites



A terrain search for tau-neutrino detectors: an elevation model in; out come the ground that qualifies, ranked sites, and the assumptions that produced them.

Install: `pip install oroscope` – then simply import `oroscope` in your code

Code: github.com/mbustama/oroscope

Docs + API ref.: mbustama.github.io/oroscope – 13 tutorial [Jupyter notebooks](#)

```
# Fetch the elevation model (a free OpenTopography key)
$ oroscope-fetch-dem --region arequipa --open_topography_api_key KEY

# Cut the Colca window out of it
$ oroscope-crop input/dem/arequipa_SRTMGL1.tif input/dem/colca.tif \
    --north -15.30 --south -15.85 --west -72.40 --east -71.55

# Run the search – the experiment is the configuration file
$ oroscope --config_path config/tambo_colca_config.json
```

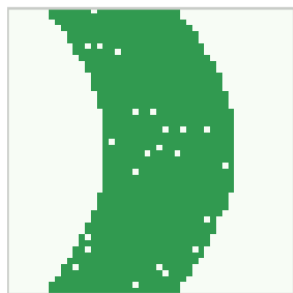
NumPy- and Numba-optimized and parallelized, with memory safeguards to prevent blow-ups: a department-sized scan at 30-m resolution runs in **26.8 minutes**; all of Peru at 90-m, in **4 minutes**.

Striding, closing, pruning

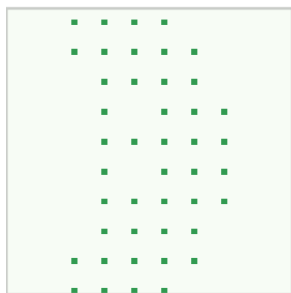
Striding: Before the expensive scan, we keep one surviving candidate pixel in N (here 1 in 5).

A memory-cost control, not a physics criterion

The price is paid downstream: **closing** must reconnect what striding cut apart, and a fragment that stays too small fails the minimum-array-size cut at **pruning**.

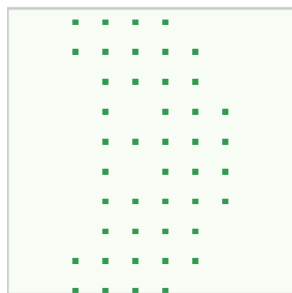


Accepted, every pixel
1.00x the accepted set



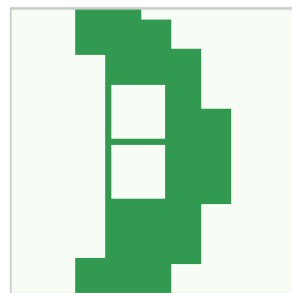
Marked on a 5 px lattice
0.04x the accepted set

□ 3 px wide
— gap 5 px

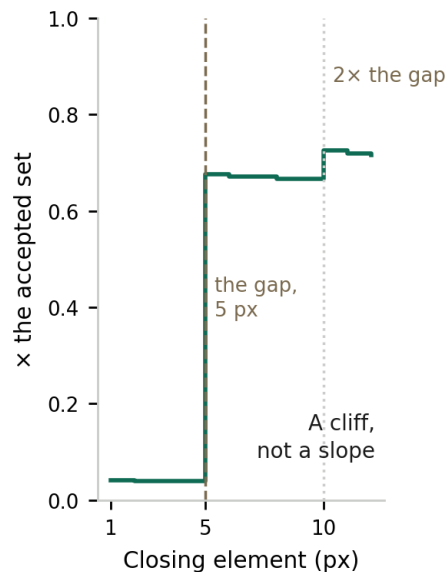


Closed, element 3 px ✗
0.04x the accepted set

□ 5 px wide
— gap 5 px



Closed, element 5 px ✓
0.68x the accepted set



Scoring: ten components, one product

Each score component grades one criterion in $[0, 1]$:

clearance: The radio path stays clear — Sight-line clearance to the target, in Fresnel radii; full marks at one clear zone.

decay: The τ decays inside the gap — $1 - \exp(-\text{gap}/\text{decay length})$, folded over a power-law spectrum.

depth: Rock enough behind the exit — A band on column depth: enough rock to produce the τ , little enough that it escapes.

distance: Exit in the distance window — 1 inside the experiment's window, falling to zero outside it.

footprint: Spacing samples the footprint — The radio footprint (cone angle $\times$ distance, narrower at altitude) vs. the detector spacing.

geomagnetic: Emission in local field — Mean angle between accepted directions and geomagnetic field; 0 along it, so N-S showers fail.

muon_shielding: Rock enough to block muons — pass or fail: the overburden must stop atmospheric muons (TAMBO has > 4 km).

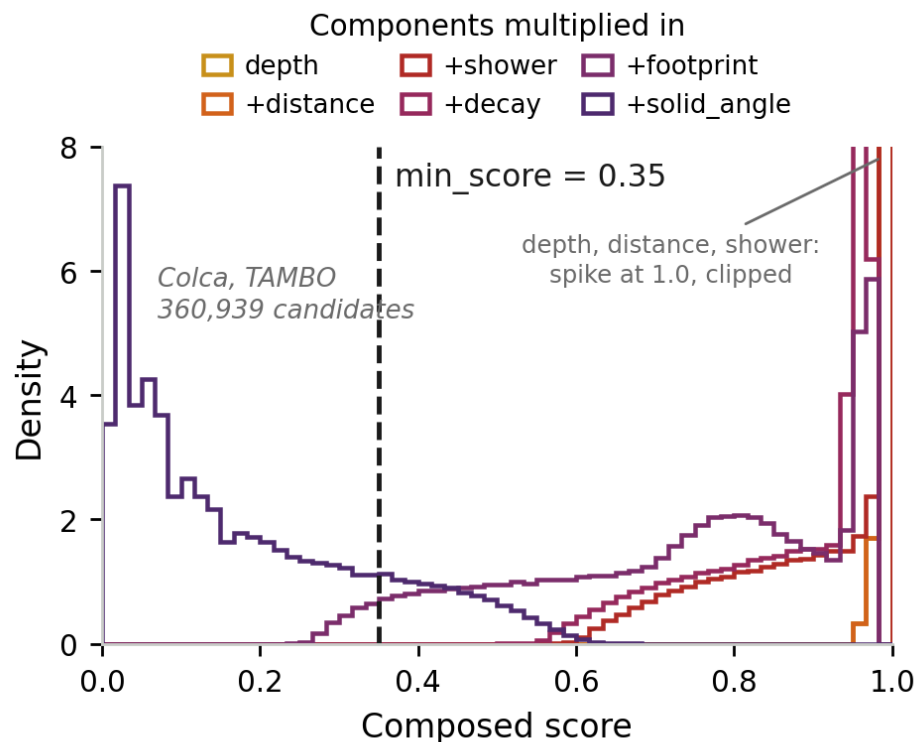
nu_survival: Survives the Earth chord — $\exp(-\text{chord}/\text{interaction length})$: the ν must reach the exit unabsorbed.

shower: Matures before arrival — Air along the path, in g cm^{-2} : the shower must reach its maximum before it reaches the detector.

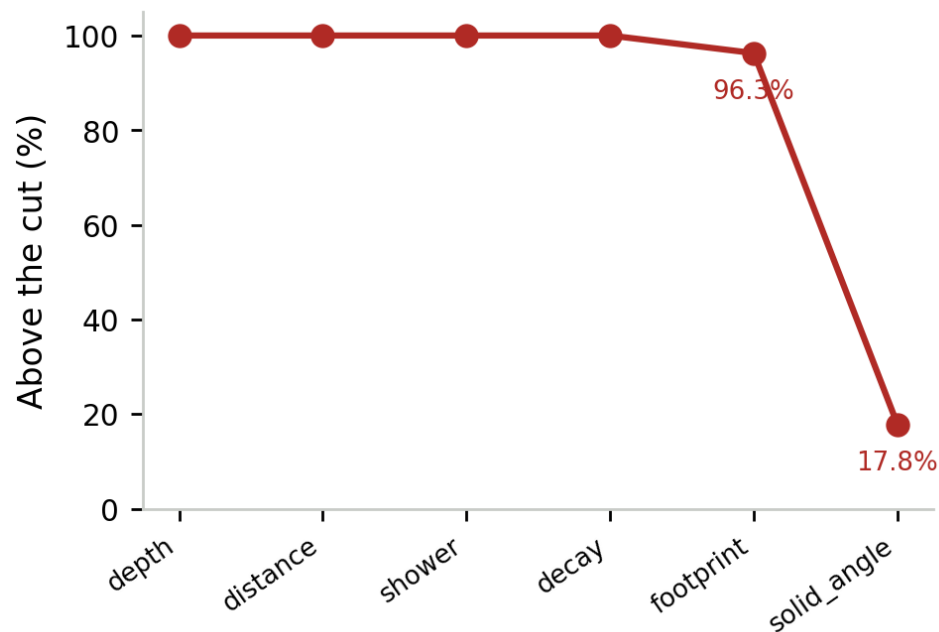
solid_angle: How much sky is accepted — Accepted solid angle as a share of what the fan could accept; saturating, half marks near 8%.

Multiplied together: a site must be good at everything; a single zero rejects it outright

Scoring: ten components, one product



We keep scores only $> \text{min_score}$:
0 (GRAND), 0.35 (TAMBO)
(Not systematically explored; other rankings available in code)



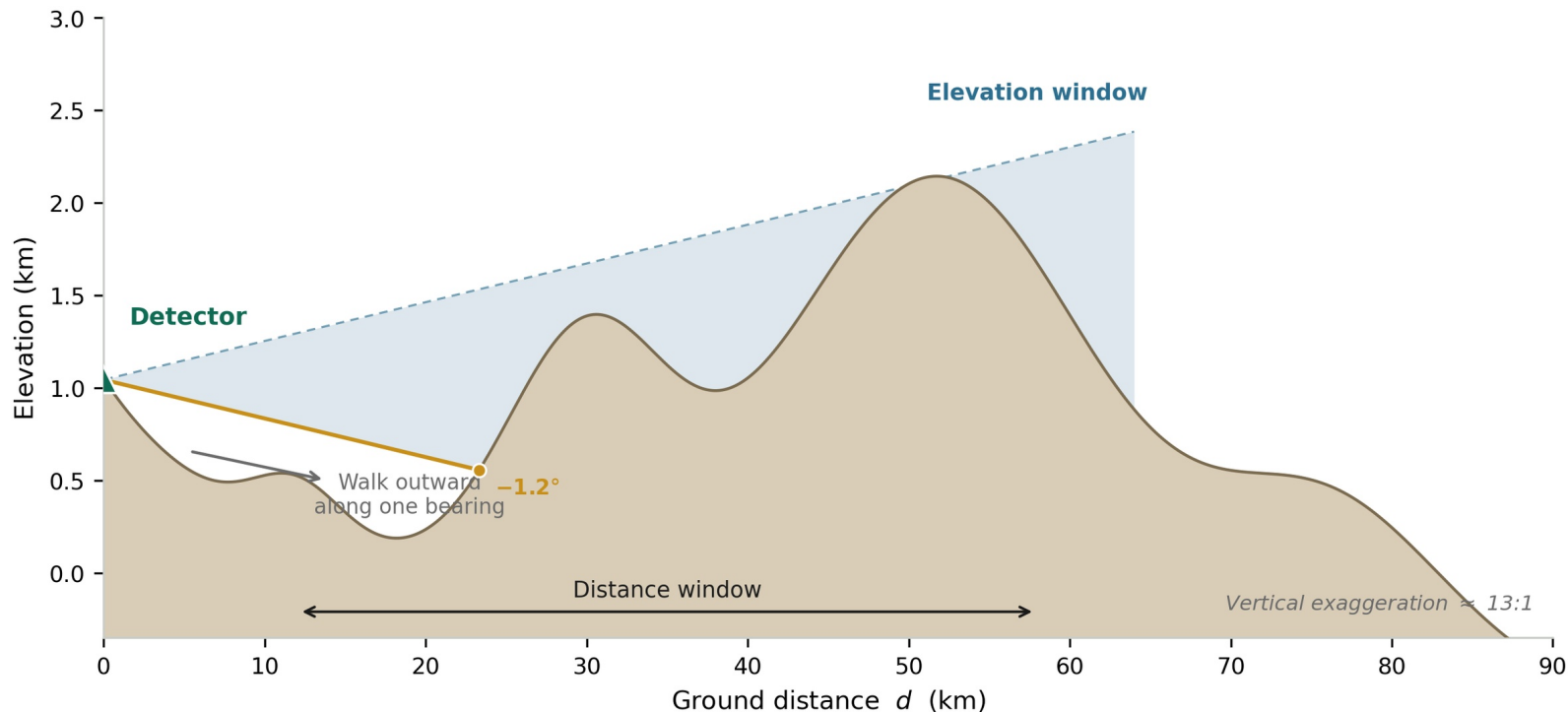
Multiplied together: a site must be good at everything; a single zero rejects it outright

Arrival scan: when does a direction count?

A direction is **accepted** when the first terrain the ray meets lands inside two windows — both drawn here — and, for TAMBO, strikes ground steep enough:

Distance window — how far away a target may sit: GRAND 10–40 km, TAMBO 2–5 km.

Elevation window — how far above or below the horizon: GRAND $\pm 3^\circ$, TAMBO $\pm 20^\circ$.

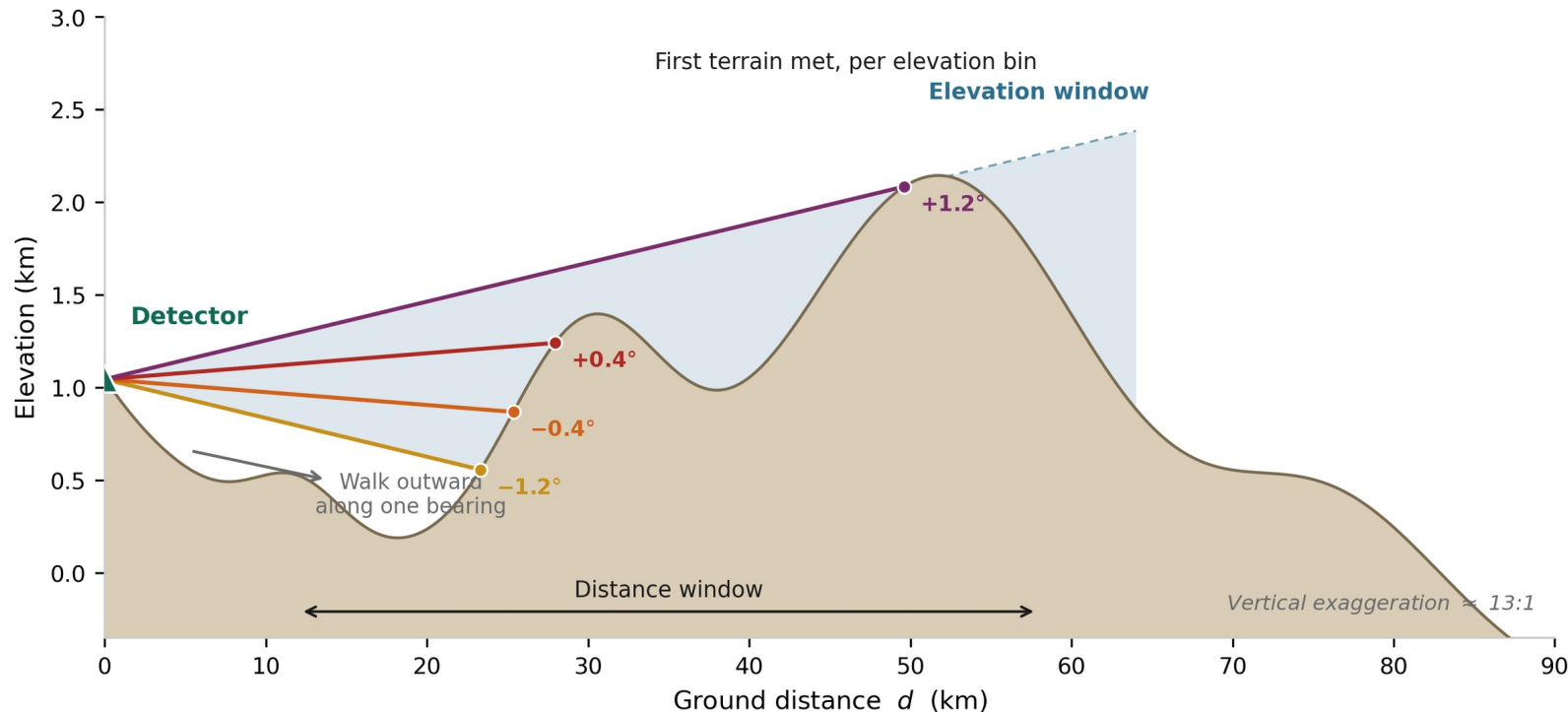


Arrival scan: when does a direction count?

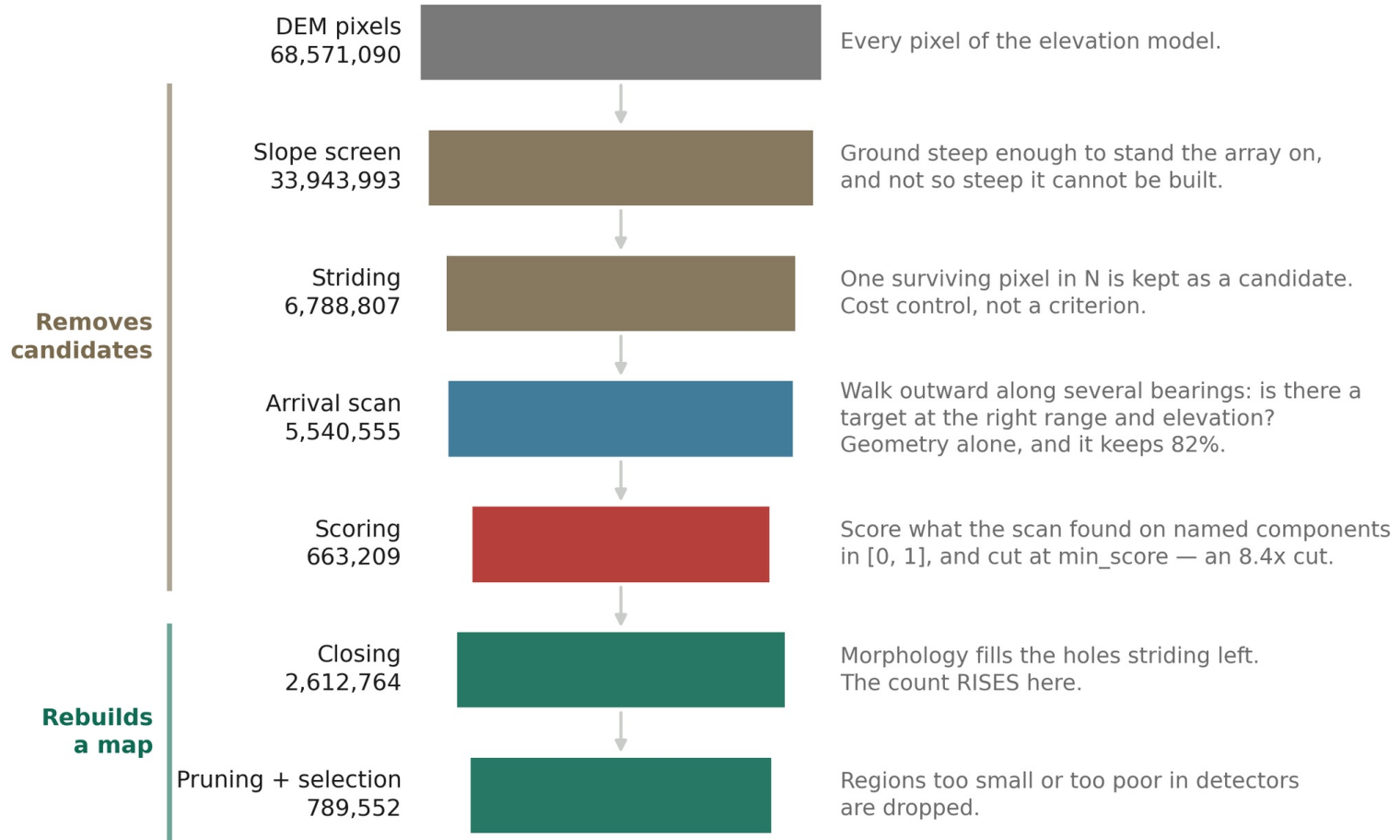
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The Oroscope funnel

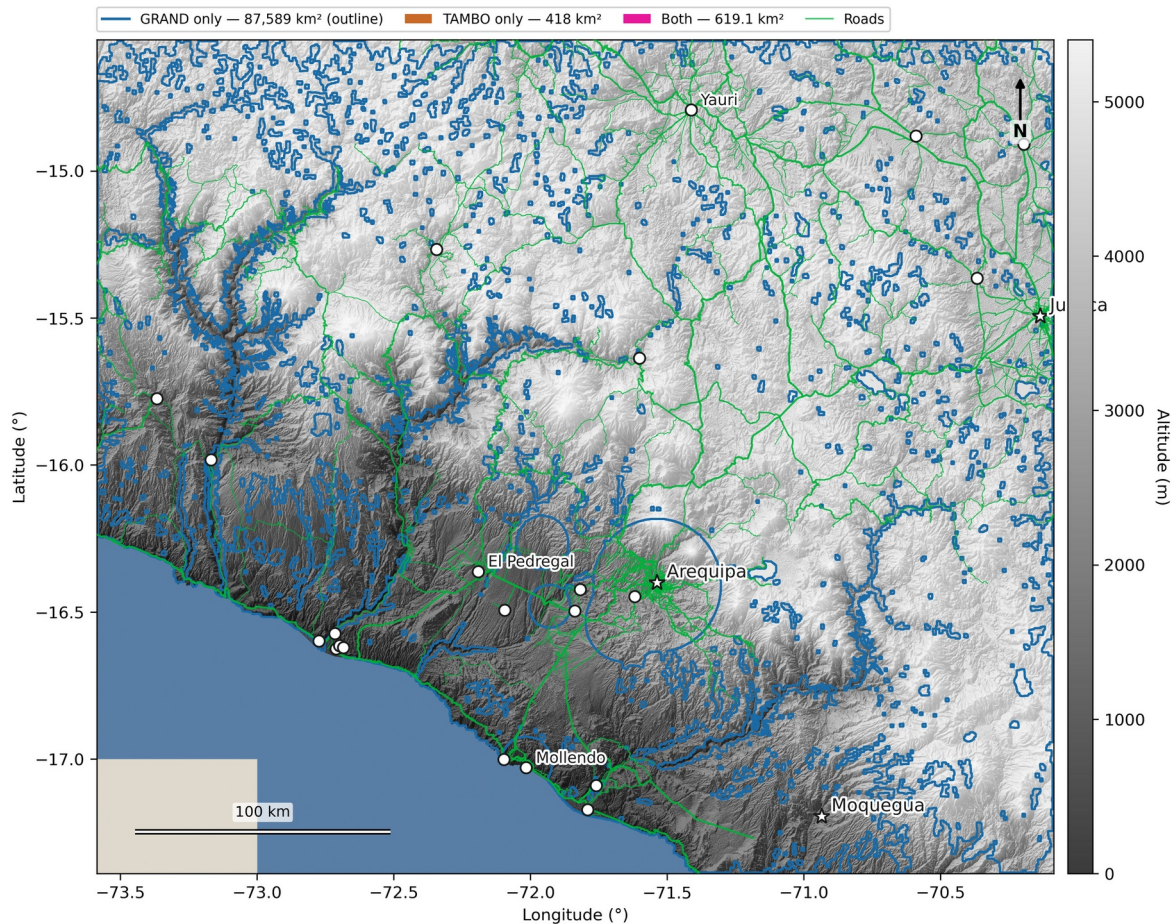


TAMBO over the full Ancash DEM. Bar widths are logarithmic in the survivor count.

What we found

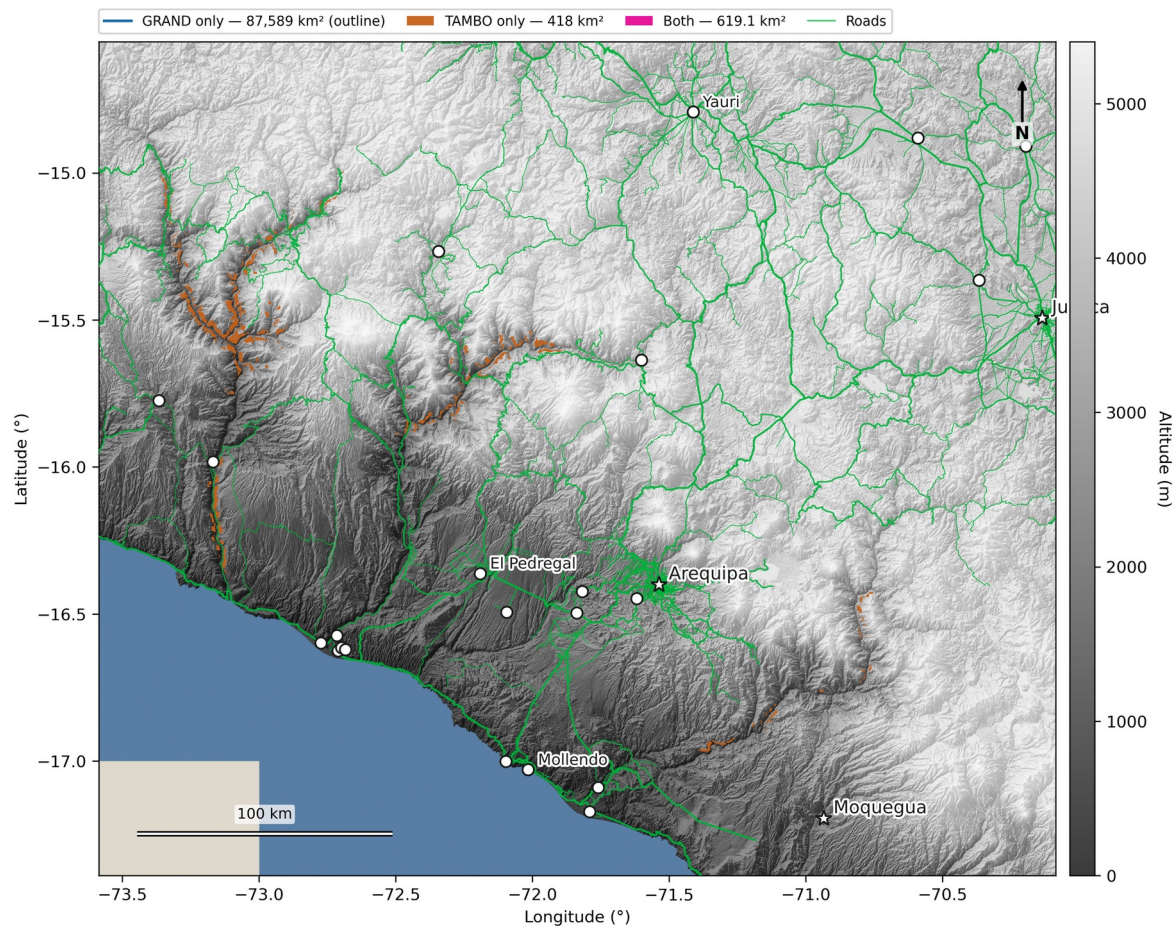
Arequipa, in full – placing GRAND

The blue boundary encloses **87,589 km²** – plateau and gentle ground, most of the department.



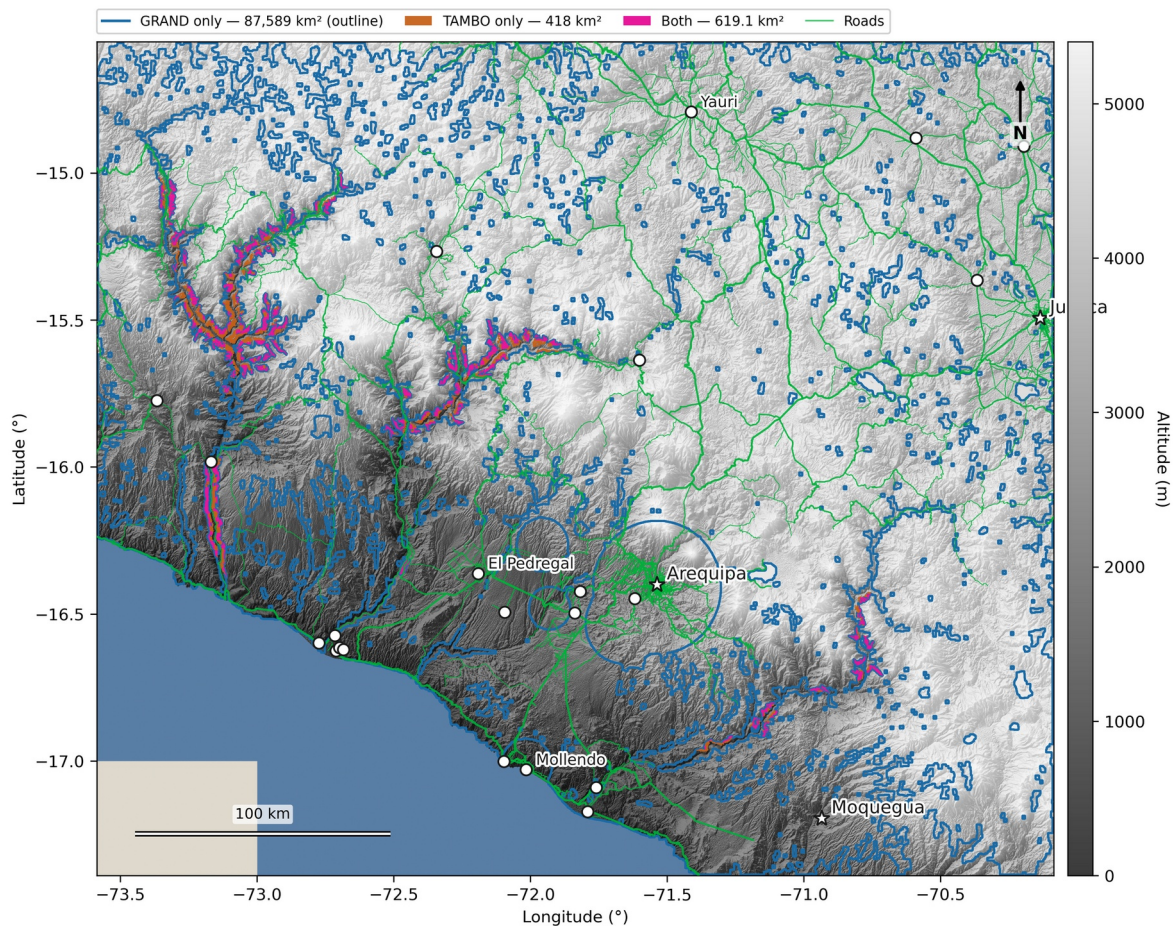
Arequipa, in full – placing TAMBO

418 km², strung along the canyons — TAMBO is the scarce partner from the start.



Arequipa, in full – placing GRAND + TAMBO

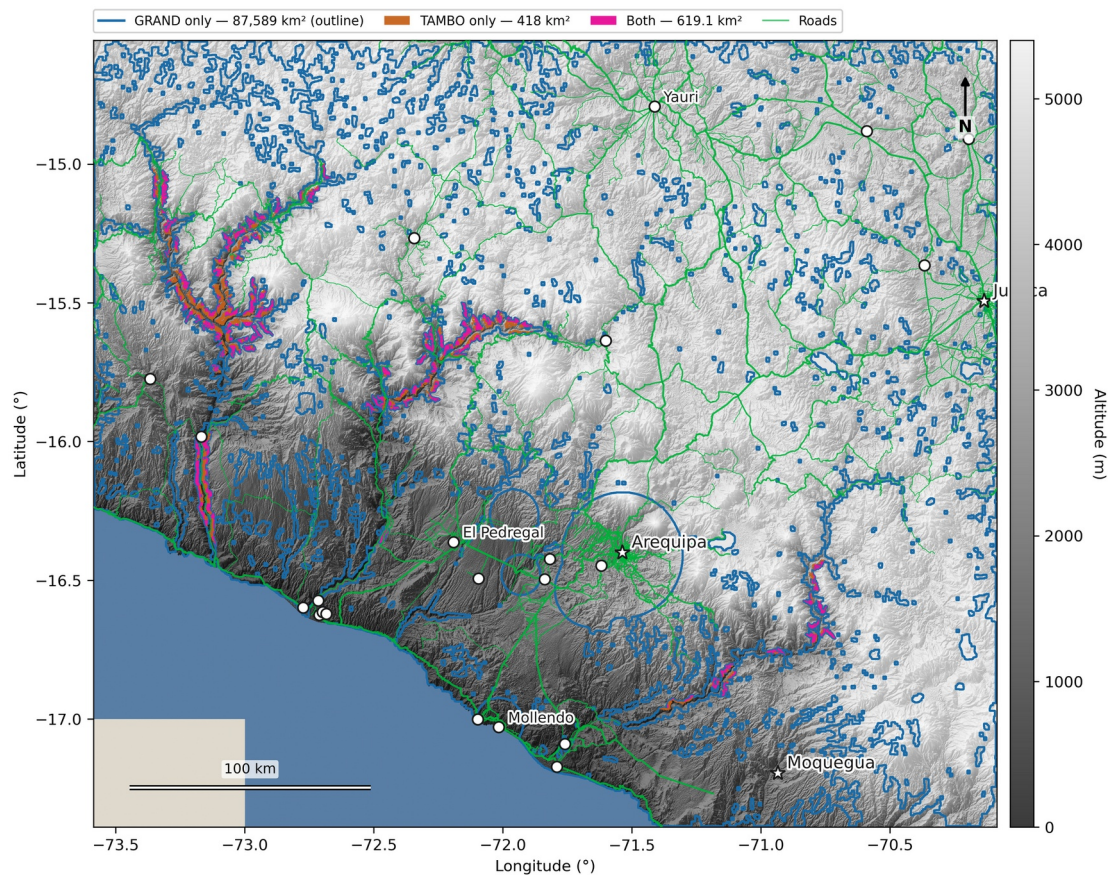
619 km² shared — 59.7% of everything TAMBO-viable (619 of 1,037 km²) is also GRAND-viable.



Every run explains itself

Oroscope output includes: a map, **GeoTIFF mask**, **KML** (for Google Earth), **results JSON**

Plus, plain-language account of the results (**explanation.txt**): what was found, which stage bound it, *etc.*



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WHERE THESE EXPERIMENTS CAN SHARE GROUND, AND WHY

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WHAT EACH BRINGS

experiment	area km²	sites	detectors	in joint
------------	----------	-------	-----------	----------

GRAND	88,208.2	1	101,584	0.7%
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TAMBO	1,036.9	85	49,271	59.7%
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Ground satisfying GRAND and TAMBO at once: 619.1 km². Ground useful to any of them: 88,626.0 km². The first is the number that matters for one site, one road and one power feed serving two experiments; the second is what the programme as a whole could use.

WHAT DECIDES IT

A pixel has one slope, one altitude and one aspect, and both experiments must accept those same values to stand on it. So co-location is settled by whichever of those bands they share least of – before any arrival geometry is considered.

band	GRAND	TAMBO	shared
deployable slope	3–25°	20–60°	20–25° (23%)
target distance	GRAND: 10–40 km, TAMBO: 2–5 km		
arrival elevation	GRAND: -3–3°, TAMBO: -20–20°		

PAIRWISE

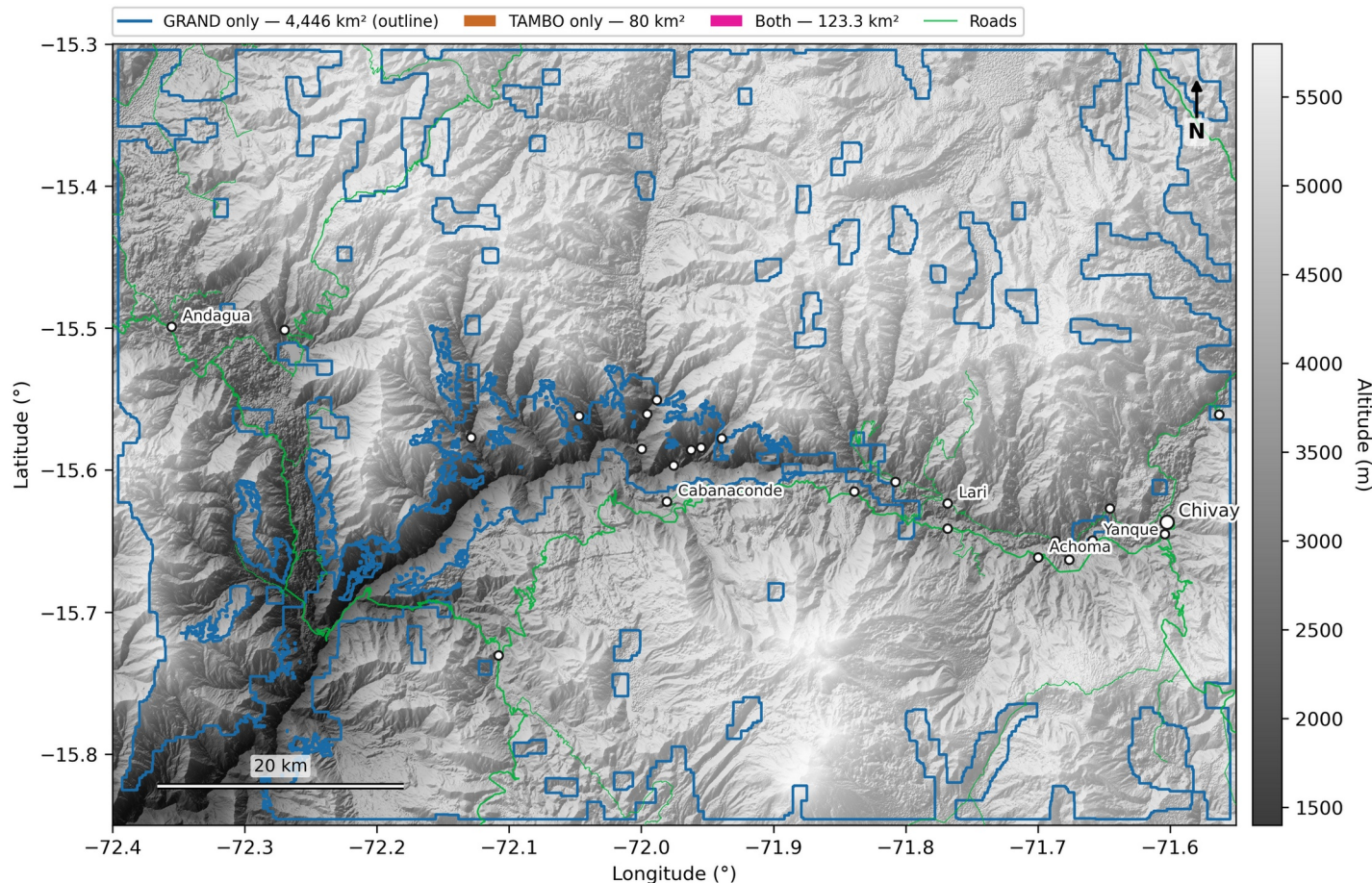
GRAND & TAMBO: 619.1 km² shared, Jaccard 0.0070.
0.7% of GRAND's own ground
59.7% of TAMBO's own ground

HOW TO READ THIS

Every caveat on the individual runs applies here and compounds: these are the same masks, so morphological closing has already moved each area, and the joint of two inflated masks is inflated twice over. The overlay is exact – the masks are pixel-aligned and the alignment is checked, not assumed – but what it overlays is only as good as each run. And co-location is a question about ground, not about detectors. Two experiments sharing a hillside still need their own arrays, their own spacing and their own trigger; what they share is the site, the access and the power.

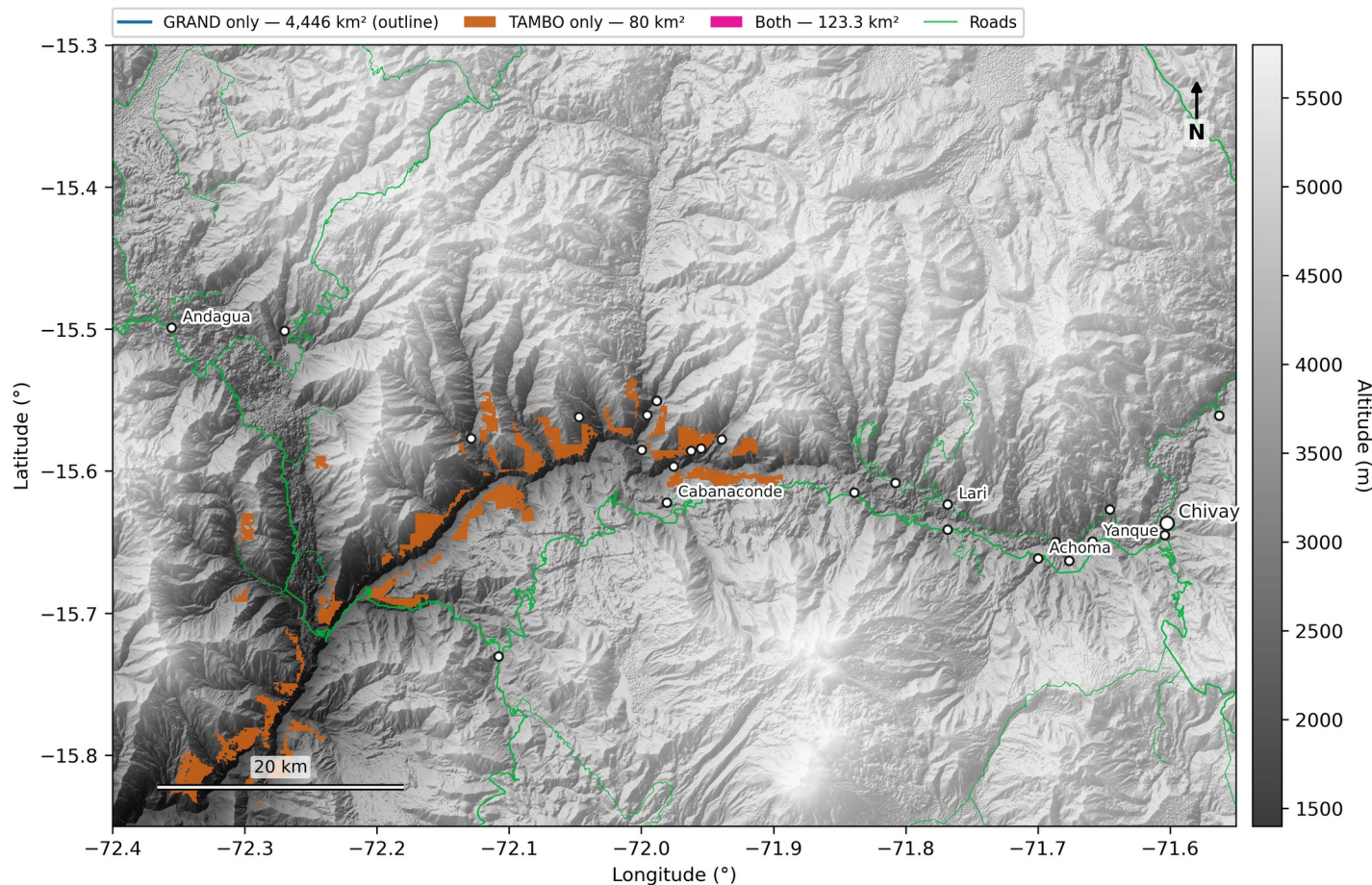
Colca (Arequipa) zoom – placing GRAND

4,446 km² — most of the frame, plateau and gentle ground.



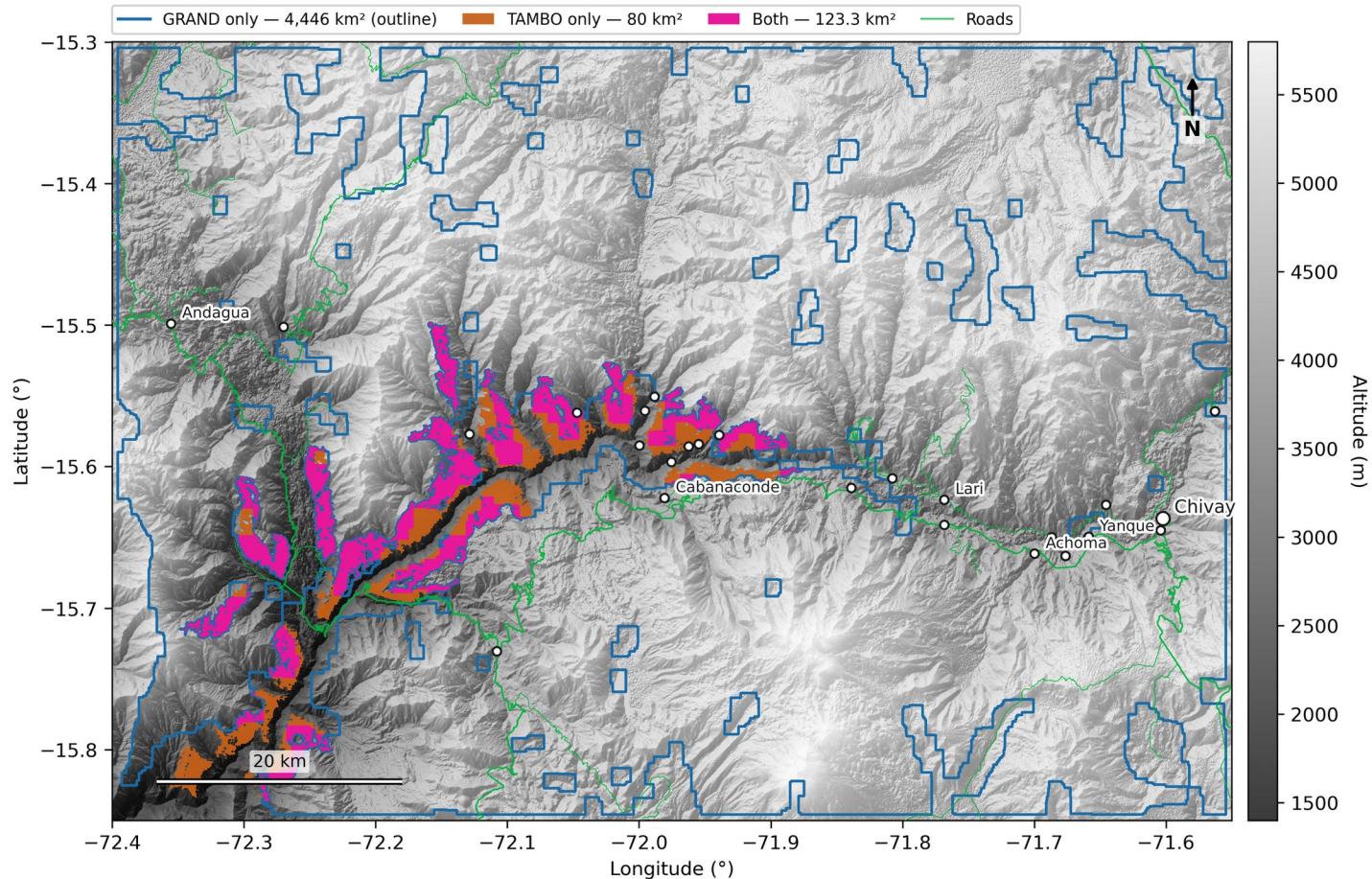
Colca (Arequipa) zoom – placing TAMBO

80 km² — a scatter of patches threading the canyon system.



Colca (Arequipa) zoom – placing **GRAND + TAMBO**

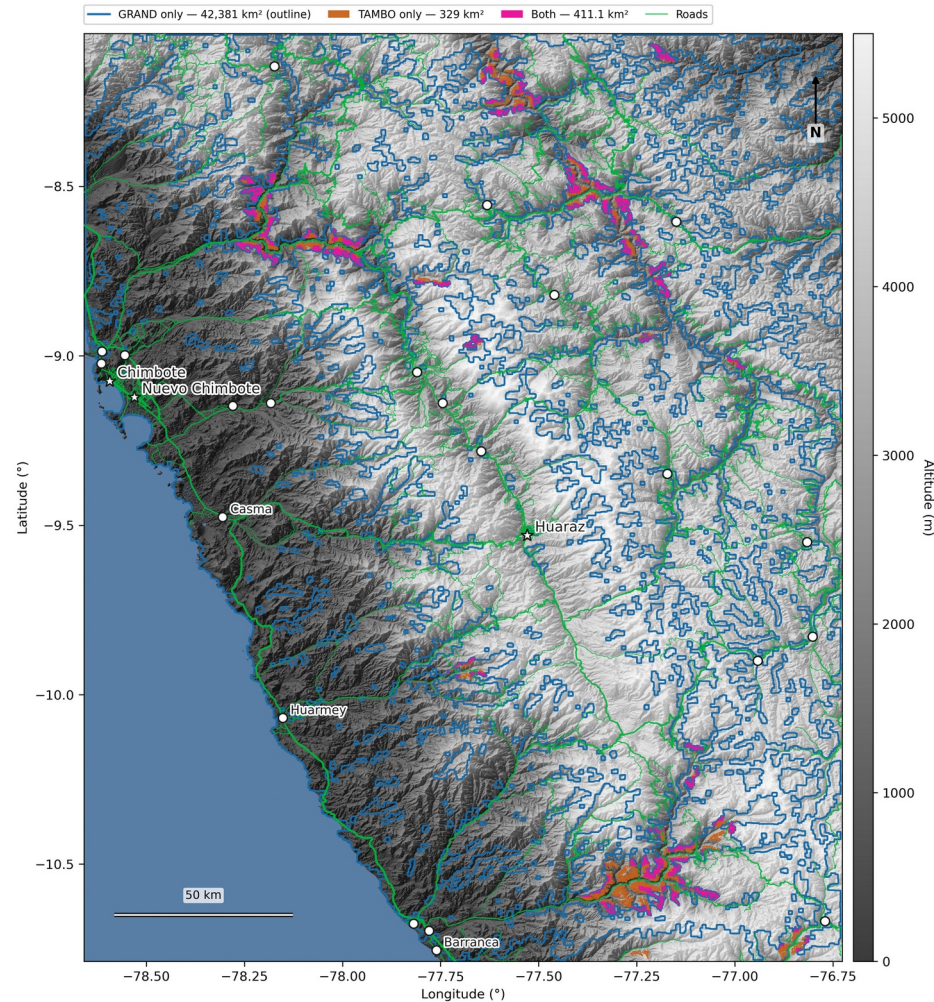
123.3 km² shared — 60.7% of TAMBO's 203 km², tracing the canyon rims — where the roads also run.



Áncash, in full – placing GRAND + TAMBO

411.1 km² shared —
55.6% of TAMBO's 740
km².

Subsampled, so a lower
bound.

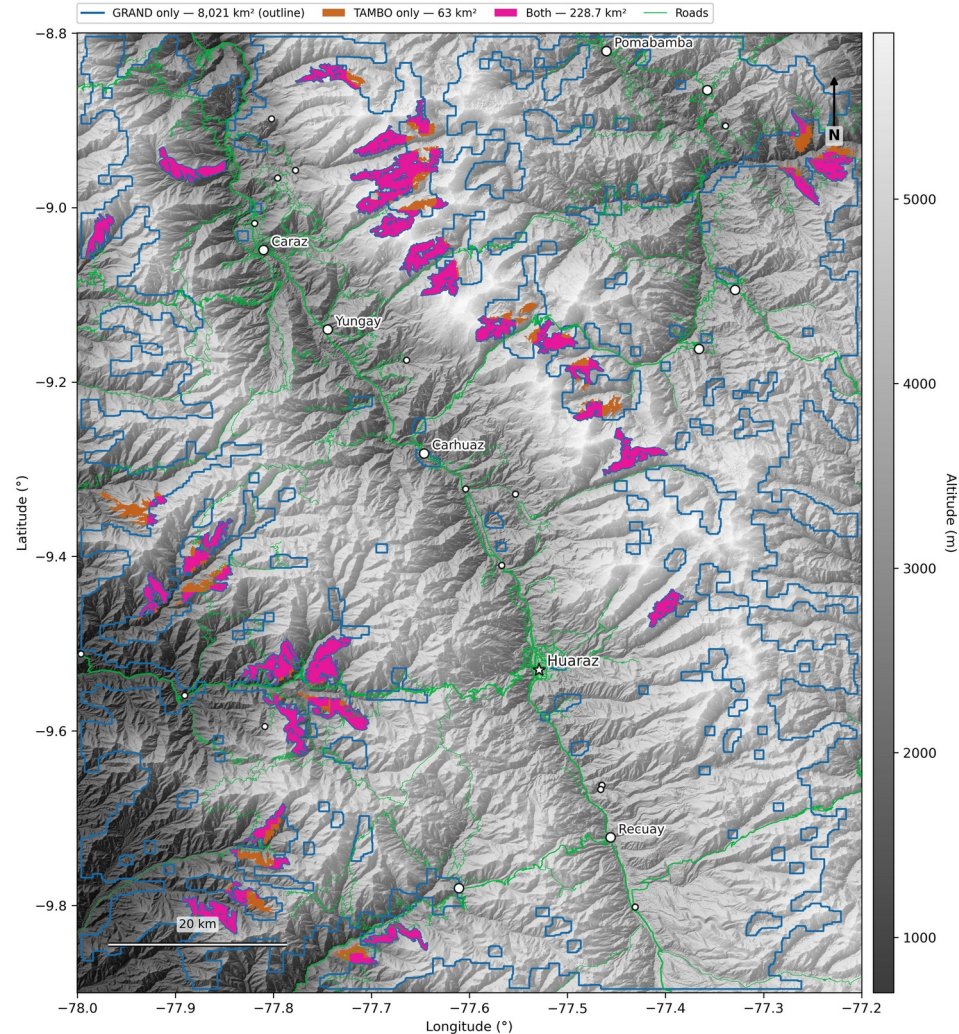


Roads and place names © OpenStreetMap contributors (ODbL). Elevation: SRTMGL1 via OpenTopography.

Callejón de Huaylas (Áncash) – placing GRAND + TAMBO

No subsampling

228.7 km² of TAMBO's 291.3 km² is also GRAND-viable:
78.5% (highest measured).

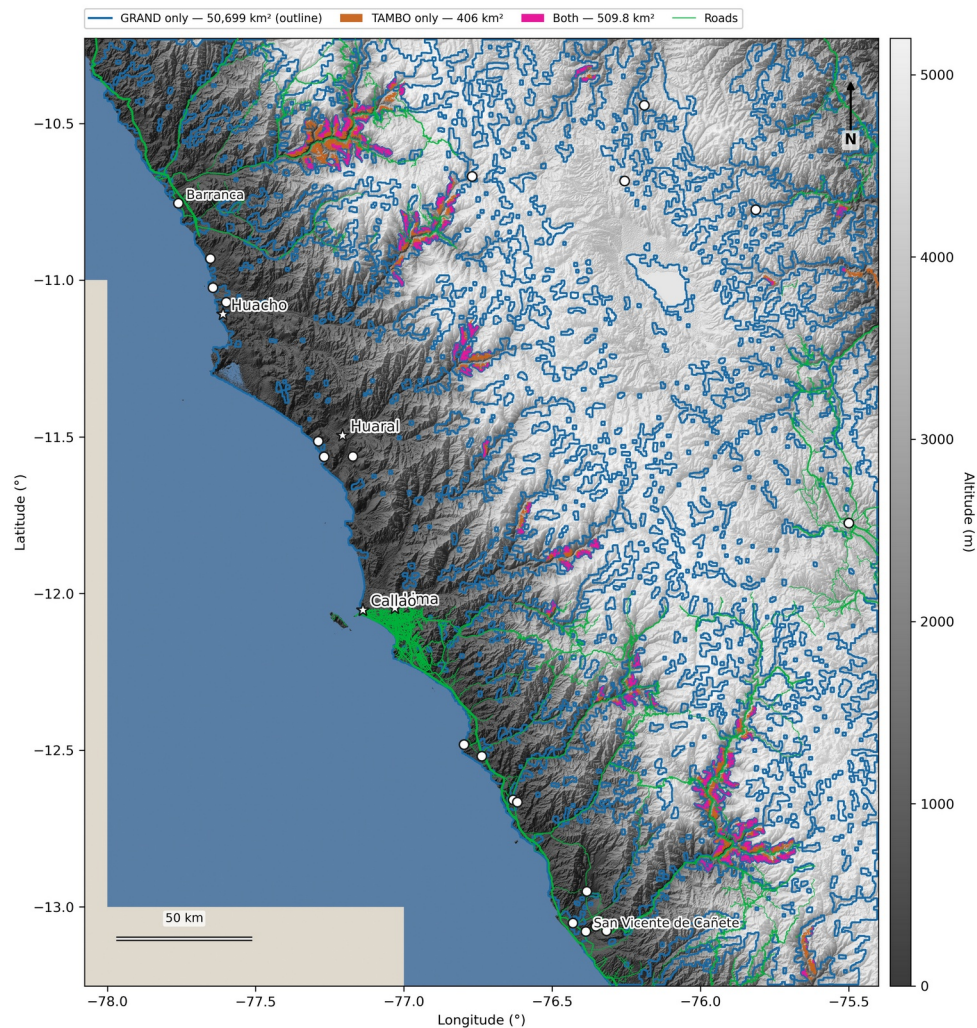


Roads and place names © OpenStreetMap contributors (ODbL). Elevation: SRTMGL1 via OpenTopography.

Lima, in full – placing GRAND + TAMBO

509.8 km² shared — 55.7%
of TAMBO's 915 km².

Subsampled, so a lower
bound.

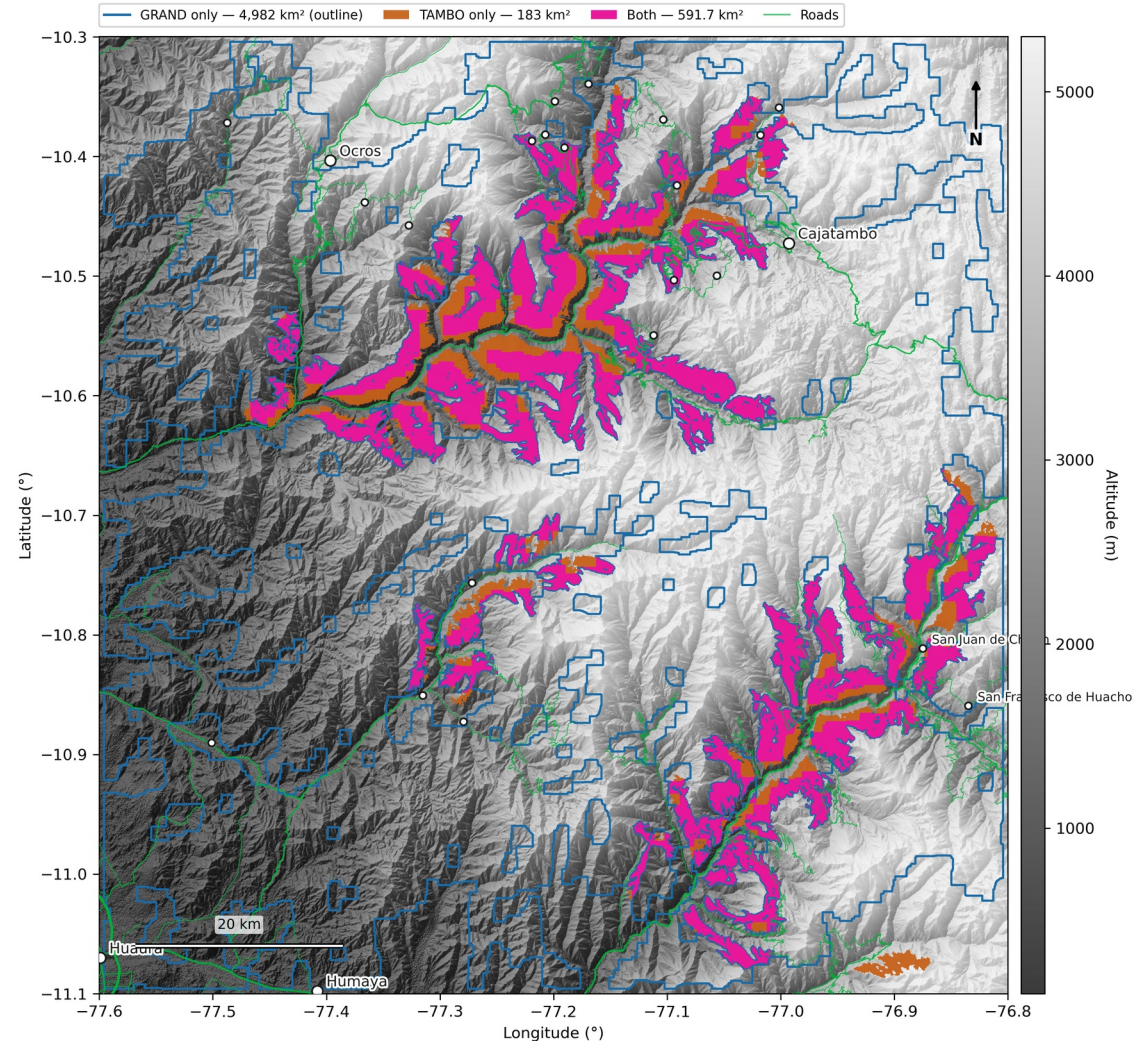


Roads and place names © OpenStreetMap contributors (ODbL). Elevation: SRTMGL1 via OpenTopography.

Cajatambo (Lima) – placing **GRAND** + **TAMBO**

509.8 km² shared — 55.7%
of TAMBO's 915 km².

Subsampled, so a lower
bound.



Roads and place names © OpenStreetMap contributors (ODbL). Elevation: SRTMGL1 via OpenTopography.

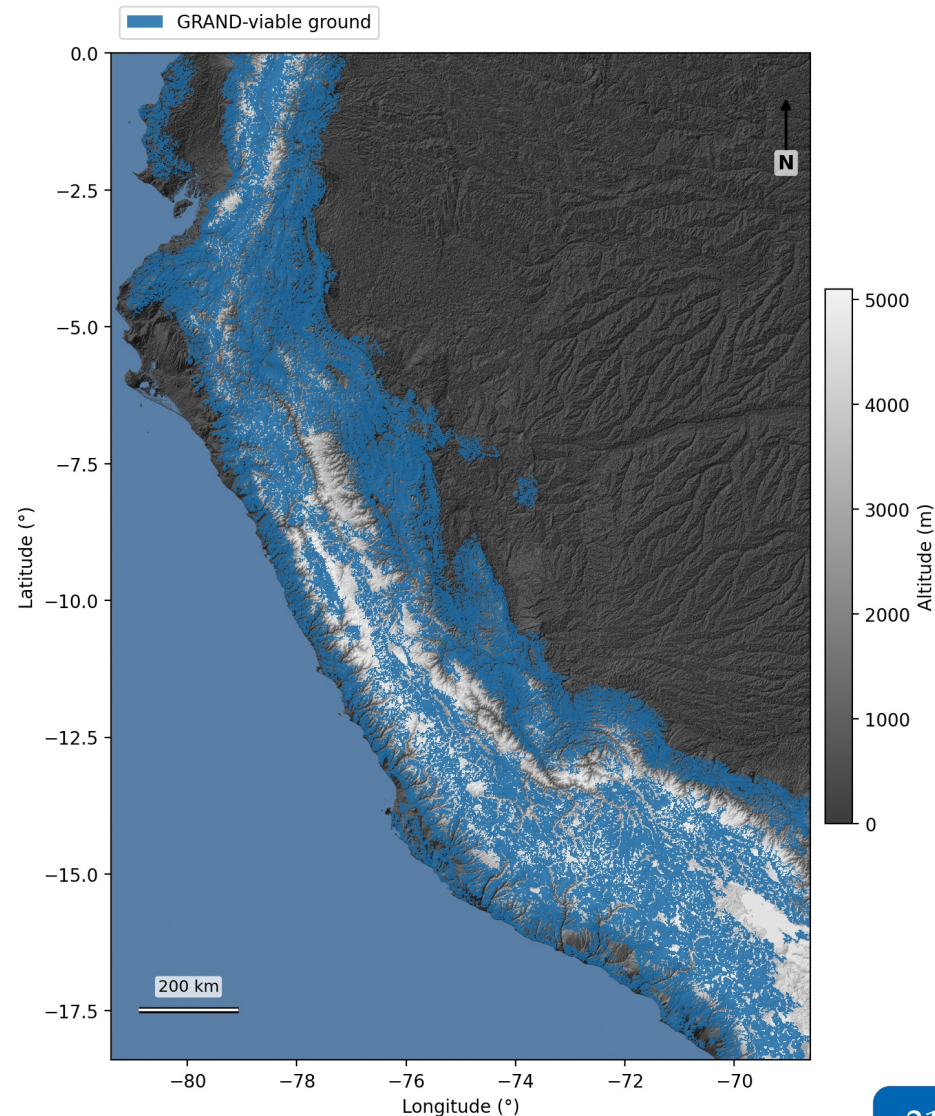
All of Peru

Only on coarse 90-m grid (SRTMGL3, 339 Mpx):

400,000–600,000 km² of GRAND-viable ground.

The slope screen is the binding constraint.

There is no national TAMBO or joint answer yet. It needs the 30-m grid over the whole country, plus tiling; that has not been run.



Shared area in Arequipa

The **619 km²** of shared GRAND + TAMBO ground, divided by each experiment's own:

0.7%

619 / 88,208 km²

of GRAND's viable ground

59.7%

619 / 1,037 km²

of TAMBO's viable ground

Co-location costs GRAND almost nothing, and consumes most of what TAMBO has.

The shared ground is TAMBO-limited in every region measured — so the site choice is TAMBO's to lead, and GRAND's to join.

TAMBO takes the canyon, GRAND stands a ways off

The shared ground is canyon wall: median slope 35° at Colca

At 150-m spacing, it holds 6,322 TAMBO units (**enough for full TAMBO**)

Ground gentle enough for GRAND is 10 km² of slivers there, about 8 antennas' worth.

GRAND was never meant to stand on it

What couples the arrays is a shared line of sight to the same massif — its lattice lies a median 1.0 km from the shared ground (83% of it within 3 km)

From each region's best TAMBO site: 100 antennas fit within ~10 km, 1,000 within 20–30 km, 5,000 within 60 km.

Six regions explored

Region	GRAND km ²	TAMBO km ²	Joint km ²	Share of TAMBO
Huaylas	8,249.5	291.3	228.7	78.5%
Cajatambo	5,573.8	774.5	591.7	76.4%
Colca	4,569.4	203.0	123.3	60.7%
Ancash	42,791.9	740.0	411.1	55.6%
Lima	51,209.0	915.4	509.8	55.7%
Arequipa	88,208.2	1,036.9	619.1	59.7%

Share of TAMBO = joint area ÷ TAMBO's own viable area — the fraction of TAMBO-viable ground that GRAND could also stand on.

Shaded rows: every pixel was tested. The rest were subsampled to fit in memory — read those as lower bounds.

Roughly three quarters to four fifths of TAMBO-viable ground is also GRAND-viable.

How many detectors is that?

Spacing is the input (GRAND 1 km, TAMBO 150 m). **Capacity is computed:** a hexagonal lattice is packed into each region's mask, position by position — not $\text{area} \div \text{spacing}^2$.

Region	GRAND antennas	TAMBO units
Huaylas	9,559	14,925
Cajatambo	6,457	39,658
Colca	5,315	10,437
Ancash	49,059	34,275
Lima	58,669	42,549
Arequipa	101,584	49,271

TAMBO's design target is ~5,000 units — Colca alone holds twice that. Every region measured clears it. **GRAND's 10,000-antenna design** fits in any of the three departments; no single crop holds it. Shaded rows unbiased, the rest subsampled — the same lower-bound caveat as the areas.

What this does
not tell you

What is missing, and what's next

Nothing has been checked against an external simulation. Everything here is internally consistent and externally unvalidated. The cheapest test is the Earth-absorption prediction — an arrival-window edge of -4.4° at 100 PeV rising to -0.9° at 10 EeV — and it has not been done.

Detector acceptance $A(E)$ is modeled poorly. Two published *integral* curves are digitized, and an array-size correction cannot correct for the site.

Neutral-current regeneration is not modeled, so Earth-chord suppression is somewhat overstated.

Magnetic declination falls back to the Arequipa value rather than being per-site.

A **national TAMBO answer** needs the 30 m grid over the whole country, and tiling — the Peru map you saw is GRAND-only.

Price tag

1 GRAND DU $\sim$ 3.5k € $\sim$ 1 TAMBO unit

What is the minimum number of GRAND DUs needed?

About 10 (for five-DU coincidence triggers), so $\sim$ 35k €

Take-home

- Two detectors that want opposite hillsides do not want opposite **views**.
- Co-location limited by **slope**, not by sight: bands overlap only over 20–25°.
- Even so, **three quarters to four fifths** of TAMBO-viable ground is also GRAND-viable.
- It costs GRAND **0.7%** of its ground and gives TAMBO a partner within **10 km** of its best sites.

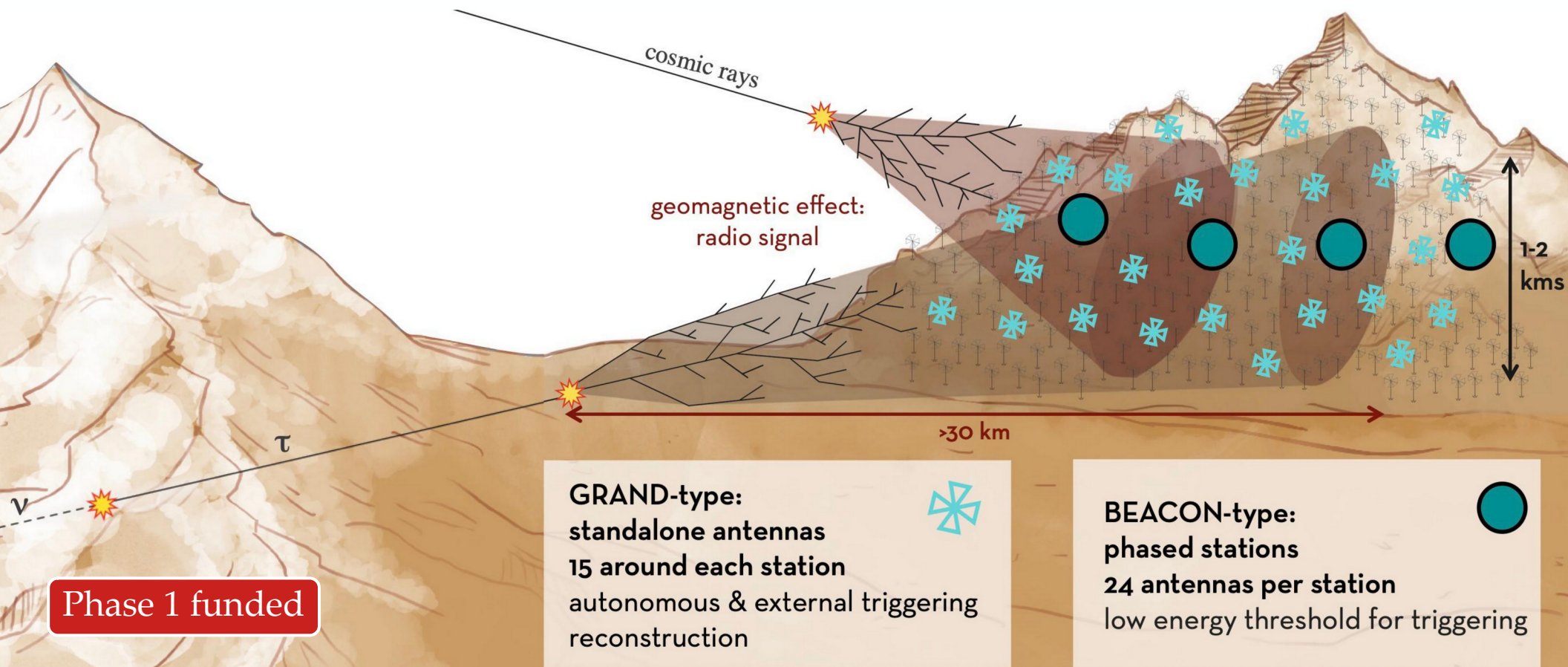
The terrain does not forbid it. The question worth asking next is not **whether** they can share — but **which massif**.

Thank you!

Backup slides

HERON: Hybrid Elevated Radio Observatory for Neutrinos

BEACON phased-array stations (low-SNR detection) + GRAND antennas (self-triggering)



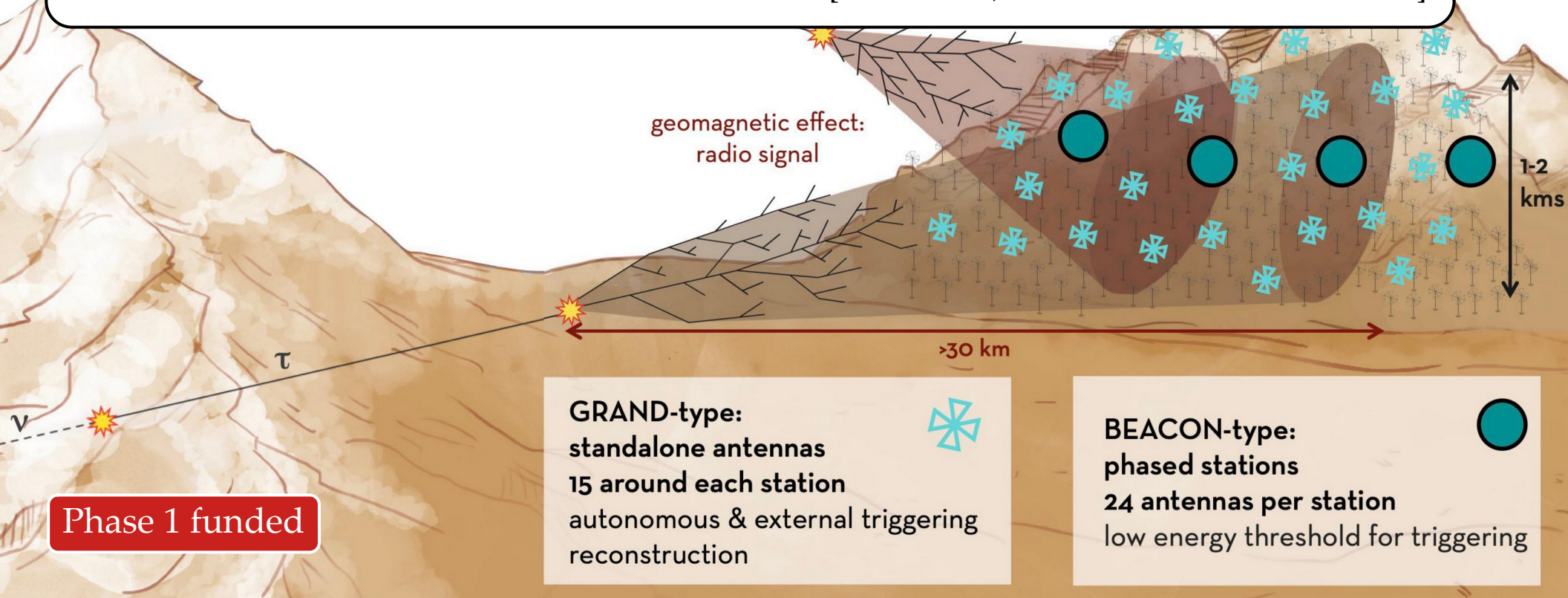
HERON: Hybrid Elevated Radio Observatory for Neutrinos

BEACON phased-array stations (low-SNR detection) + GRAND antennas (self-triggering)

Narrow field of view, high sensitivity:

Phase 1: 24 BEACON stations + 360 GRAND antennas [$\approx$ GRAND10k sens. for point sources]

Phase 2: 200 BEACON stations + 3k GRAND antennas [$\approx$ GRAND, IC-Gen2 sens. for diffuse flux]



Partner-array distances, per region

How far from each region's best TAMBO site a GRAND array of a given size fits.

From each region's best TAMBO site	100 antennas	1,000	5,000
Arequipa	10 km	20 km	60 km
Ancash	10 km	30 km	60 km
Lima	10 km	30 km	60 km

The limit is only that 1,000 antennas at 1 km spacing need 866 km², and steep country does not offer that within a few kilometres. These are capacity counts from packing a lattice, not optimised layouts.

Four caveats

- 1 Reported area is an upper bound.** Closing — the stage that fills the holes striding leaves — inflates the mask before area is measured: **2.35×** at Colca, against a stride-1 control.
- 2 A subsampled TAMBO area is a lower bound** — by **1.51×** at Colca, **23.0×** on the Callejón de Huaylas. Striding does not bias which pixels pass; the loss is in rebuilding the map, where a fragmented mask fails the minimum-array-size cut at pruning.
- 3 min_score is the dominant assumption — and it is TAMBO's alone.** GRAND's configurations cut nothing. The shipped 0.35 sits at the **17.8th percentile** of a product that piles up near zero, and a sweep across the cut finds **no knee anywhere** — every TAMBO area here moves if you move it.
- 4 solid_angle is the weakest component at every selected TAMBO site in every region** — 16/16 Colca, 62/62 Ancash, 84/84 Lima, 85/85 Arequipa. Under a product, the weakest term bounds the site — so the TAMBO result rests on that one component.